

# Fundamental Interactions

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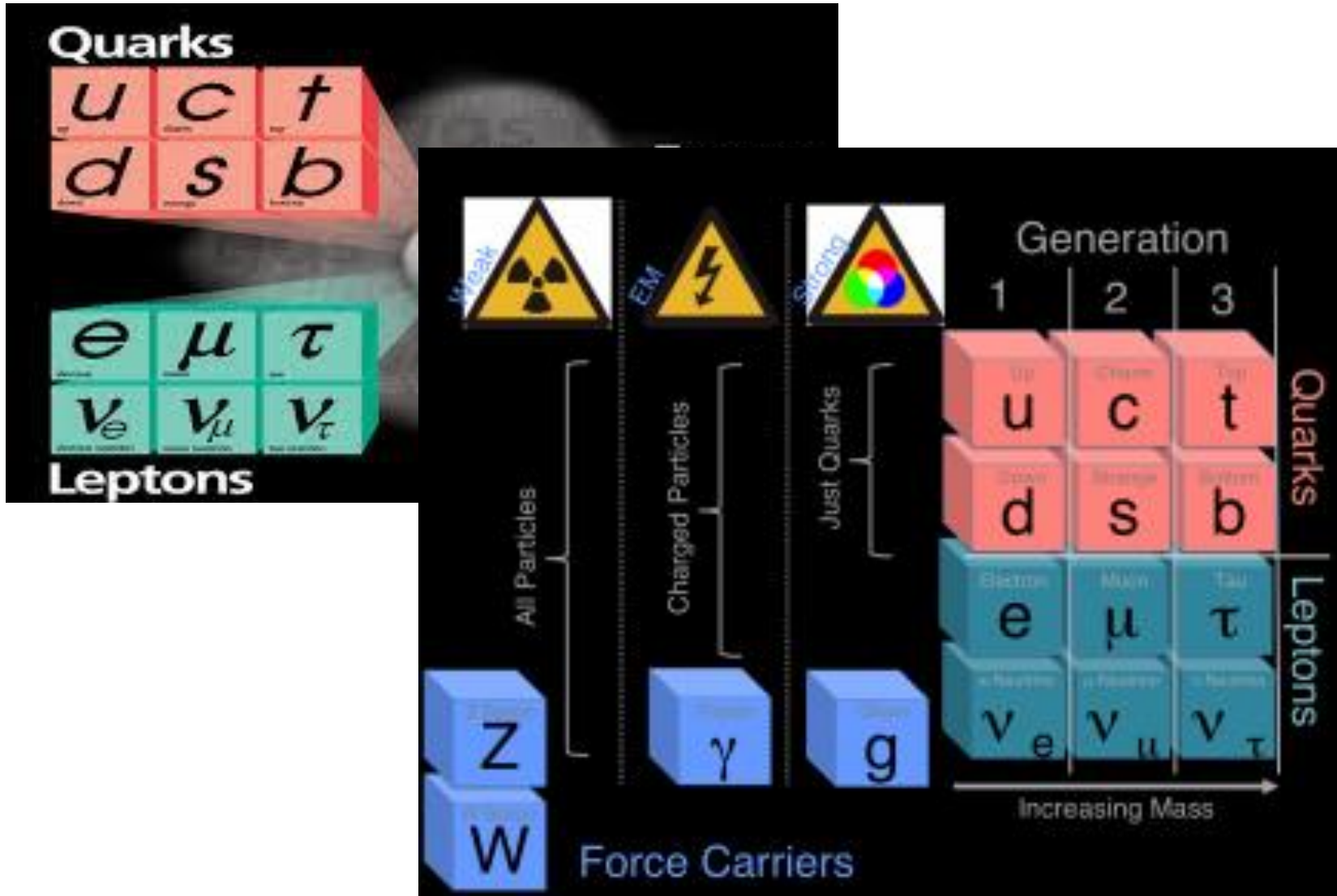
# Outline of 2 lectures

- **Standard Model**
  - Very basic introduction
  - Recent changes
- **Fundamental interactions at low energy**
  - General approach
    - Symmetries, conservation laws
    - Nuclei as laboratory
    - Tools → traps
  - Detailed look at present and planned work
    - Superallowed Fermi decay
    - Angular correlations in  $\beta$ -decay
    - PNC in atoms
    - EDM
    - Others ...
- **FRIB and fundamental interactions**



2nd lecture

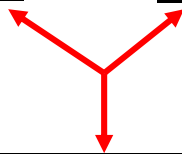
# Standard Model constituents



# Fundamental particles and interactions

FERMIONS			matter constituents spin = 1/2, 3/2, 5/2, ...		
Leptons spin = 1/2			Quarks spin = 1/2		
Flavor	Mass GeV/c <sup>2</sup>	Electric charge	Flavor	Approx. Mass GeV/c <sup>2</sup>	Electric charge
$\nu_e$ electron neutrino	$<1 \times 10^{-8}$	0	<b>u</b> up	0.003	2/3
<b>e</b> electron	0.000511	-1	<b>d</b> down	0.006	-1/3
$\nu_\mu$ muon neutrino	$<0.0002$	0	<b>C</b> charm	1.3	2/3
<b><math>\mu</math></b> muon	0.106	-1	<b>S</b> strange	0.1	-1/3
$\nu_\tau$ tau neutrino	$<0.02$	0	<b>t</b> top	175	2/3
<b><math>\tau</math></b> tau	1.7771	-1	<b>b</b> bottom	4.3	-1/3

BOSONS			force carriers spin = 0, 1, 2, ...		
Unified Electroweak spin = 1			Strong (color) spin = 1		
Name	Mass GeV/c <sup>2</sup>	Electric charge	Name	Mass GeV/c <sup>2</sup>	Electric charge
$\gamma$ photon	0	0	<b>g</b> gluon	0	0
<b>W<sup>-</sup></b>	80.4	-1			
<b>W<sup>+</sup></b>	80.4	+1			
<b>Z<sup>0</sup></b>	91.187	0			



High energy

Low energy

PROPERTIES OF THE INTERACTIONS					
Property \ Interaction	Gravitational	Weak (Electroweak)		Strong	
		Flavor	Electric Charge	Fundamental	Residual
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge	See Residual Strong Interaction Note
Particles experiencing:	All	Quarks, Leptons	Electrically charged	Quarks, Gluons	Hadrons
Particles mediating:	Graviton (not yet observed)	<b>W<sup>+</sup> W<sup>-</sup> Z<sup>0</sup></b>	$\gamma$	Gluons	Mesons
Strength relative to electromag for two u quarks at:	$10^{-41}$ m	0.8	1	25	Not applicable to quarks
	$3 \times 10^{-17}$ m	$10^{-4}$	1	60	
	for two protons in nucleus	$10^{-36}$	$10^{-7}$	1	Not applicable to hadrons



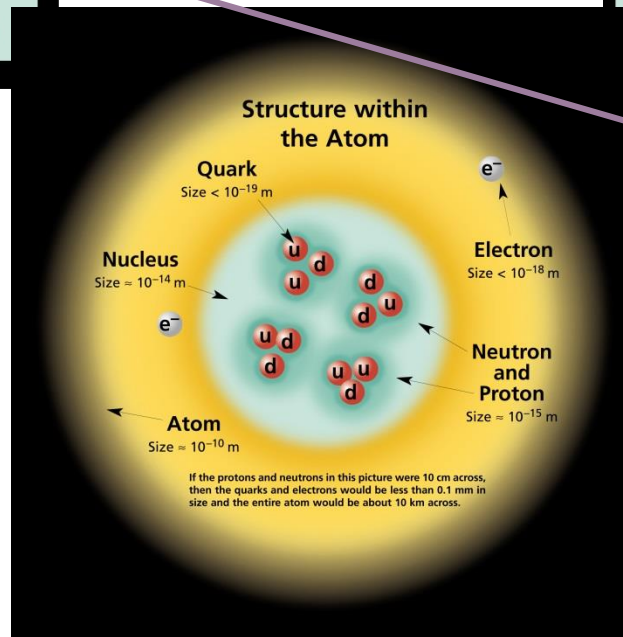
# Resulting composite particles

Mesons $q\bar{q}$					
Mesons are bosonic hadrons. There are about 140 types of mesons.					
Symbol	Name	Quark content	Electric charge	Mass GeV/c <sup>2</sup>	Spin
$\pi^+$	pion	$u\bar{d}$	+1	0.140	0
$K^-$	kaon	$s\bar{u}$	-1	0.494	0
$\rho^+$	rho	$u\bar{d}$	+1	0.770	1
$B^0$	B-zero	$d\bar{b}$	0	5.279	0
$\eta_c$	eta-c	$c\bar{c}$	0	2.980	0



Composite particles must be colour neutral

Baryons $qqq$ and Antibaryons $\bar{q}\bar{q}\bar{q}$					
Baryons are fermionic hadrons. There are about 120 types of baryons.					
Symbol	Name	Quark content	Electric charge	Mass GeV/c <sup>2</sup>	Spin
$p$	proton	$uud$	1	0.938	1/2
$\bar{p}$	anti-proton	$\bar{u}\bar{u}\bar{d}$	-1	0.938	1/2
$n$	neutron	$udd$	0	0.940	1/2
$\Lambda$	lambda	$uds$	0	1.116	1/2
$\Omega^-$	omega	$sss$	-1	1.672	3/2



Composite particles are mostly “fields” ... distinction between particles and fields not so simple.

# Recent SM update --- Higgs boson discovery

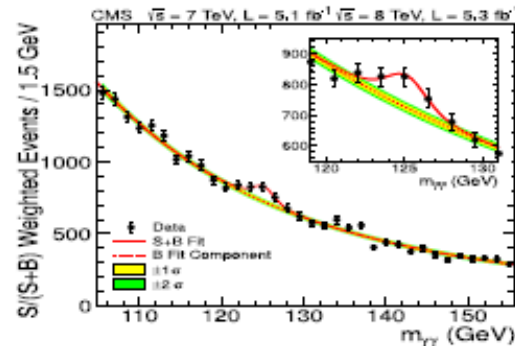
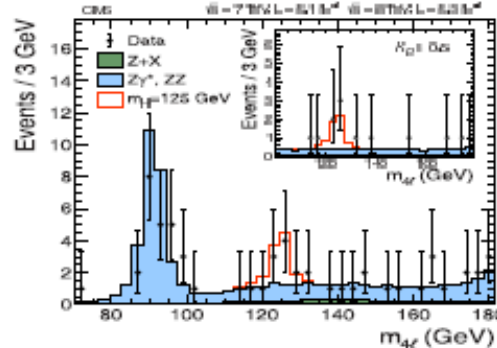
... actually not an update, more of a confirmation ... fully consistent with SM

## July 2012: Observation of a New Boson

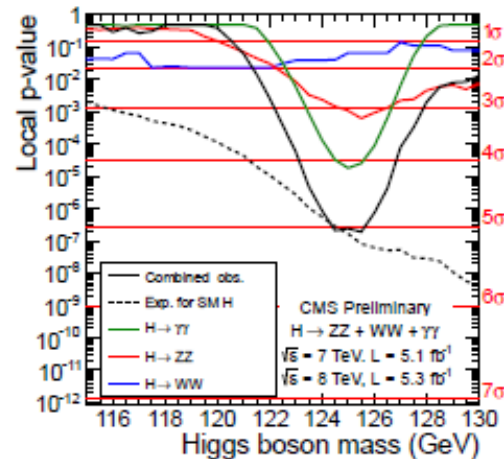
- Observation of a New Boson on CMS:  $5\sigma$  excess

$$X \rightarrow Z(*)Z(*)$$

$$X \rightarrow \gamma\gamma$$

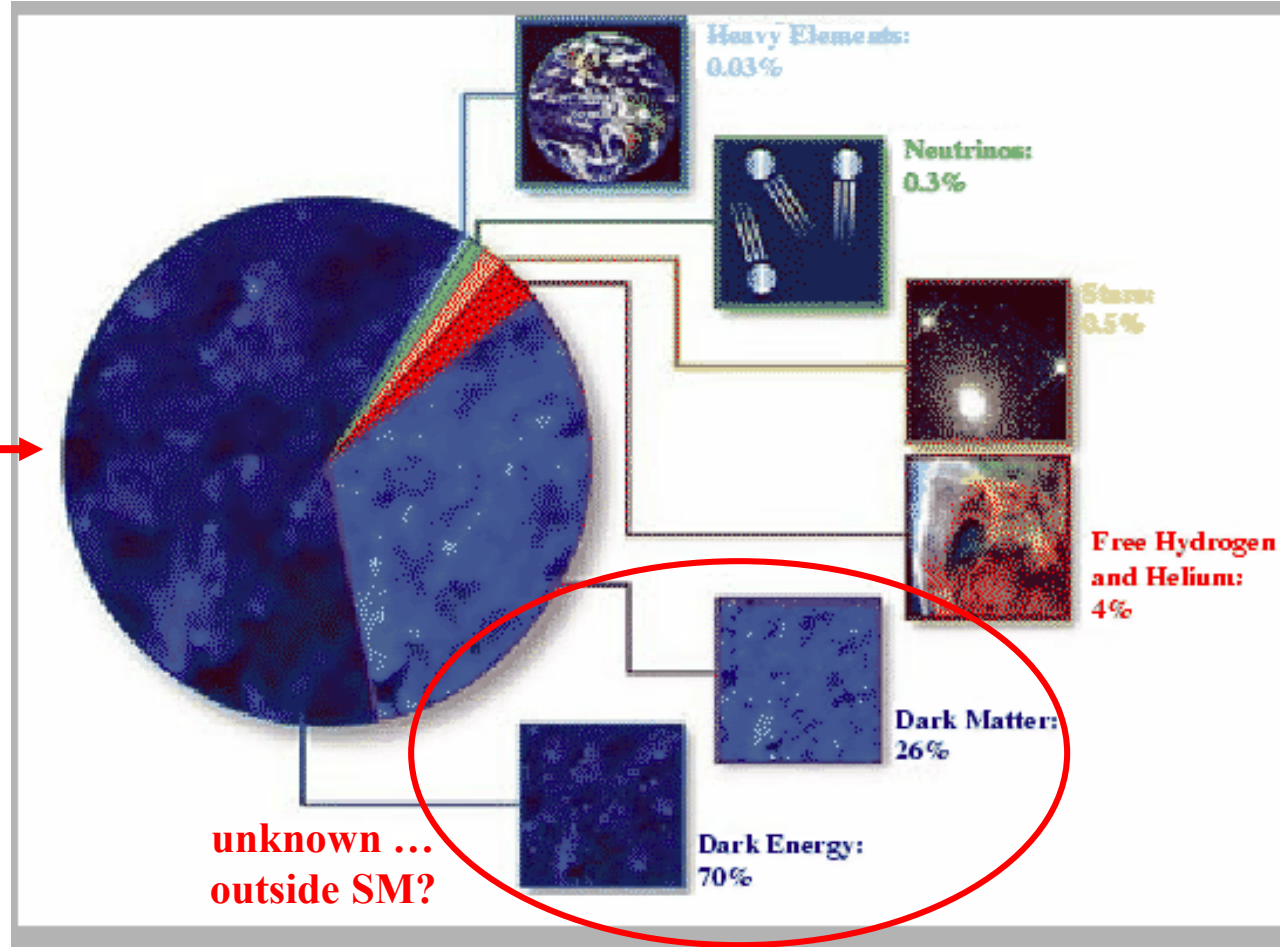
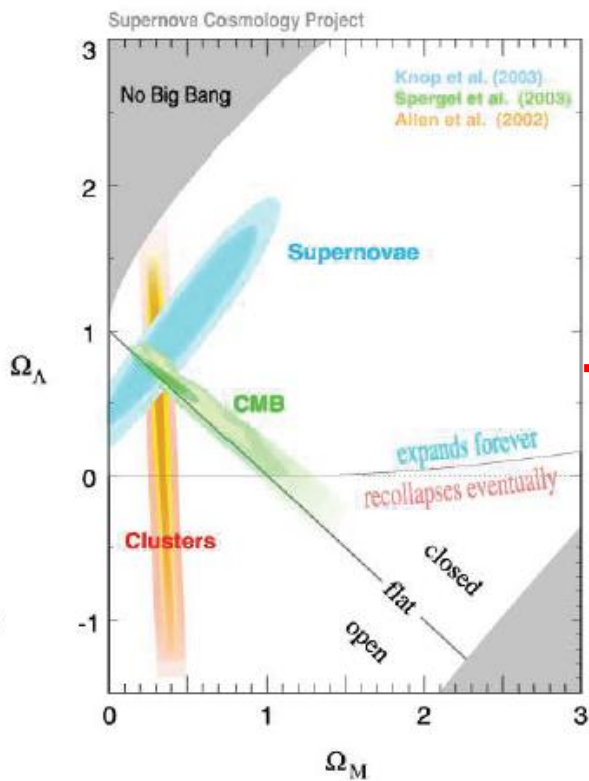


- Probability of background  
 $\sim 0.2 \times 10^{-6}$

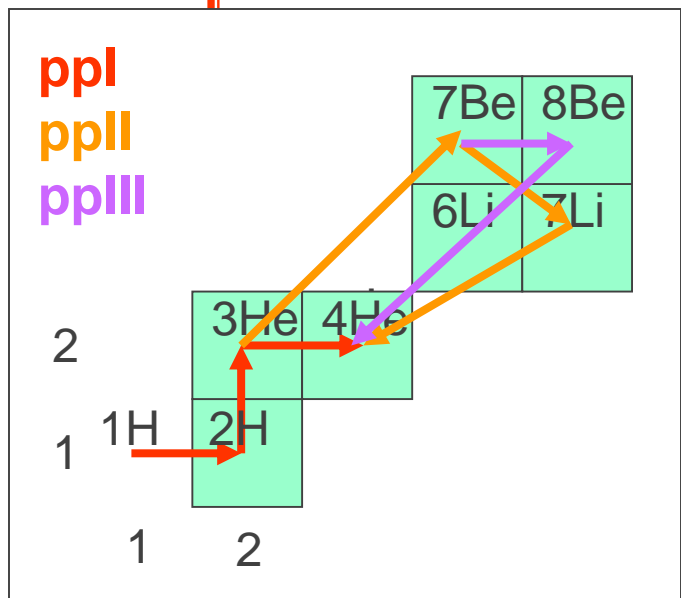
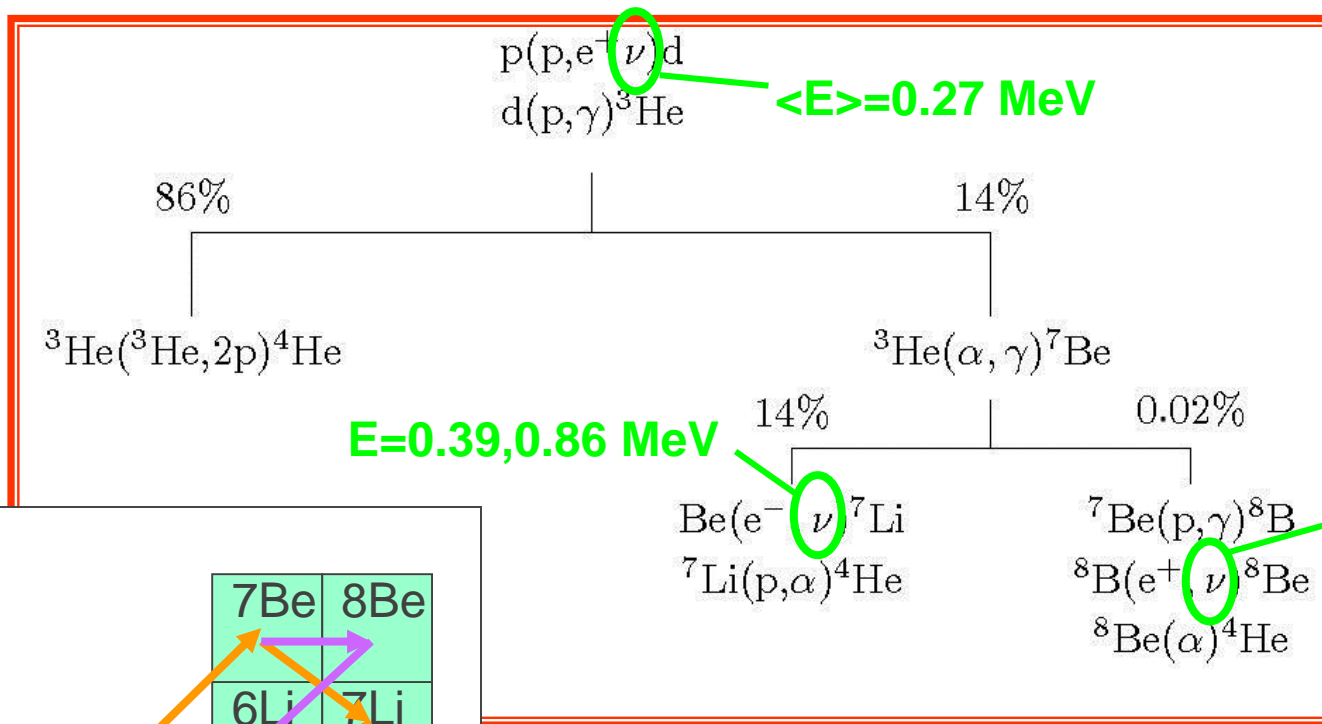


# Recent SM update --- Dark matter and Dark Energy

## The Standard Model only explains 5% of what makes up our universe



# Recent SM update --- Neutrino oscillations: Solar neutrinos

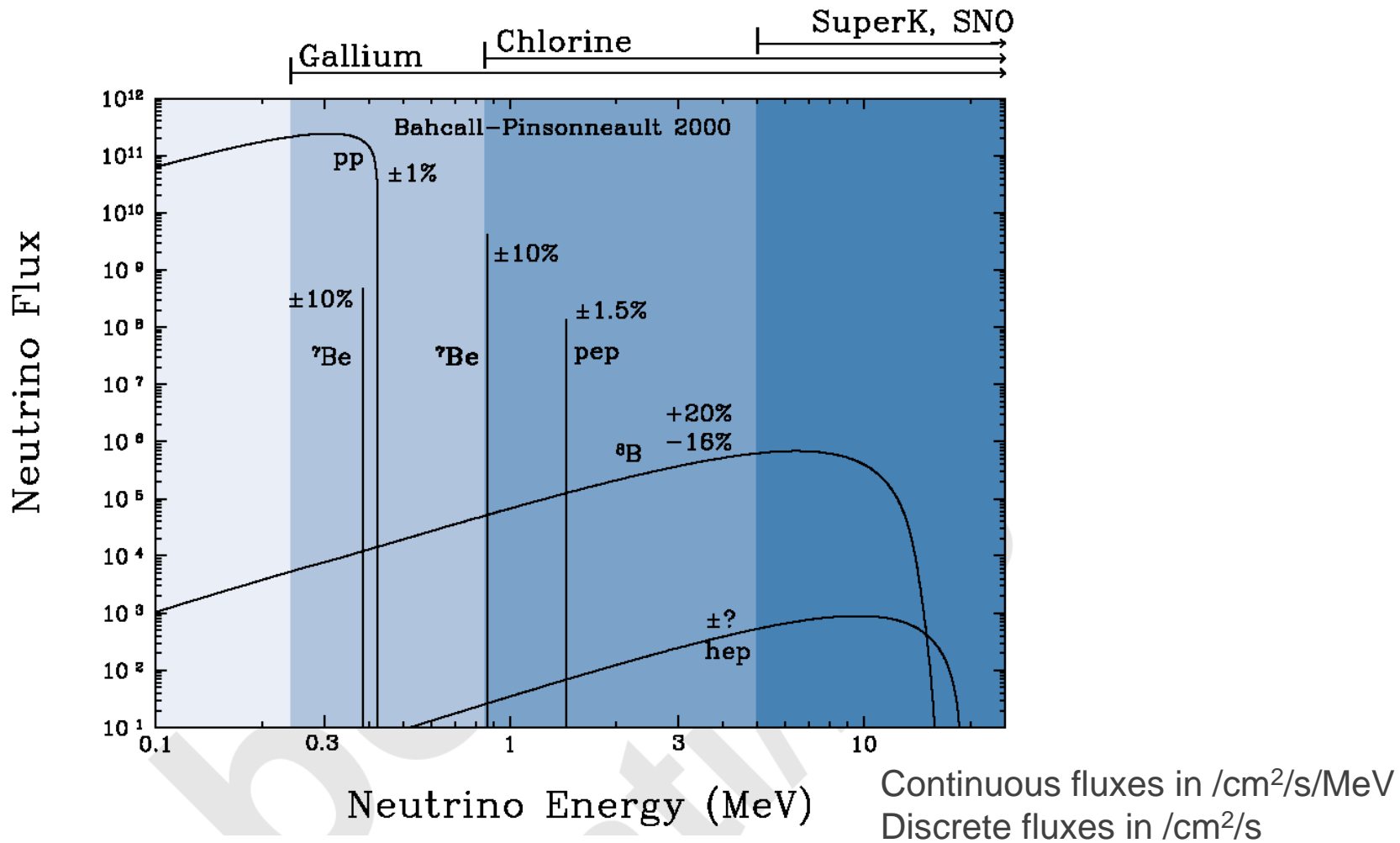


Stellar burning

Total energy lost to neutrinos : 2.3%

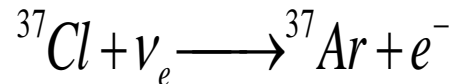


# Recent SM update --- Neutrino oscillations: Solar neutrinos



# Recent SM update --- Neutrino oscillations: Solar neutrino detection

- **1964** John Bahcall and Ray Davis have the idea to detect solar neutrinos using the reaction:



- **1967** Homestake experiment starts taking data

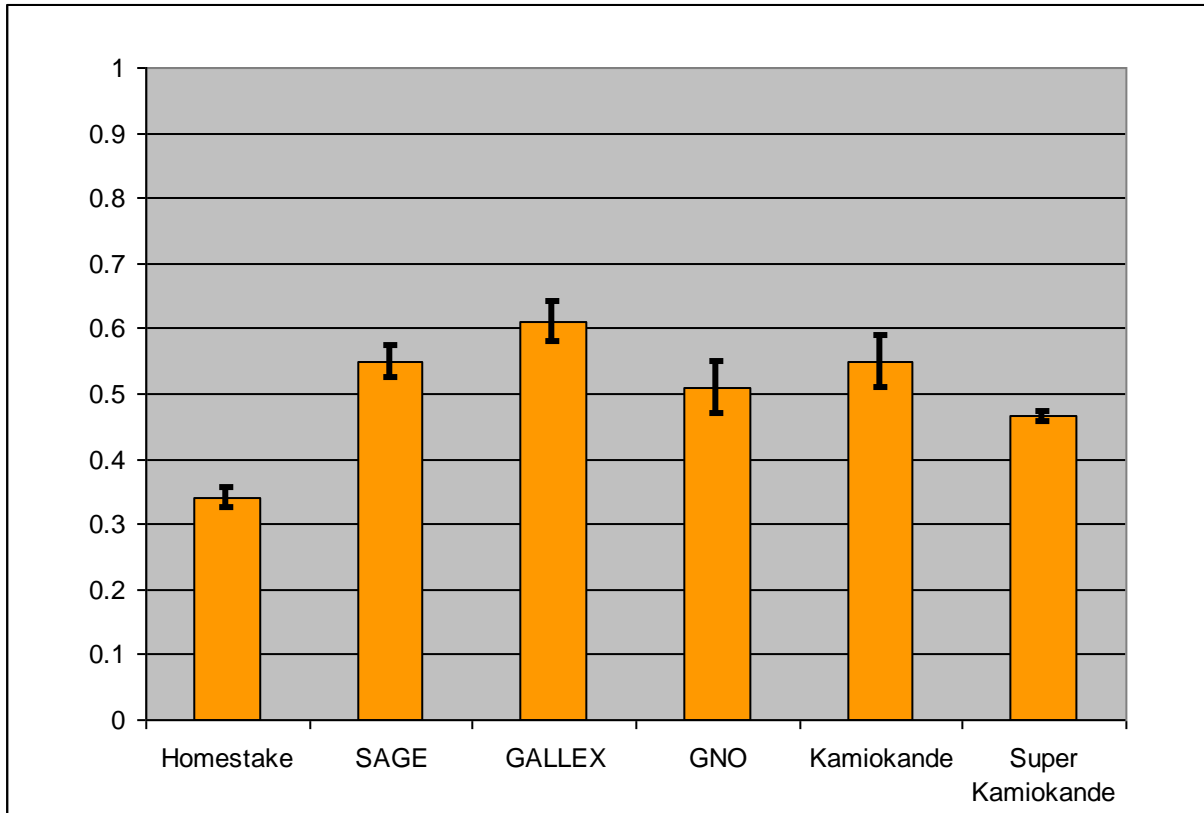
- 100,000 Gallons of cleaning fluid in a tank 4850 feet underground
- ${}^{37}\text{Ar}$  extracted chemically every few months (single atoms !)  
and decay counted in counting station (35 days half-life)
- event rate: ~1 neutrino capture per day !

- **1968** First results: only 34% of predicted neutrino flux !

solar neutrino problem is born - for next 20 years there is no other detector !



# Recent SM update --- Neutrino oscillations: Solar neutrino detection



many more experiments over the years with very different energy thresholds:  
all show a deficit vs standard solar model



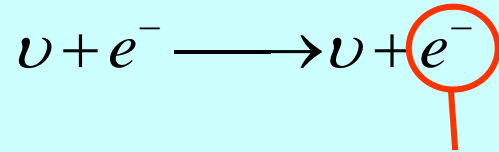
$\nu_e$  only



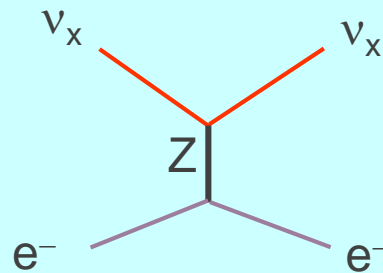
all flavors, but  $\nu_\tau, \nu_\mu$  only 16% of  $\nu_e$  cross section because no CC, only NC

# Recent SM update --- Neutrino oscillations: Solar neutrino detection

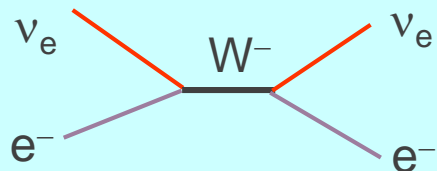
Water Cherenkov detector:



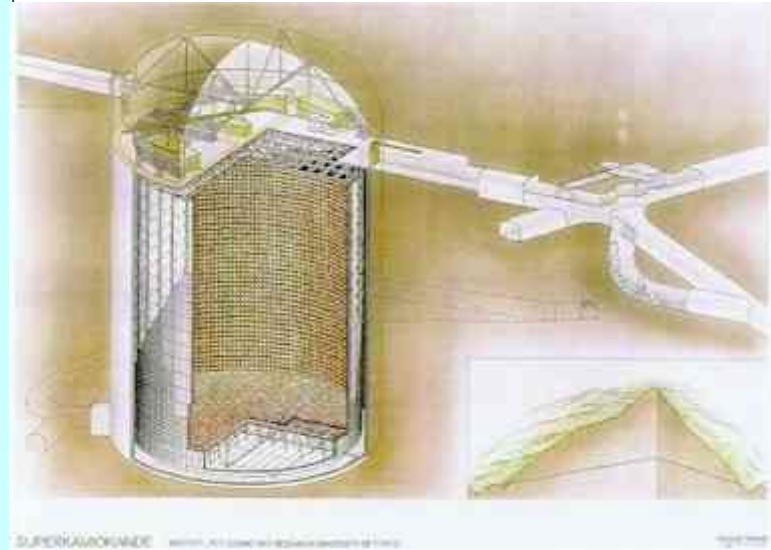
high energy (compared to rest mass)  
- produces Cherenkov radiation when traveling in water (**can get direction**)



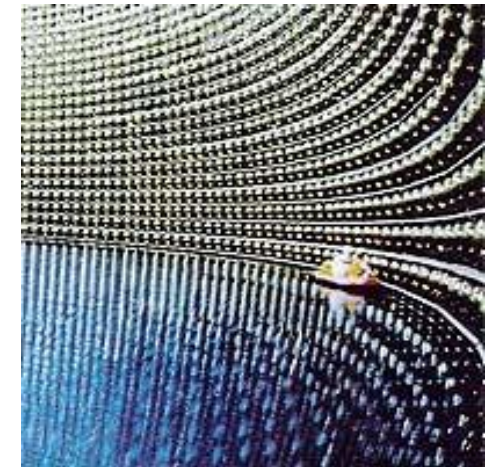
neutral current (NC)



charged current (CC)



Super-Kamiokande Detector

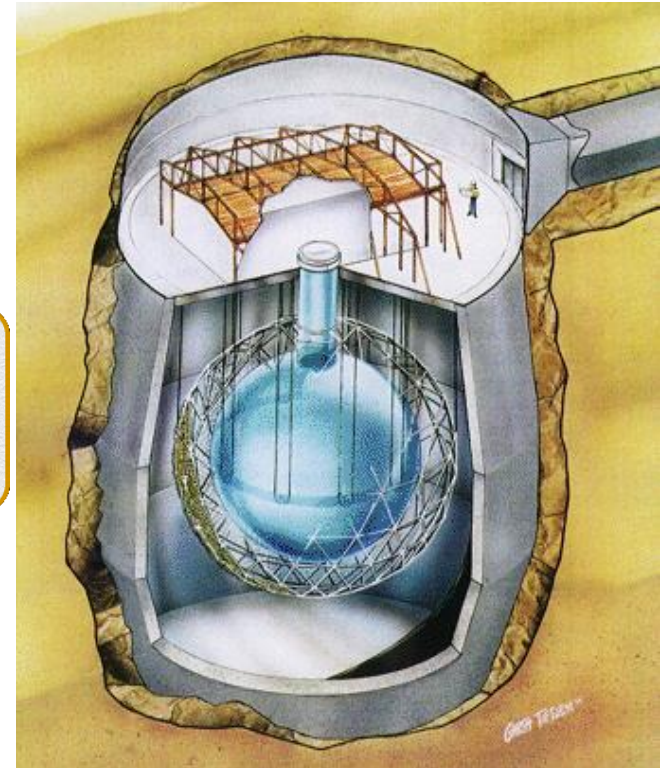
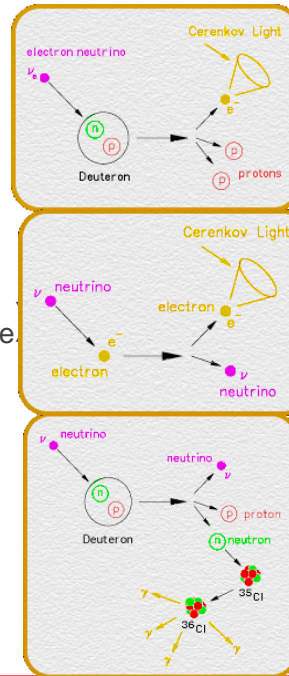
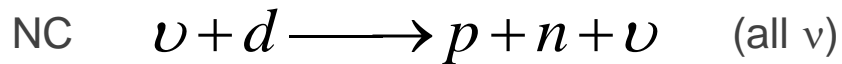
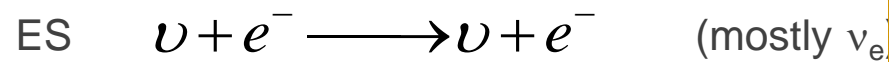
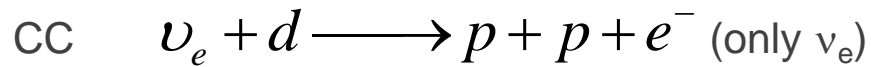


# Recent SM update --- Neutrino oscillations: Solar neutrino detection

**Possible explanation: neutrinos can change flavor while traveling from sun to earth**

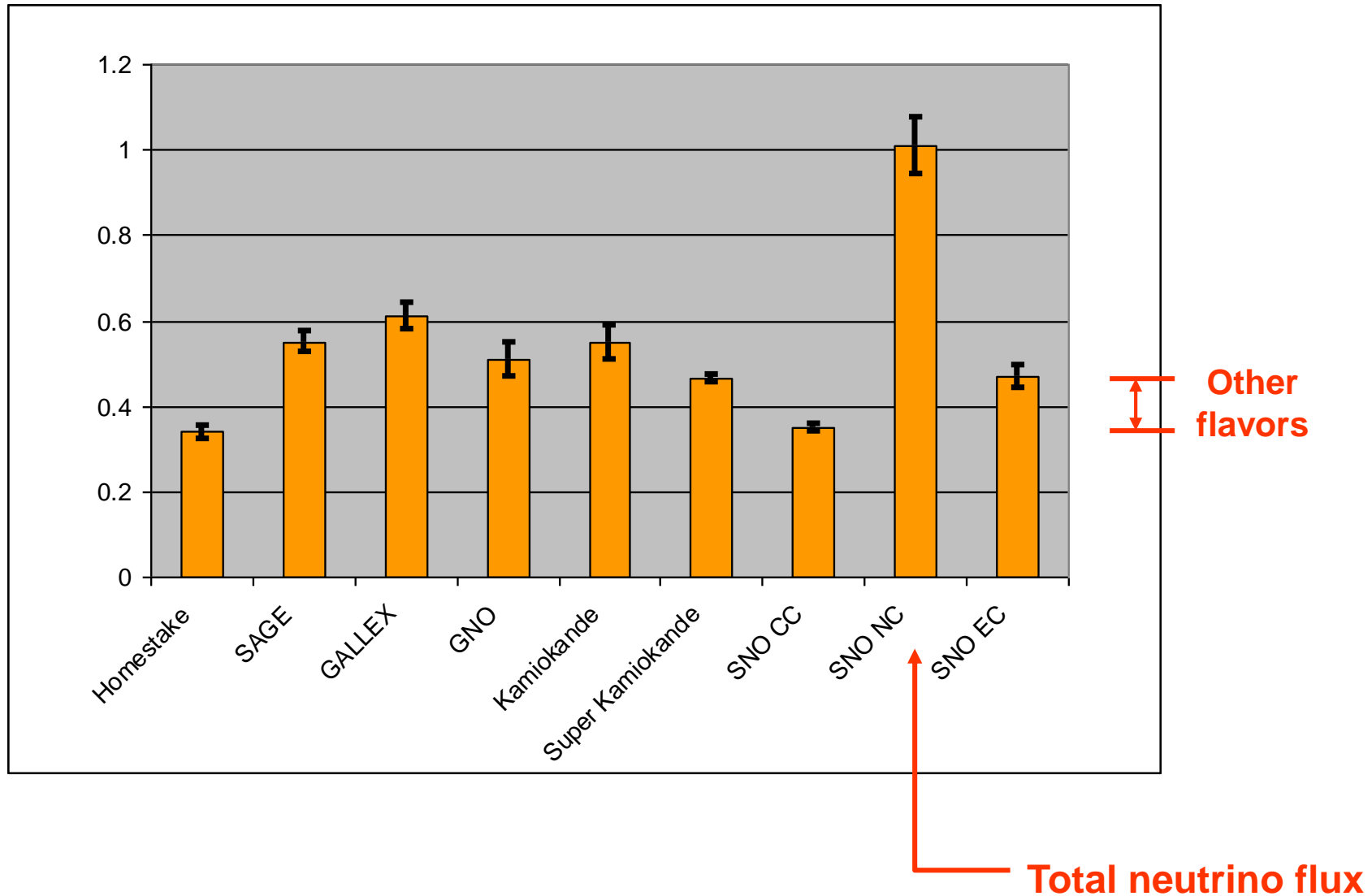
## SNO solar neutrino experiment

three reactions in heavy water:

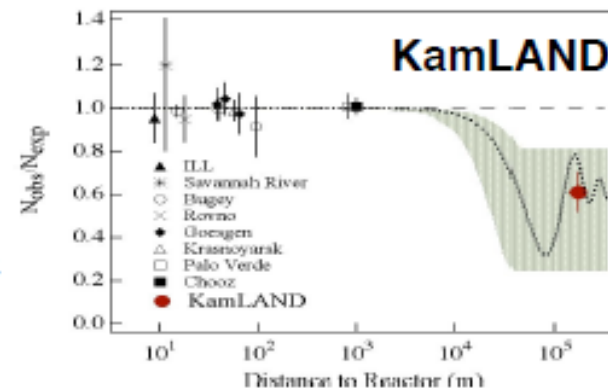
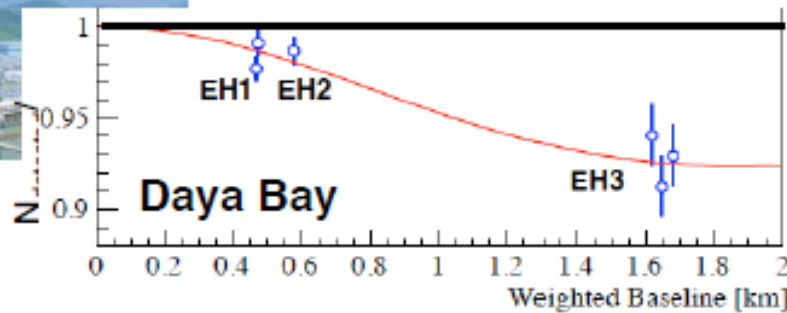
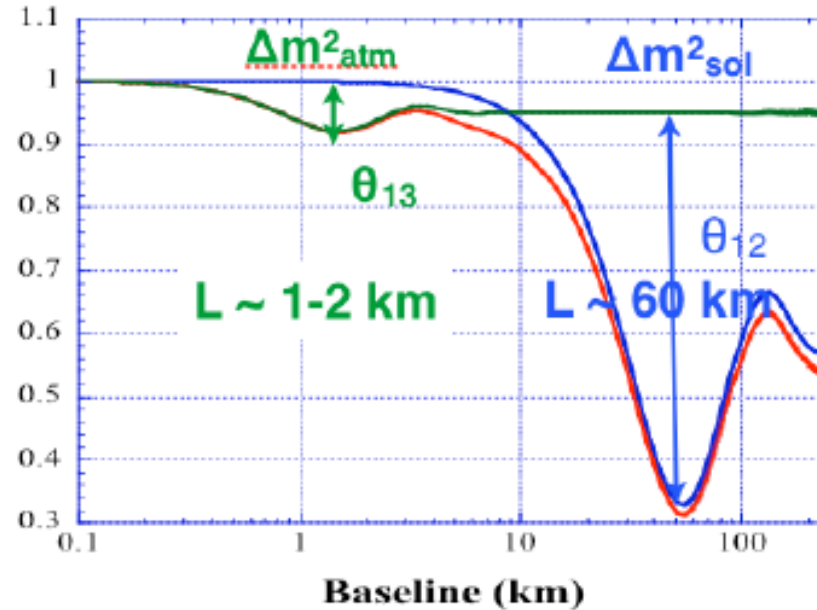
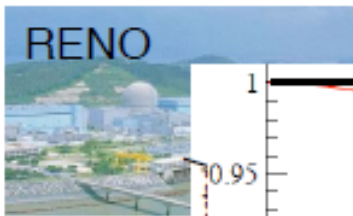


- signature:
- NC independent of flavor - should always equal solar model prediction if oscillations explain the solar neutrino problem
  - Difference between CC and ES indicates additional flavors present

# Recent SM update --- Neutrino oscillations: Solar neutrino detection



# Recent SM update --- Neutrino oscillations: Reactor neutrino



Beginning of era of precision neutrino physics :

$$\sin^2 2\theta_{13} = 0.084^{+0.005}_{-0.005}$$

$$|\Delta m^2_{ee}| = 2.44^{+0.10}_{-0.11} \times 10^{-3} \text{eV}^2$$

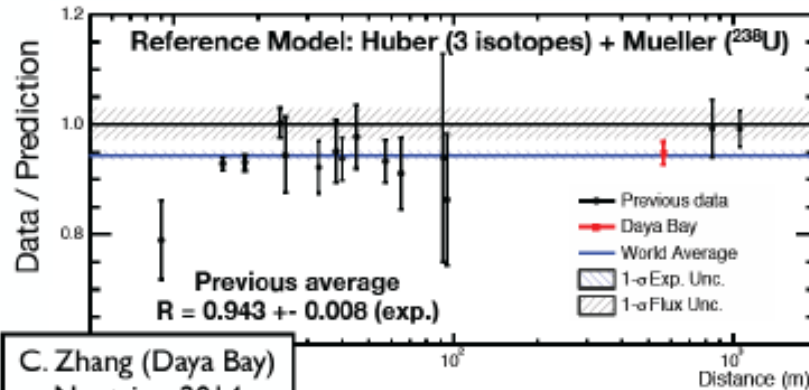
# Recent SM update --- but not all is perfect in neutrino land



## Reactor Flux and Spectrum “Anomalies”

### Flux Deficit

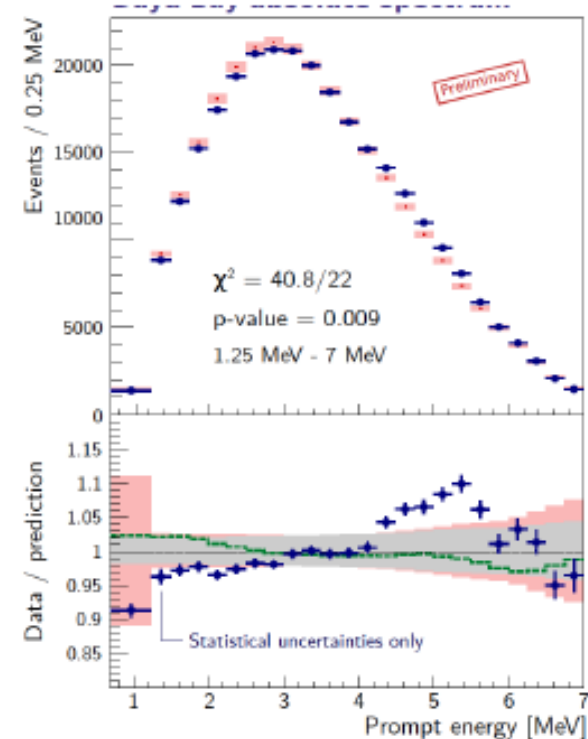
Consistent with previous experiments



Extra neutrino oscillations or artifact of flux predictions?

More data needed to better understand these observations

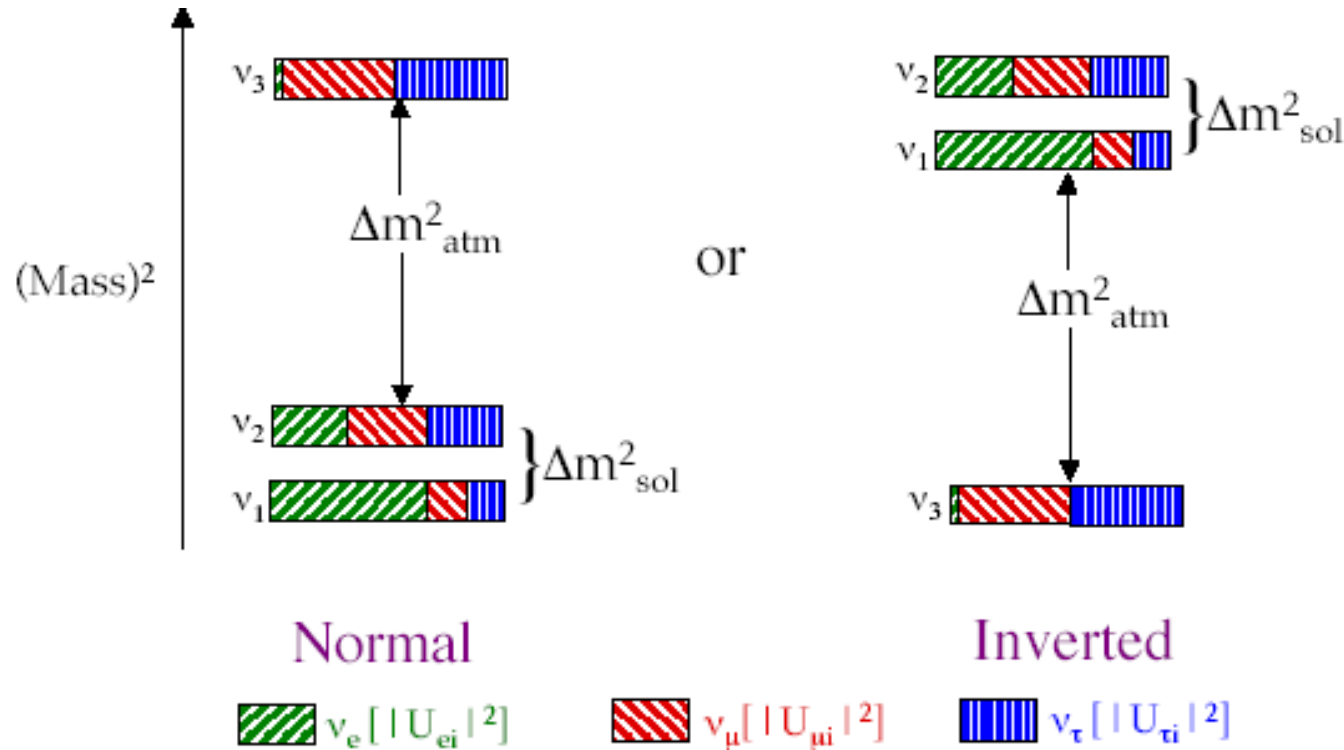
### Spectral Deviation



New feature in 4-6 MeV region of spectrum. Seen by Daya Bay, Double Chooz, and Reno.



# Recent SM update --- Neutrino oscillations: Neutrino mass status



Don't know yet absolute offset for m:

$$\langle m_\beta \rangle = \sqrt{\sum_{i=1}^3 |U_{ei}|^2 m_i^2} < 2.2 \text{ eV} \quad (\text{direct searches from tritium } \beta\text{-decay})$$

$$\sum m_i < (0.4 - 1.0) \text{ eV} \quad (\text{cosmological data+model})$$

# Recent SM update --- Neutrino oscillations: Neutrino nature status

Neutrinos not quite like other leptons:

- $\nu$  mass requires addition of new fields to SM Lagrangian

e.g.  $L \sim m_D \overline{\nu}_L \nu_R$

- $\nu$  mass allows  $\overline{\nu}_i = \nu_i$  (Majorana neutrinos)

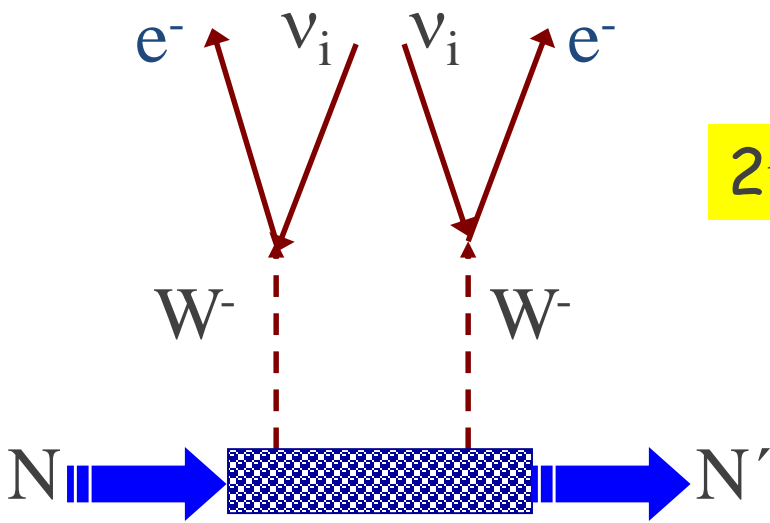
Which in turn allows new CP-violating phases:

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}}_{\text{atmospheric, K2K}} \times \underbrace{\begin{pmatrix} \cos \theta_{13} & 0 & e^{-i\delta_{CP}} \sin \theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin \theta_{13} & 0 & \cos \theta_{13} \end{pmatrix}}_{\text{reactor and accelerator}} \times \underbrace{\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{SNO, solar SK, KamLAND}} \times \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha/2} & 0 \\ 0 & 0 & e^{i\alpha/2+i\beta} \end{pmatrix}}_{\text{Majorana phases}}$$

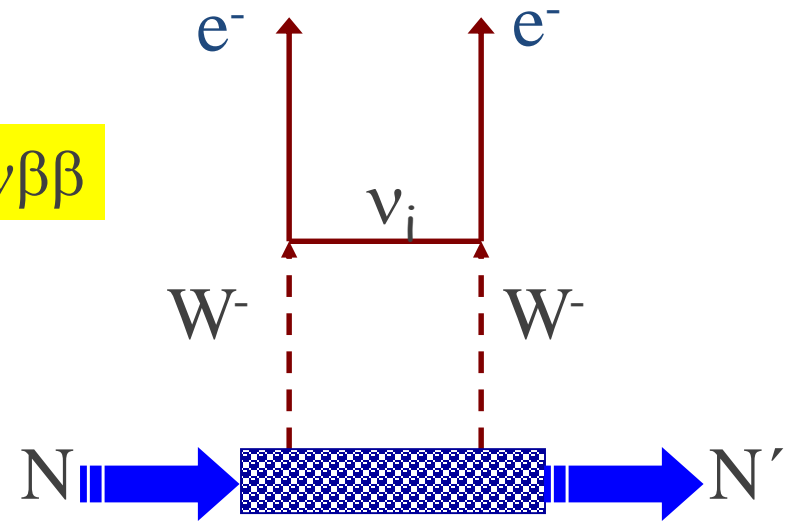
$\sin^2 \theta_{23}$ 
 $\sin^2 \theta_{13}$ 
 $\sin^2 \theta_{12}$ 
 $0\nu\beta\beta$



# Recent SM update --- Neutrino oscillations: Testing neutrino nature (Majorana vs Dirac)



$2\nu\beta\beta$  vs.  $0\nu\beta\beta$



$$\Gamma_{2\nu} = G_{2\nu} |M_{2\nu}|^2$$

$$\langle m_\beta \rangle = \sqrt{\sum_{i=1}^3 |U_{ei}|^2 m_i^2}$$

observed      not seen yet

Nuclear matrix element (~ OK)  
Phase space (OK)

$$T_{1/2} \propto m_\nu^2$$

$$\Gamma_{0\nu} = G_{0\nu} |M_{0\nu}|^2 m_\nu^2$$

Mass is mixed average

$$\langle m_{\beta\beta} \rangle = \sum_{i=1}^3 |U_{ei}|^2 m_i \epsilon_i$$



# Recent SM update --- Neutrino oscillations: Testing neutrino nature

Cosmology  $\Sigma = m_1 + m_2 + m_3$   $\delta m_{ij}^2 = m_j^2 - m_i^2$  Oscillations

Beta decay  $\langle m_\beta \rangle = \sqrt{\sum_{i=1}^3 |U_{ei}|^2 m_i^2}$   $\langle m_{\beta\beta} \rangle = \sum_{i=1}^3 |U_{ei}|^2 m_i \epsilon_i$   $\beta\beta$

	<b>Absolute Mass Scale</b>	<b>Relative Mass Scale</b>	<b>Mixing Matrix Elements</b>	<b>CP nature of <math>\nu</math></b>
$\beta\beta$	✓			✓
$\beta$ , cosm	✓			
<b>Oscil.</b>		✓	✓	

# Standard Model as it stands now

The Standard Model has been very successful at describing the world around us in terms of its basic constituents and interactions.

Recent changes have however made the number of free parameters in the SM grow to at least 26:

- 12 Fermion masses:  $m_e, m_\mu, m_\tau$ , neutrino masses, quark masses
- 2 Gauge boson masses:  $M_Z$  and  $M_W$  (2 parameters)
- 2 Coupling constants:  $\alpha_{\text{strong}}, \alpha_{\text{electroweak}}$
- 2 Vacuum energies: Higgs mass,  $\theta_{\text{QCD}}$
- 4 quark mixing angles ( $\theta_{12}, \theta_{13}, \theta_{23}, \delta_{\text{CP}}$ ) Cabibbo-Kobayashi-Maskawa
- 4 lepton mixing angles ( $\theta_{e\tau}, \theta_{e\mu}, \theta_{\mu\tau}, \delta$ ) (neutrino oscillations, etc.)

more if the neutrinos are Majorana particles or if sterile neutrinos. It does not tell us what dark energy is and cannot accommodate us being here (not enough CP violation to explain the dominance of matter over antimatter).

A more complete (and elegant) scheme must exist.

# Low energy tests of the Standard Model: accurate physics in an inaccurate system

The nucleus is a complex quantum system **not generally amenable to an exact description.**

However, by a proper choice of nuclear system and observable, **specific physical processes can be isolated** and determined to high precision.

e.g.:  $0^+$  to  $0^+$  superallowed decays or PNC in atoms

Requirements are:

- An identifiable and separable **observable** (e.g.: a P-violating signal in an otherwise P-conserving experiment)
- A suitable **laboratory** (nucleus where the specific observable is enhanced and unperturbed)
- The proper **experimental tools**

# Testing the Standard Model at low energy: use the symmetries

For  $\mathbf{H}$  an Hamiltonian obeying a symmetry  $\mathbf{S}$  :

$$[\mathbf{H}, \mathbf{S}] = 0$$

If  $|\Psi\rangle$  is an eigenstate of  $\mathbf{H}$  with eigenvalue  $E$ , then

$$(\mathbf{H}\mathbf{S} - \mathbf{S}\mathbf{H}) |\Psi\rangle = 0 \quad \rightarrow \quad \mathbf{H} (\mathbf{S} |\Psi\rangle) = E (\mathbf{S} |\Psi\rangle)$$

Thus  $\mathbf{S}|\Psi\rangle$  is an eigenstate with the same energy  $E$  and if the spectrum is nondegenerate,  $\mathbf{S}|\Psi\rangle = s |\Psi\rangle$  with  $s$  a number.

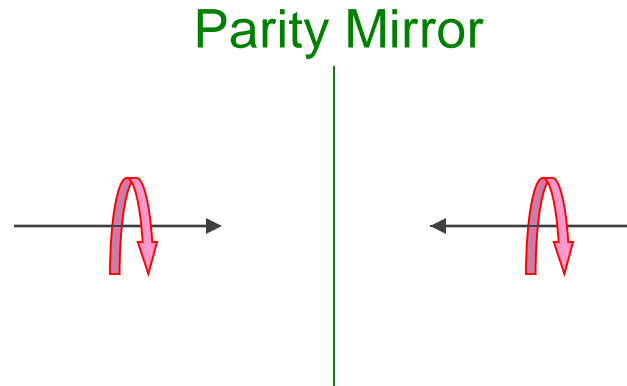
$\rightarrow |\Psi\rangle$  is an eigenstate of  $\mathbf{S}$  if  $\mathbf{H}$  obeys the symmetry  $\mathbf{S}$ .

**You can learn about the properties of the interactions by studying the wavefunction properties.**

# Symmetry example : Parity

Under the parity operation:

- $r \rightarrow -r$
- $p \rightarrow -p$
- $r \times p \rightarrow r \times p$
- $J \rightarrow J$



If you call the two states  $|\Psi_R\rangle$  and  $|\Psi_L\rangle$ , neither is an eigenstate of an inversion symmetric  $H$ . The proper states are:

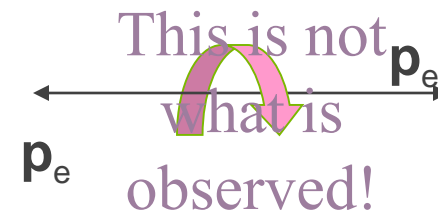
$$|\Psi_S\rangle = (2)^{-1/2} (|\Psi_L\rangle + |\Psi_R\rangle)$$

$$|\Psi_A\rangle = (2)^{-1/2} (|\Psi_L\rangle - |\Psi_R\rangle)$$

Madame Wu's experiment:

Polarize  $^{60}\text{Co}$  and look at the direction of the emitted  $\beta$ 's.

In a **parity-symmetric world** we would see as many electrons emitted in the direction of  $\mathbf{J}$  as opposite  $\mathbf{J}$ .





# Other useful fundamental symmetries

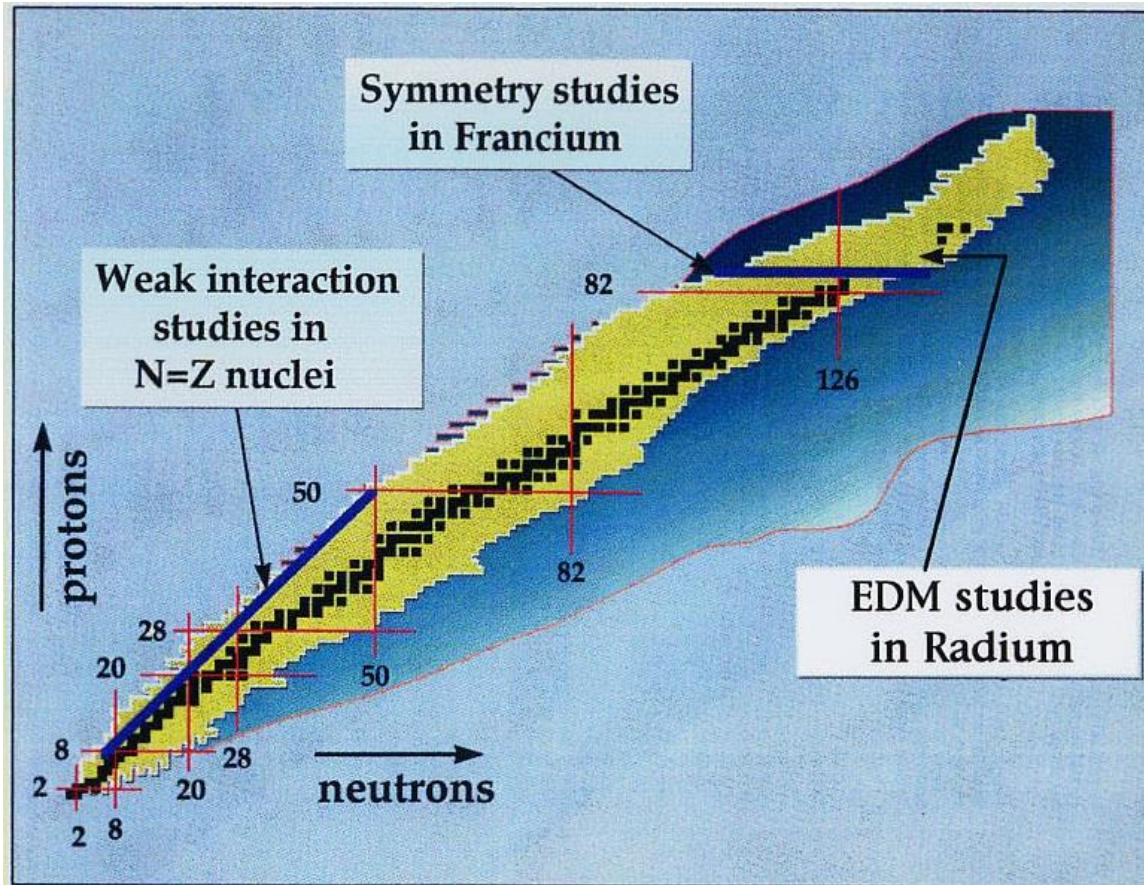
Other useful fundamental symmetries:

- P (space reversal)
- T (time reversal)
- C (charge conjugation .... particle  $\leftrightarrow$  antiparticle )
- CP ... must be broken for us to exist
- **CPT** ... this one must be conserved ... Lorentz invariance
- ...





# Testing fundamental symmetries at low energy: the laboratory



Specific nuclei have advantages: good isospin symmetry, enhanced effects because of relativity or deformation ... just by statistics, you would expect most of the best candidates will be radioactive.

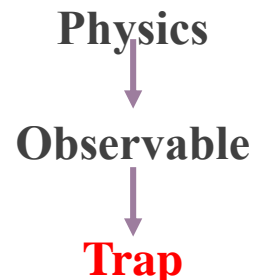
# Testing fundamental symmetries at low energy: traps ... the precision tools

Traps allow the confinement of particles in a well-controlled environment, free from outside perturbations, where they are available for precision measurements.

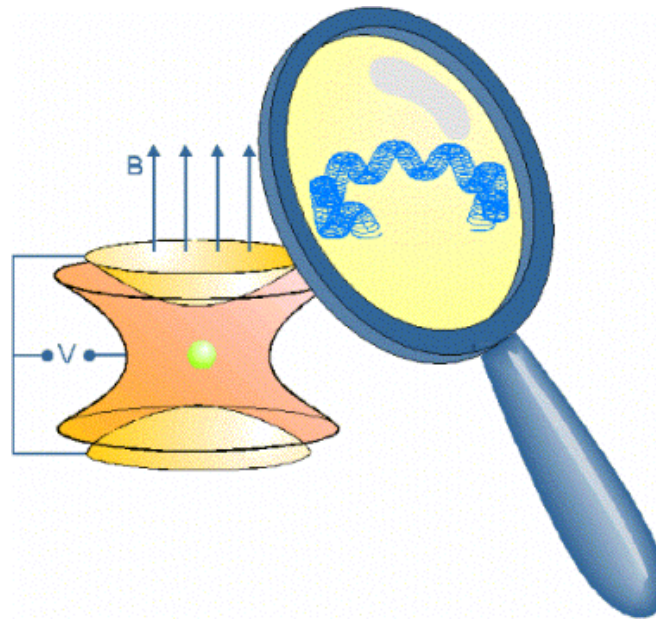
A large variety of traps is available:

- Penning trap
- Optical trap
- Magnetic trap
- Dipole trap
- FORT trap
- Radio-frequency trap
- Magneto-optical trap
- Cyclotron trap
- ...

Conception of experiment:



# Measurements in traps

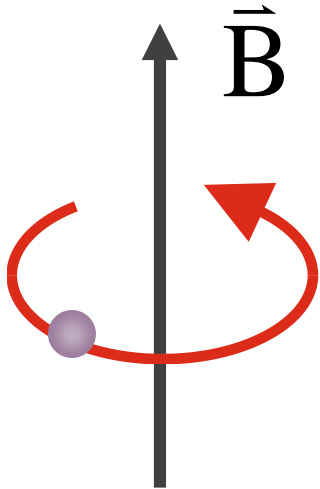


**Traps can be used as a storage device where the intrinsic properties of the confined particle can be studied under optimal conditions**

**- or -**

**Information on the captured species can be obtained from the eigenmodes of the trapped particles in a precision trap**

# Most simple trap: 2d-confinement of charged particle in a magnetic field

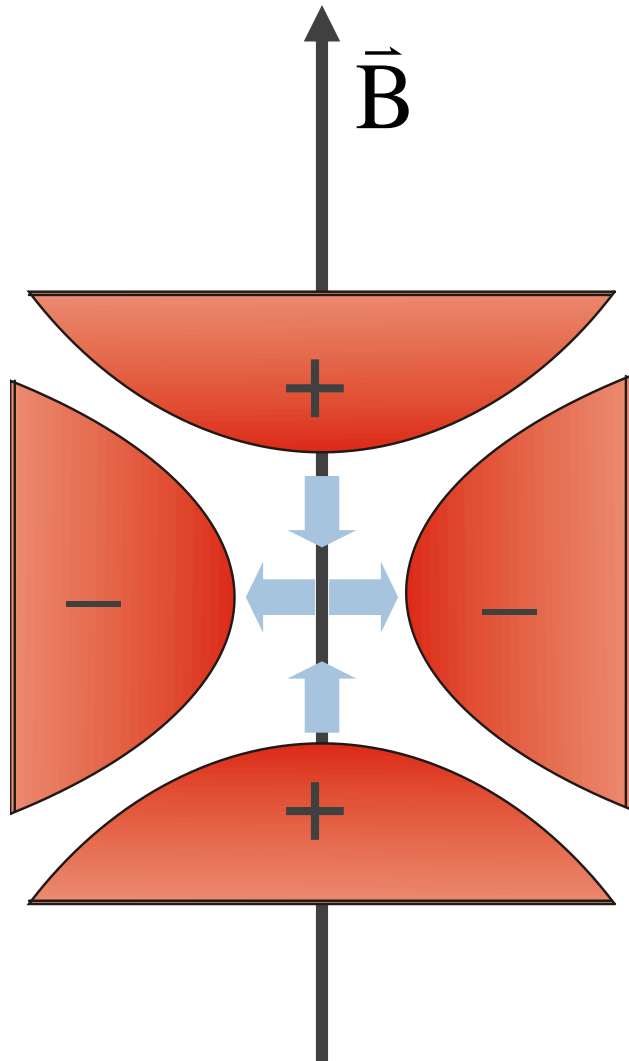


- Constant axial magnetic field
- particle orbits in horizontal plane

$$\omega_c = \frac{qB}{m}$$

- free to escape axially

# Most simple trap: 3d-confinement in the Penning trap



- Add an axial harmonic electric field to confine particles

- axial oscillations:

$$\omega_z = \sqrt{\frac{eV}{md^2}}$$

- Radial motion split into two components by electric field:

- $\omega_+$  : reduced cyclotron freq.

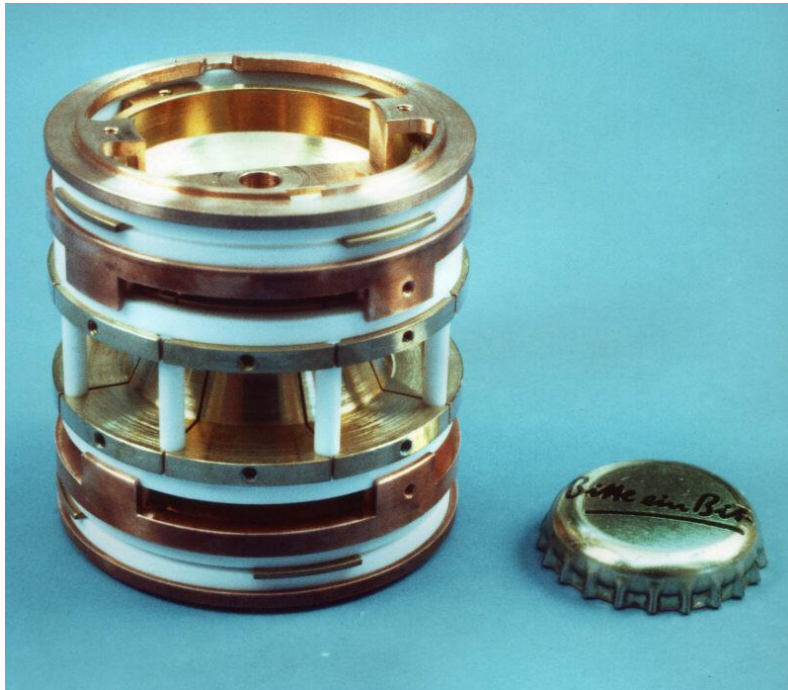
- $\omega_-$  : magnetron frequency

with:

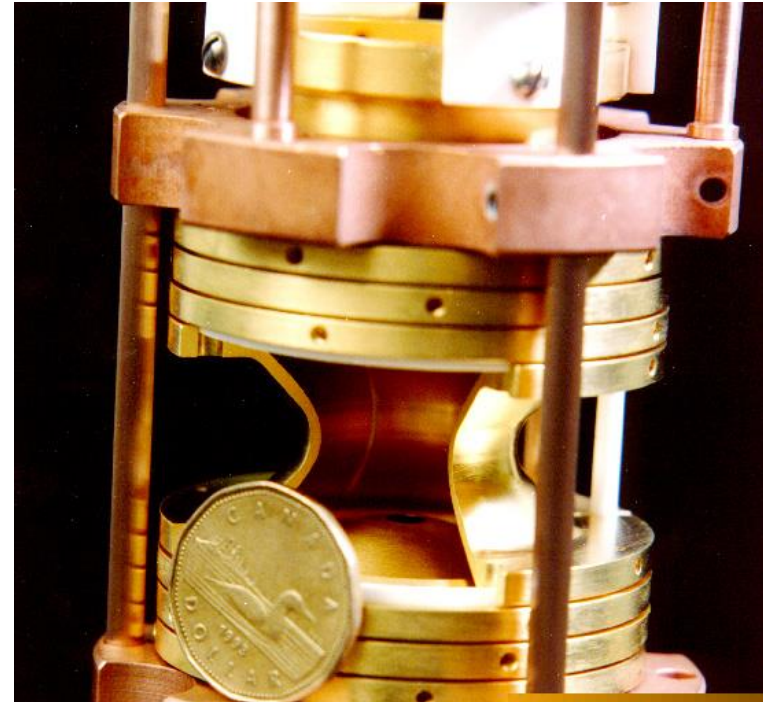
$$\omega_+ + \omega_- = \omega_c$$

**Can trap any charged species!**

# Early on-line Penning traps



ISOLTRAP at CERN



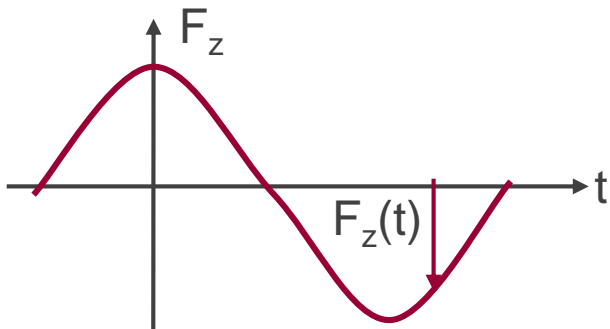
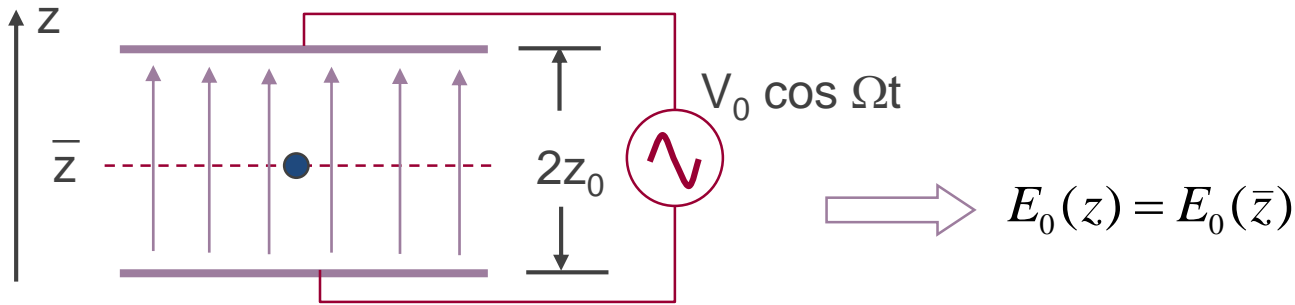
CPT at Argonne

... now a lot more of them



# RF focusing(1)

Uniform  
RF field



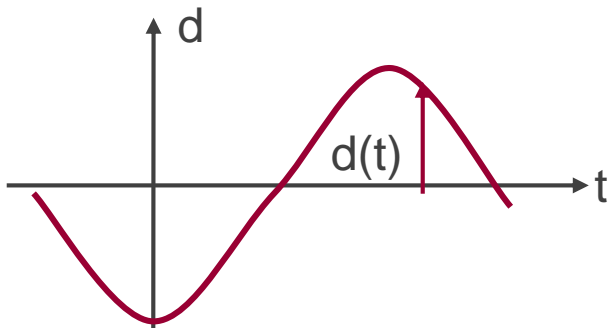
Basic harmonic motion

$$m\ddot{z} = F_z(t) = eE_0(z) \cos \Omega t$$

$$z(t) = \bar{z} + d(t)$$

$$d(t) = -d_0 \cos \Omega t$$

$$d_0 = \frac{eE_0(\bar{z})}{m\Omega^2}$$

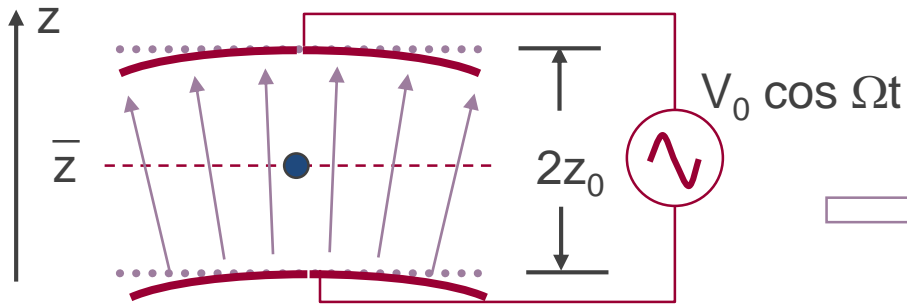


$$\langle F_z(t) \rangle_{av} = eE_0(\bar{z}) \langle \cos \Omega t \rangle = 0$$

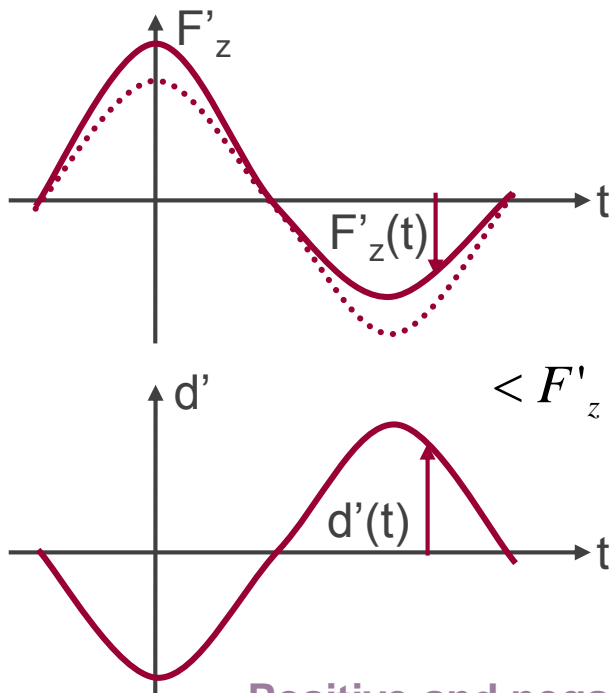
No net force when averaged over full RF cycle!

# RF focusing(2)

Non-Uniform  
RF field



$$\Rightarrow E_0(z) = E_0(\bar{z}) + \left( \frac{\partial E_0(\bar{z})}{\partial \bar{z}} \right) d$$



$$m\ddot{z} = F'_z(t) = eE_0(z) \cos \Omega t$$

$$z(t) = \bar{z} + d'(t)$$

$$d'(t) \cong d(t) = -d_0 \cos \Omega t$$

$$d_0 = \frac{eE_0(\bar{z})}{m\Omega^2}$$

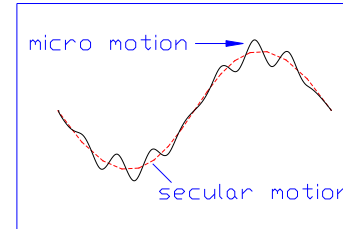
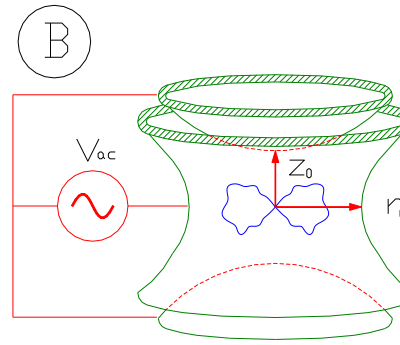
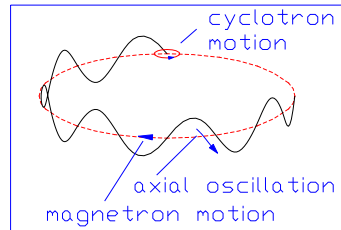
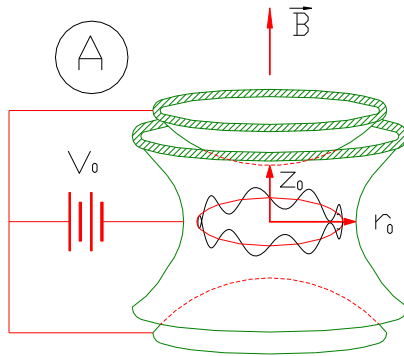
$$\langle F'_z(t) \rangle_{av} = e \left( \frac{\partial E_0(\bar{z})}{\partial \bar{z}} \right) \langle d' \cos \Omega t \rangle = \frac{-e^2 E_0(\bar{z})}{2m\Omega^2} \left( \frac{\partial E_0(\bar{z})}{\partial \bar{z}} \right)$$

$$F' = -e \frac{\partial V_{ps}}{\partial z} \Rightarrow V_{ps} = \frac{eE_0^2(z)}{4m\Omega^2}$$

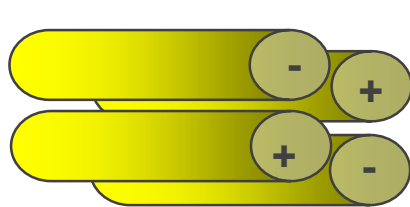
Positive and negative charges attracted to regions of lower RF amplitude!

# Guiding/trapping structures

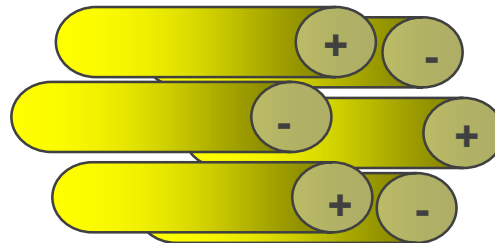
Penning Trap



RFQ Trap

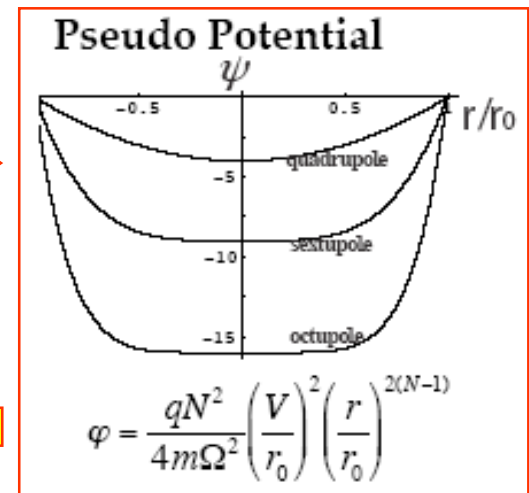
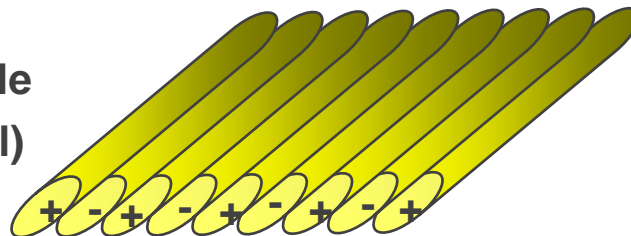


RF quadrupole



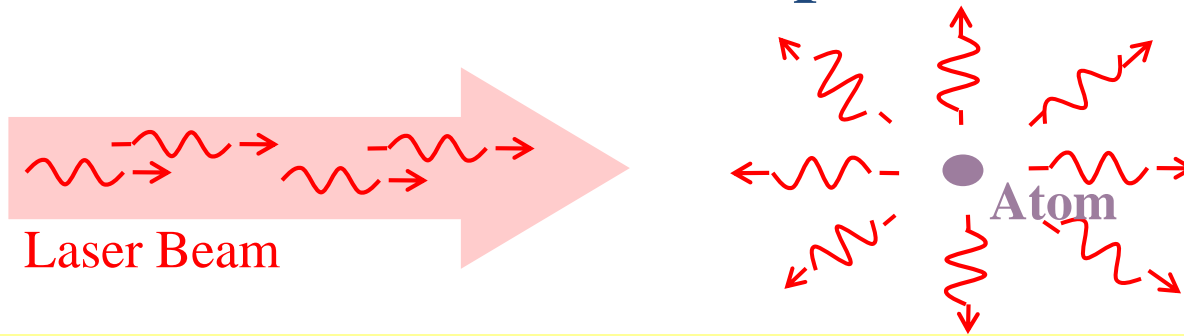
RF hexapole

Centipole (RF wall)



# Optical traps: spontaneous scattering light force

## ... resonance and repetition

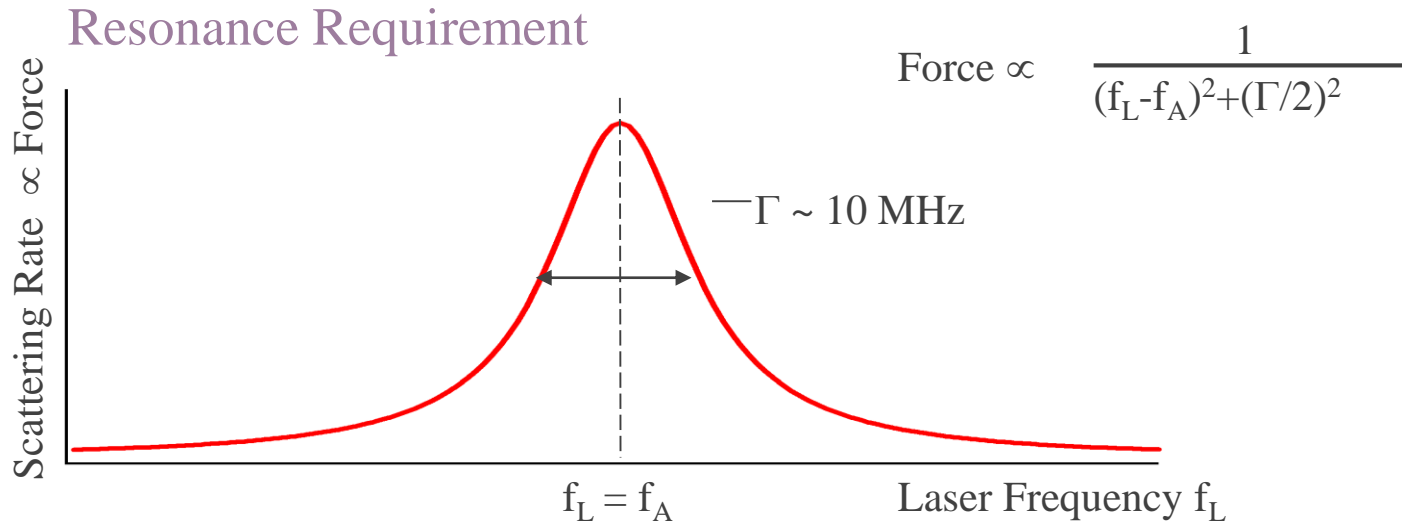


Krypton Atom:

$1s_5 \rightarrow 2p_9$  wavelength = 811 nm

Single photon kick  $\delta v = 6$  mm/sec

Transition rate  $\sim 1 \times 10^7$  /sec Acceleration  $\sim 6 \times 10^4$  m/sec<sup>2</sup>

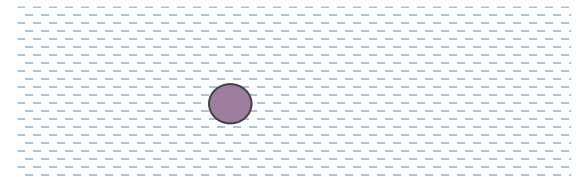
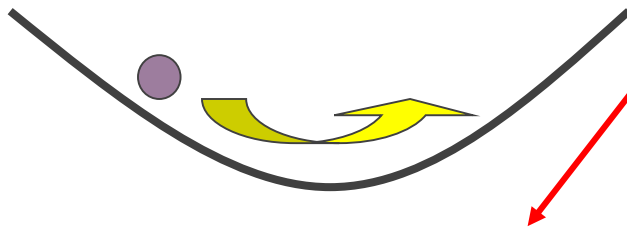


# Optical traps: requirements

**Trapping**  
 $F = -kx$

*position  
dependent*

**Cooling**  
 $F = -av$



Magnetic Field  $B(x)$

Atom Velocity



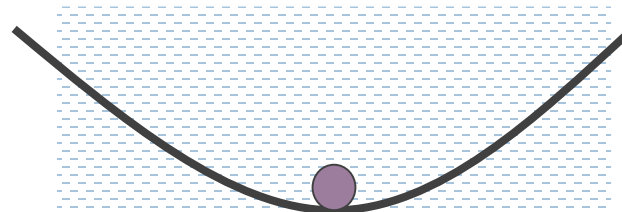
$f_A(x)$

$f_L(v)$

Zeeman Shift

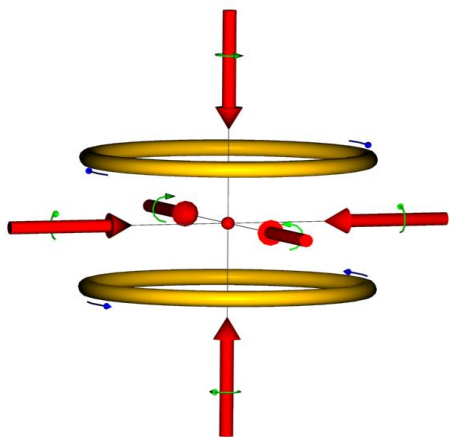
Doppler Shift

**A Trap with Cooling**



# Magneto-Optical Trap (MOT)

Raab, Prentiss, Cable, Chu, Pritchard, Bell Lab & MIT, 1987



## Ingredients:

Laser beams --- alignment, frequency, polarization;

Quadrupole B-field --- 20 G/cm, anti-Helmholtz;

Ultra-high vacuum ---  $\tau_{\text{trap}} \sim 1 \text{ sec @ } 1 \times 10^{-8} \text{ Torr.}$

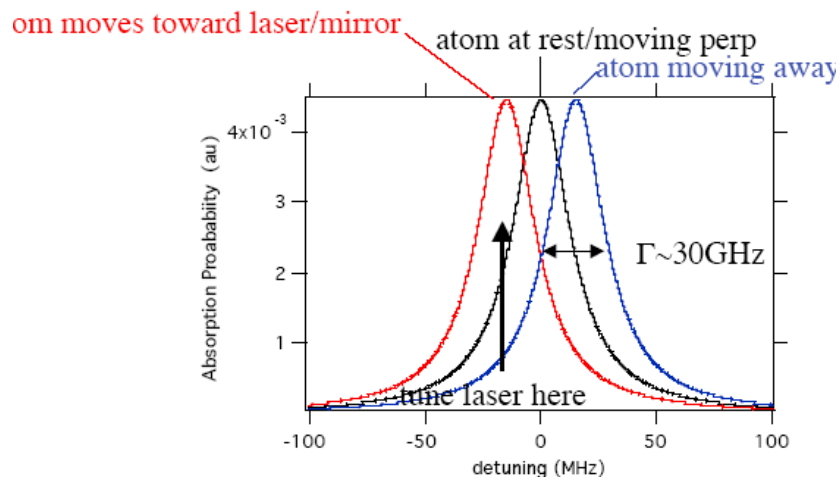
## Typical Parameters:

Number ---  $10^{10}$ ;

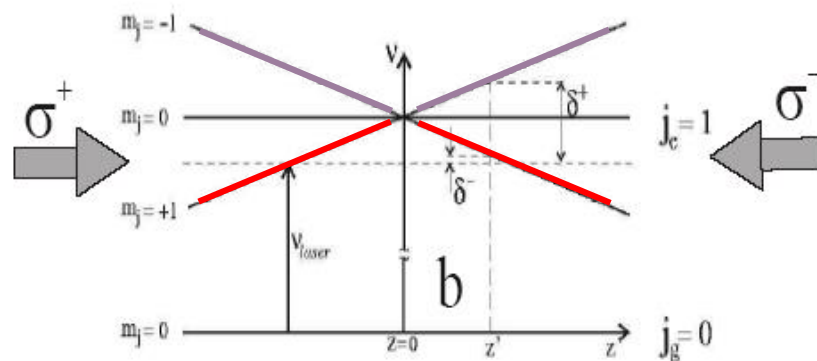
Density ---  $10^8 \text{ mm}^{-3}$ ;

Temperature ---  $\sim \text{mK}, < 1 \text{ m/sec}$ ;

Capture speed --- 20 m/sec.



$$R = \int d\sigma(\nu) \Phi_{\nu}(\nu) d\nu$$



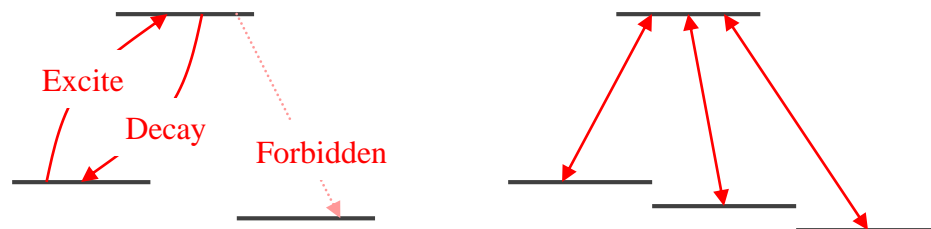
# Trappable Atoms

## Transition Requirements:

Cycling transition

High transition rate ( $10^7 \text{ sec}^{-1}$ , allowed E1)

Practical wavelength ( $\lambda > 200 \text{ nm}$ )



1 H																	2 He	
3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne	
11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	
55 Cs	56 Ba	57 La	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn	
87 Fr	88 Ra	89 Ac	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Uun	111 Uuu	112 Uub		114 Uuq		116 Uuh		118 Uuo	
			58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu		
			90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr		

36  
Kr Demonstrated

86  
Rn Trappable



# Two-sigma user's guide to traps

- **Penning trap**

- Trapping potential  $\sim 1-1000$  eV
- Universal
- Precision device
- Space-charge limitation
- Expensive

- **Radio-frequency trap**

- Trapping potential  $\sim 1-1000$  eV
- Universal
- Storage device
- Space-charge limitation
- Inexpensive

- **Magneto-optical trap**

- Trapping potential  $\sim \mu\text{eV}$
- Alkali with good efficiency, others with more difficulty
- Not a precision device, but a **great cooler**
- No space-charge limitation
- Expensive

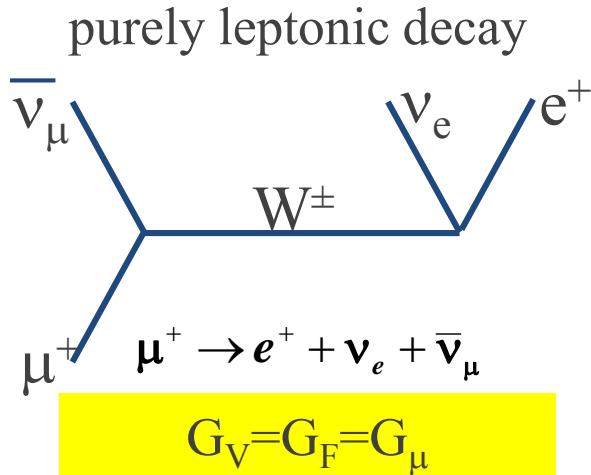


# Example of low energy “opportunities”

- Detailed look at present and planned work
  - Superaligned Fermi decay
  - Angular correlations in  $\beta$ -decay
  - PNC in atoms
  - EDM
  - others ...

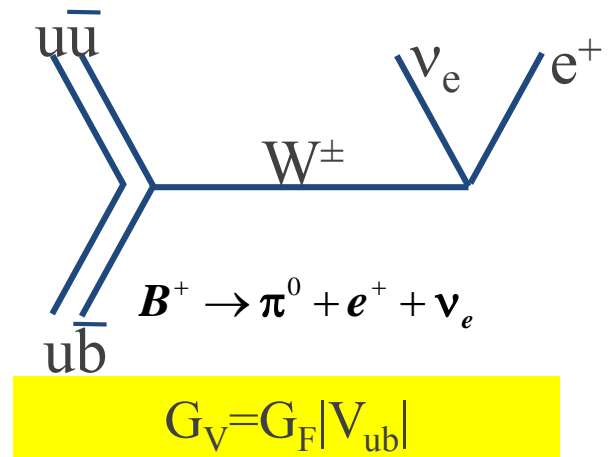
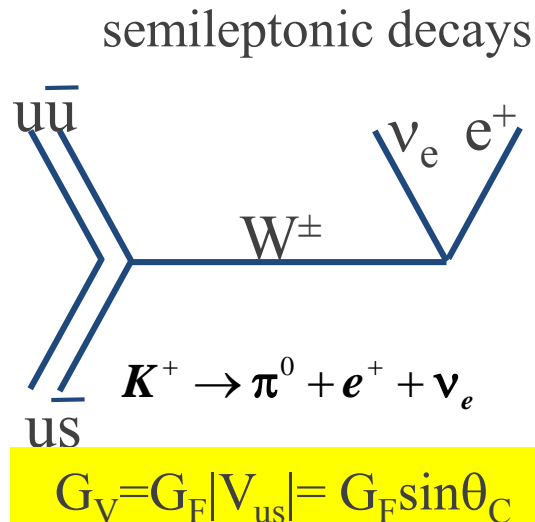
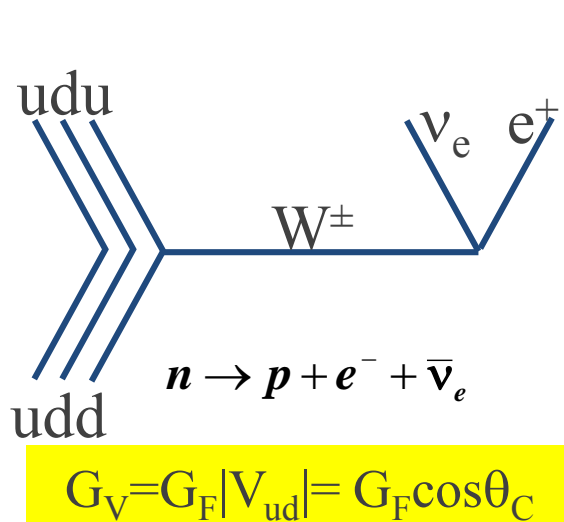


# CKM matrix and universality of weak interaction



$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

weak eigenstates      rotation matrix      mass eigenstates



# Superallowed Beta Decay

- » Precision tests of CVC
- » Determination of weak vector coupling constant and  $V_{ud}$
- » Unitary tests of the CKM matrix

- Basic  $\beta$ -decay rate equation for allowed decay

$$ft = \frac{K}{G_V^2 \left( |M_F|^2 + g_A^2 |M_{GT}|^2 \right)}$$

- Superallowed transitions between  $0^+$  T=1 states
  - $0^+ \rightarrow 0^+$  is a pure vector (Fermi) decay within SM
  - CVC tells us that all such decays should have same ft value

$$ft_{0^+ \rightarrow 0^+} = \frac{K}{G_V^2 |M_F|^2} = \text{const.}$$

with, within isospin symmetry,  $|M_F|^2 = T(T+1) = 2$  for T=1 analog states.

# Weak interaction rates

Rate by Fermi's golden rule:

$$d\lambda(E) = \left( \frac{2\pi}{\hbar} \right) |H_{fi}|^2 \rho(E)$$

density of final states strongly influences rate.

- beta decay (3 bodies in the final state):

$$\lambda(p_e) = \frac{g^2 |M_{if}|^2}{2\pi^3 \hbar^7 c^3} p_e^2 F(Z, E_e) (E_0 - E_e) \left[ (E_0 - E_e)^2 - (m_\nu c^2)^2 \right]^{1/2} dp_e$$

$$\lambda = \frac{\ln 2}{t_{1/2}} = \frac{m_e^5 g^2 c^4 |M_{if}|^2}{2\pi^3 \hbar^7} f(Z, E_0) \quad (\text{with } f \sim E^5)$$

- neutrino capture (2 bodies in final state):

$$\sigma_\nu = \frac{g^2 |M_{if}|^2 F(Z, E) p_e E_e}{\pi \hbar^4 c^3} \quad (\sigma \sim E^2)$$

# Superaligned Beta Decay

- » Precision tests of CVC
- » Determination of weak vector coupling constant and  $V_{ud}$
- » Unitary tests of the CKM matrix

- Basic  $\beta$ -decay rate equation for allowed decay

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with, within isospin symmetry,  $|M_F|^2 = T(T+1) = 2$  for T=1 analog states.

# Superaligned Beta Decay

- Demonstrating that all such  $ft$  values (accounting for small corrections) are constant tests CVC, puts stringent limits on scalar currents, and yields best value of  $G_V$ .

- Basic test for physics beyond the standard model

$G_V$  together with  $G_\mu$  yield  
the  $V_{ud}$  quark mixing element  
of the CKM matrix

$$\begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix}$$

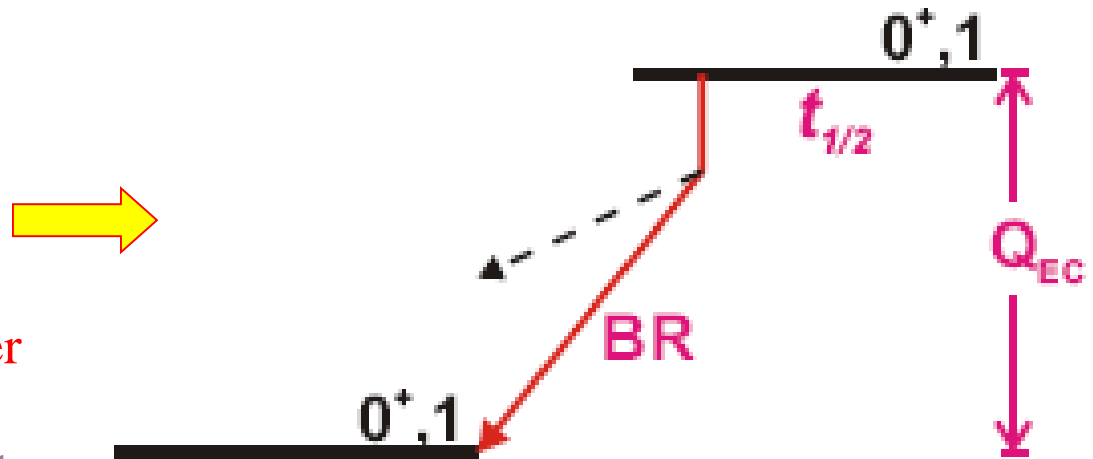
- If matrix is not unitary then we need new physics  
Additional Z bosons, Right-handed currents, ...

# Experimental inputs to Superallowed Fermi Decay

We require:

- Q-value
- lifetime
- branching ratio

Goal is for  $ft$  value at better than 0.1%. Last two measurements are required to that level, first one to 5 times better. All are at the limit of what present technology allows.



# Theoretical inputs to Superallowed Fermi Decay

WEAK DECAY EQUATION

$$ft = \frac{K}{G_V^2 \langle \tau \rangle^2}$$

$$f = f(Z, Q_{EC})$$

$$t = f(t_{1/2}, BR)$$

$G_V$  = coupling constant

$\langle \tau \rangle$  = matrix element

RADIATIVE CORRECTIONS

$$t \rightarrow t(1 + \delta_R' + \delta_{NS})$$

$$G_V^2 \rightarrow G_V^2(1 + \Delta_R)$$

$$\delta_R' = f(Z, Q_{EC})$$

$$\delta_{NS} = f(\text{nuclear structure})$$

$$\Delta_R = f(\text{interaction})$$

T SYMMETRY-BREAKING CORRECTIONS

$$\langle \tau \rangle^2 \rightarrow 2(1 - \delta_C)$$

$$\delta_C = f(\text{nuclear structure})$$

CVC

$$\mathcal{F}t = ft(1 + \delta_R' + \delta_{NS})(1 - \delta_C) = \frac{K}{2G_V^2(1 + \Delta_R)}$$

CONSTANT



# Superaligned Fermi decay candidates

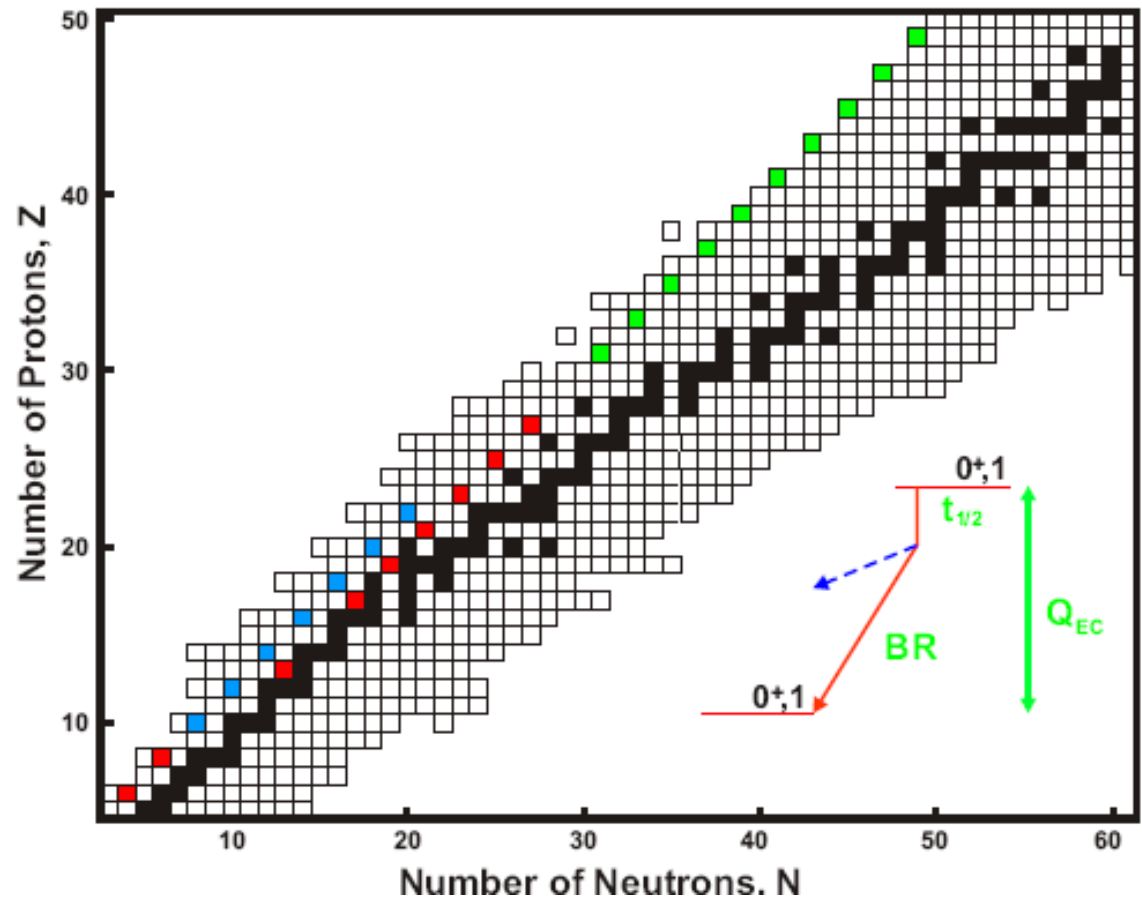
## 8 “golden” cases:

$^{14}\text{O}$  to  $^{54}\text{Co}$

- superaligned Fermi branch  $> 99\%$

- daughter is a stable nucleus

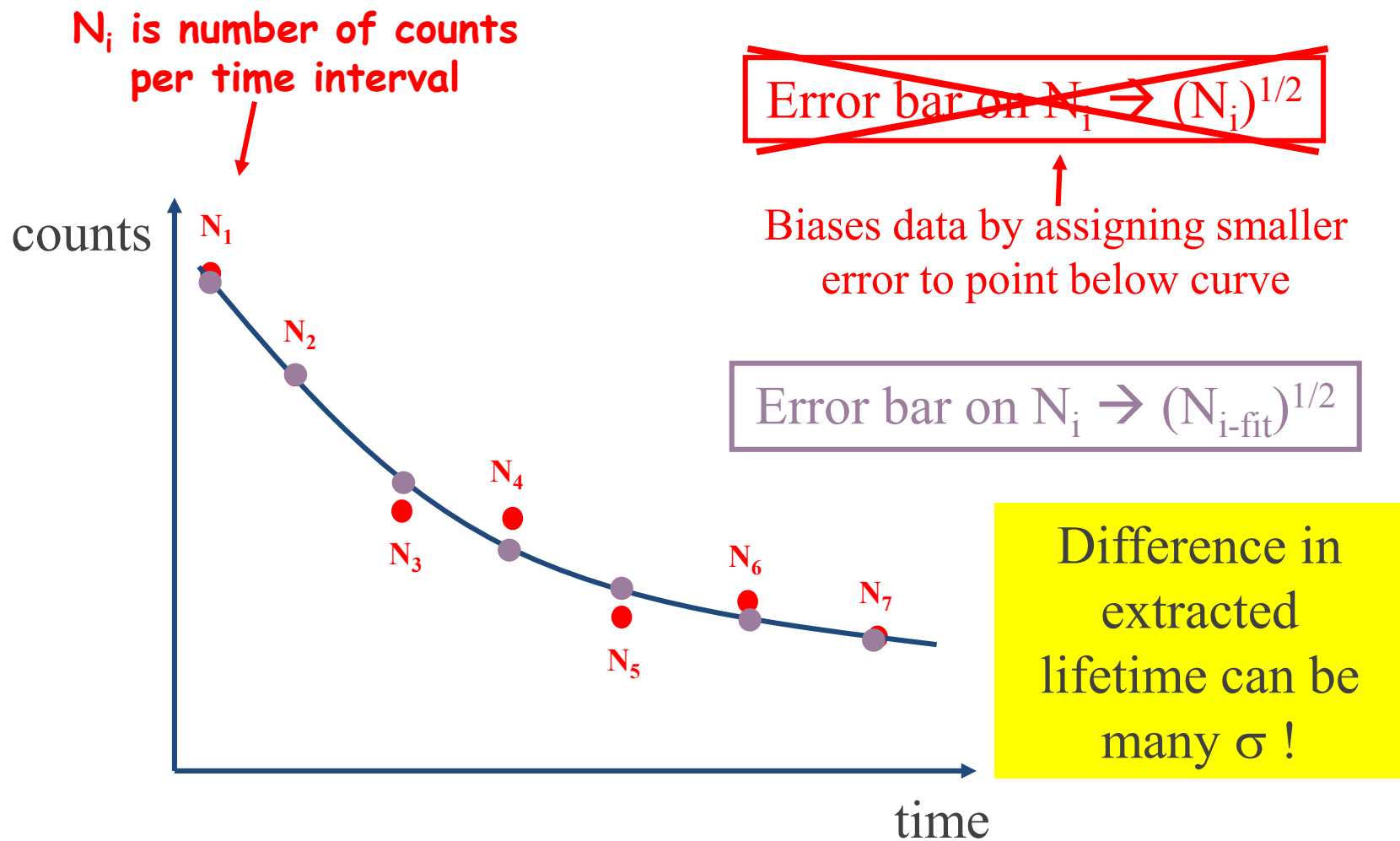
- And a whole bunch of good but difficult cases



# Precision measurements --- lifetime

- In principle **simplest measurement there is** ... yet this is where the **biggest mistakes** have been made
- At the required accuracy of about 0.05%, no measurement done before 1970 (maximum likelihood introduced) is correct. Analysis procedure needs to be tested with pseudo-data and Poisson statistics.
- This is not high energy: signature is one low energy  $\beta$  ... backgrounds are present. Need **to collect data for about 14 to 20 lifetimes** to fix background ... analysis cannot do it reliably otherwise.
- In practice, use mass separated samples if available, almost 100% efficient detector (to eliminate pile-up), and fixed deadtime in the electronics (we often used two parallel acquisitions with different deadtime to check systematics).

# Precision lifetime measurements can easily be corrupted in the experiment or the analysis



# Precision measurements --- lifetime

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# Precision measurements --- branching ratio

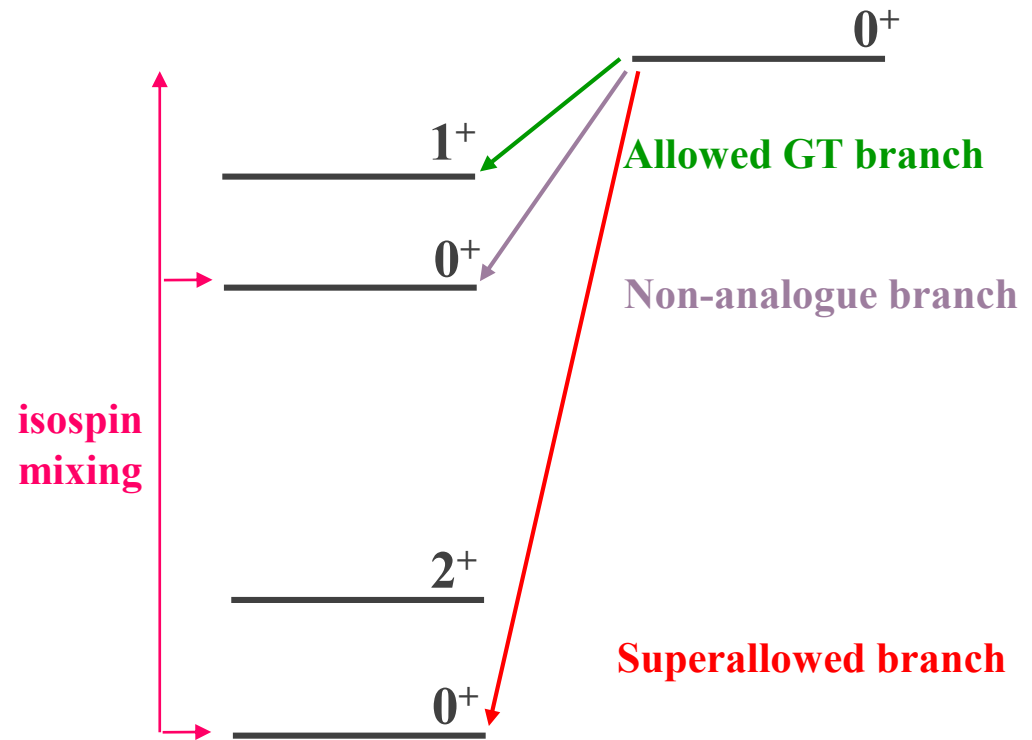
- parent and daughter are not perfect analogues

- measured rate lower than expected

- isospin mixing one contributing factor

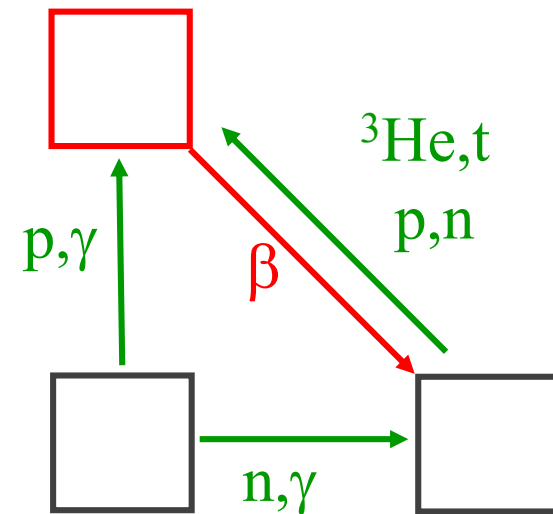
- necessary experimental quantity

- aid theory (calculated correction terms)



# Precision measurements --- Q-value

- required precision (0.1 – 1 keV) cannot be reached by endpoint measurement
- Use reaction threshold (neutron yield vs energy for p,n reaction) or gamma-ray energy
- Difficulty then lies in proton energy calibration
- Can only be applied to nuclei whose daughter is a stable target



# For Isotopes whose daughter is not stable: Precision measurements in a Penning trap

Can use:

$$\omega_c = \omega_+ + \omega_-$$
$$\omega_c^2 = \omega_+^2 + \omega_-^2 + \omega_z^2$$

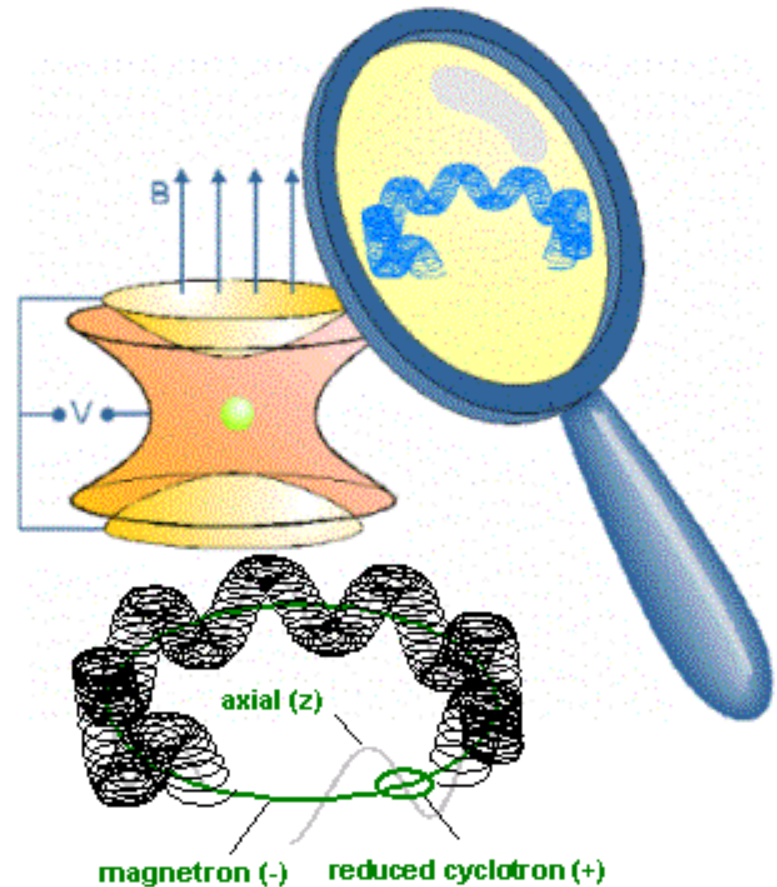
Recall:

$$\omega_c = \frac{qB}{\gamma m}$$

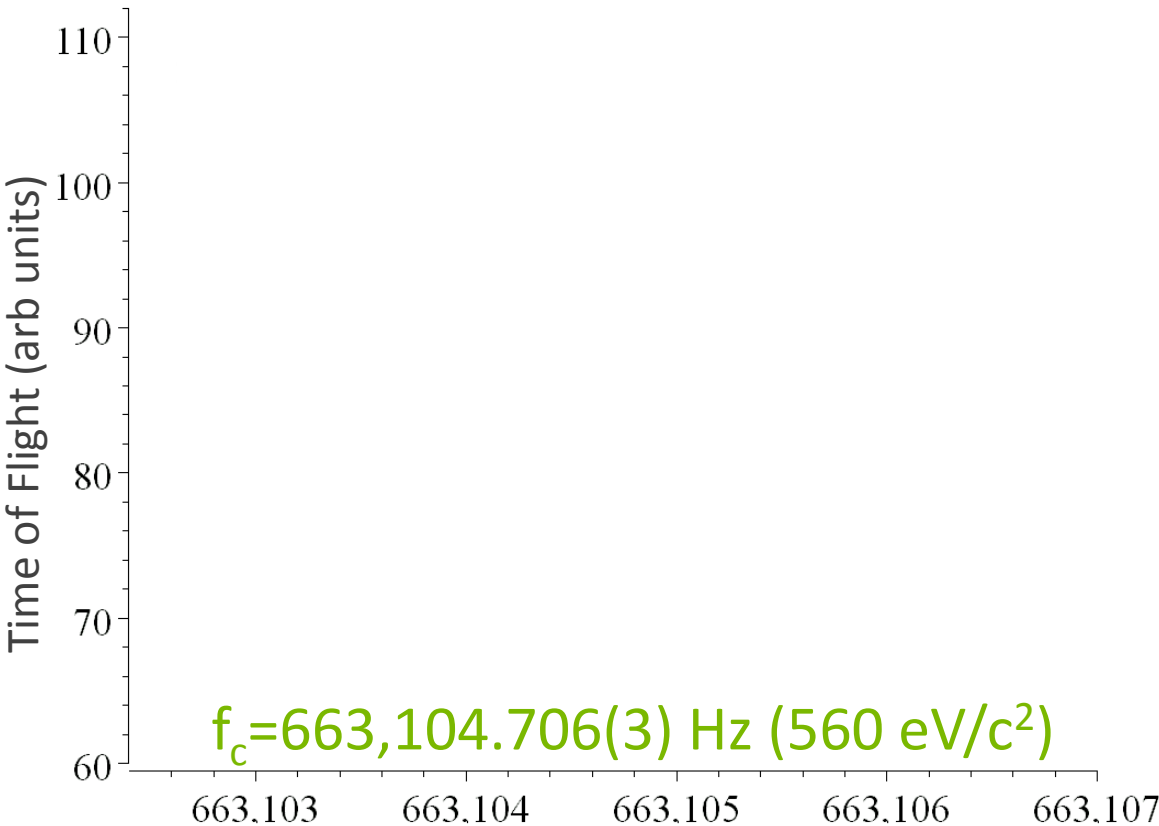
$\omega_c$  depends only on:

- the mass
- the magnetic field
- not on the electric fields or the energy as long as  $\gamma$  is small

Can use  $\omega_c$  to make accurate and precise mass measurements



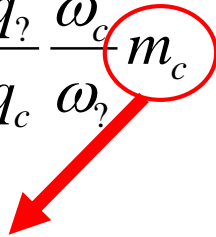
# Sample time-of-flight (TOF) spectrum



$$\omega_c = \frac{q_c B_c}{m_c}$$

Unknown:  $\omega_\gamma = \frac{q_\gamma B_c}{m_\gamma}$

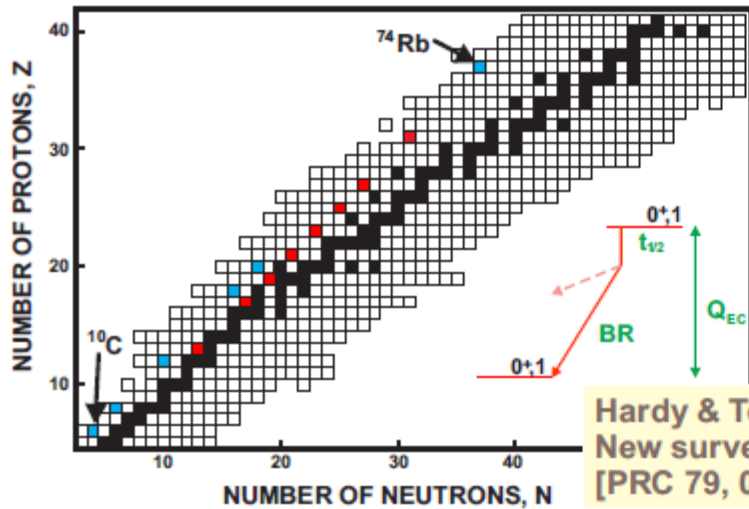
$$\frac{\text{Unknown}}{\text{Calibration}} \Rightarrow m_\gamma = \frac{q_\gamma}{q_c} \frac{\omega_c}{\omega_\gamma} m_c$$



Well-known calibrant mass is a requirement for accurate measurements, use  $^{133}\text{Cs}$  (known to  $\sim 0.01 \text{ keV}$ ) in this region.



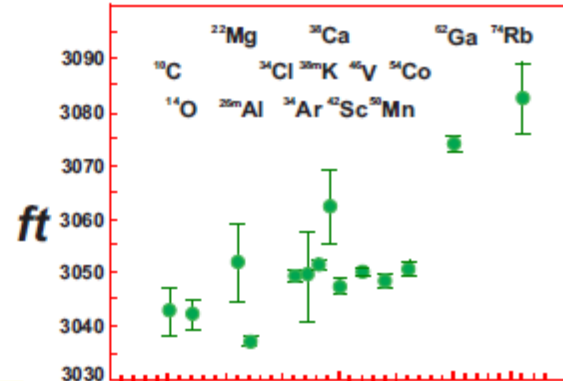
# 0+ → 0+ status as of 2014



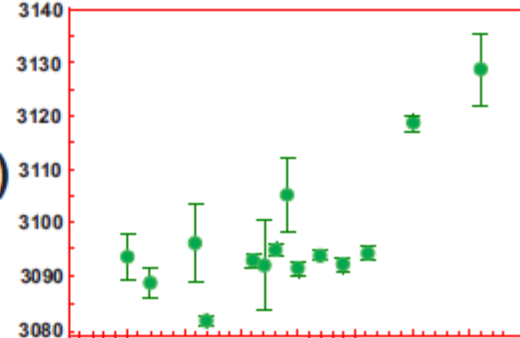
Hardy & Towner  
New survey (2014)  
[PRC 79, 055502 (2009)]

- 8 cases with  $ft$ -values measured to **<0.05% precision**; 6 more cases with **0.05-0.3% precision**.
- ~220 individual measurements with compatible precision

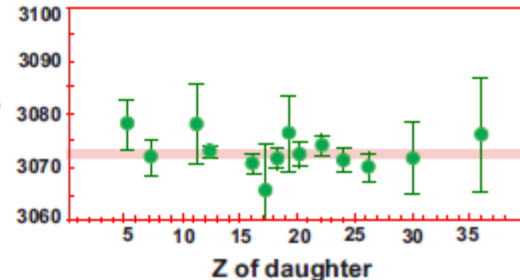
$$\mathcal{F}t = ft(1 + \delta'_R)[1 - (\delta_C - \delta_{NS})] = \frac{K}{2G_V^2(1 + \Delta_R)}$$



$ft(1 + \delta'_R)$



$\mathcal{F}t$



# Results from $0^+ \rightarrow 0^+$ decays

## FROM A SINGLE TRANSITION

Experimentally  
determine  $G_V^2(1 + \Delta_R)$

$$\tau t = ft (1 + \delta'_R) [1 - (\delta_C - \delta_{NS})] = \frac{K}{2G_V^2 (1 + \Delta_R)}$$

## FROM MANY TRANSITIONS

Test Conservation of  
the Vector current (CVC)

Validate correction terms ✓

Test for Scalar current

$G_V$  constant to  $\pm 0.013\%$

limit,  $C_S/C_V = 0.0014 (13)$

## WITH CVC VERIFIED

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

weak  
eigenstates

mass  
eigenstates

Cabibbo-Kobayashi-Maskawa matrix

Obtain precise value of  $G_V^2(1 + \Delta_R)$

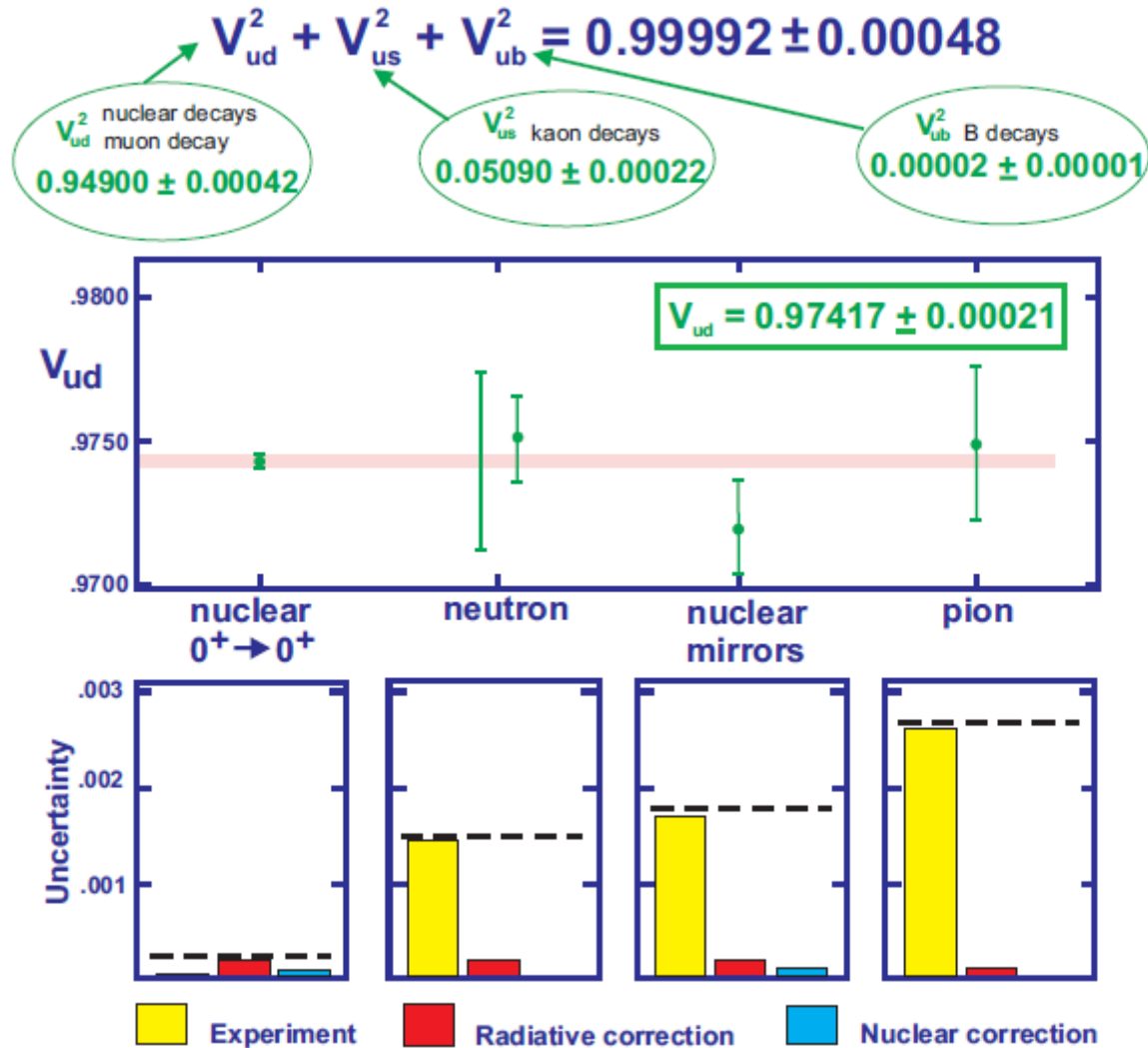
Determine  $V_{ud}^2$

$$V_{ud}^2 = G_V^2/G_\mu^2 = 0.94900 \pm 0.00042$$

Test CKM unitarity

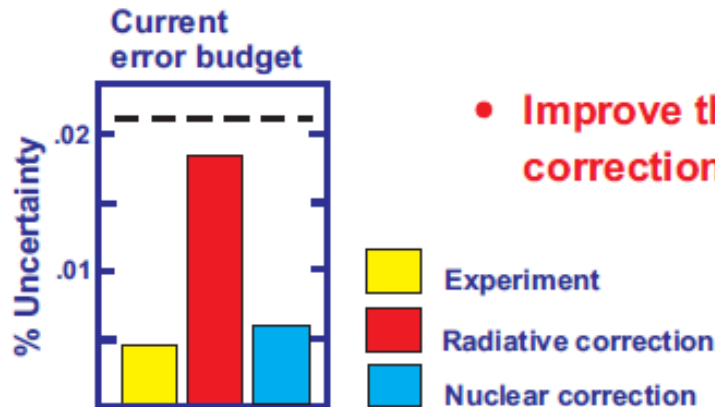
$$V_{ud}^2 + V_{us}^2 + V_{ub}^2 = 0.99992 \pm 0.00048$$

# Status of CKM unitarity



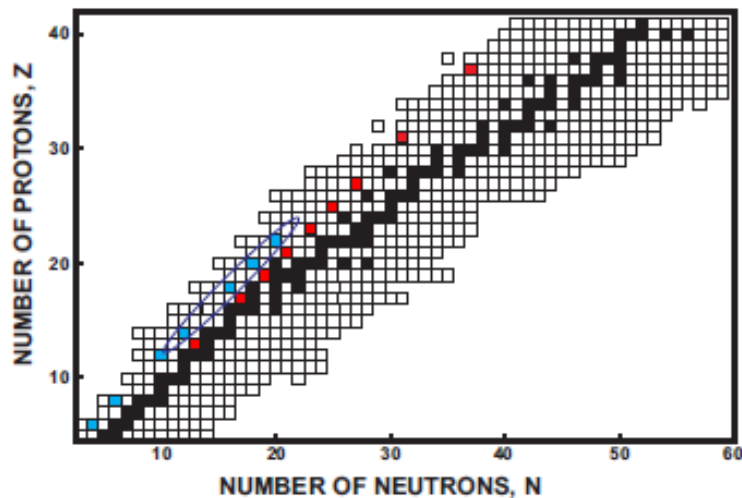
Courtesy of J.C. Hardy

# Next step: Improving $V_{ud}$ and CKM matrix unitarity



- Improve the calculation of the “inner” radiative correction  $\Delta_R$ .

Currently  $\Delta_R = 2.361(38) \%$   
Marciano & Sirlin, PRL 96, 032002 (2006)



- Potentially improve nuclear correction by completing the pairs of mirror superallowed decays with  $A \leq 42$ .

$^{38}\text{Ca}$  done  
H.I. Park *et al.*, PRL 112, 102502 (2014)  
 $^{26}\text{Si}$ ,  $^{34}\text{Ar}$  and  $^{42}\text{Ti}$  remain

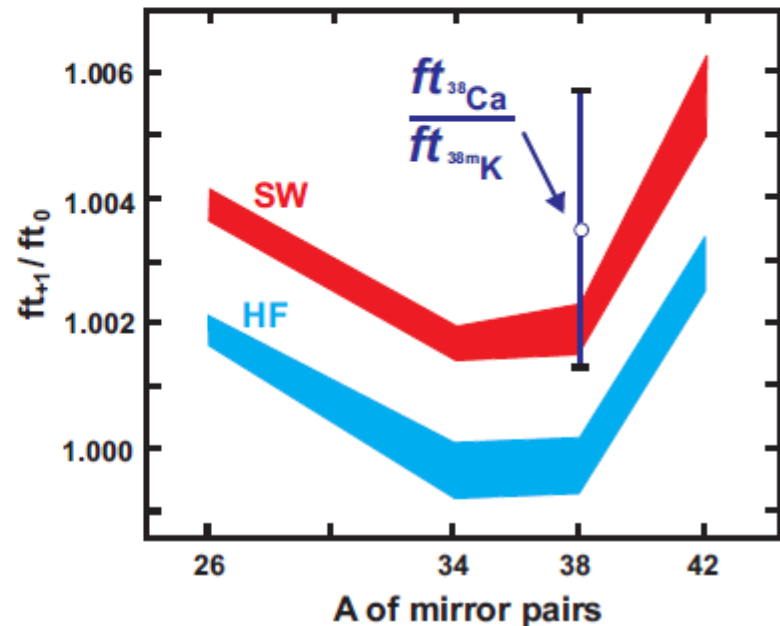
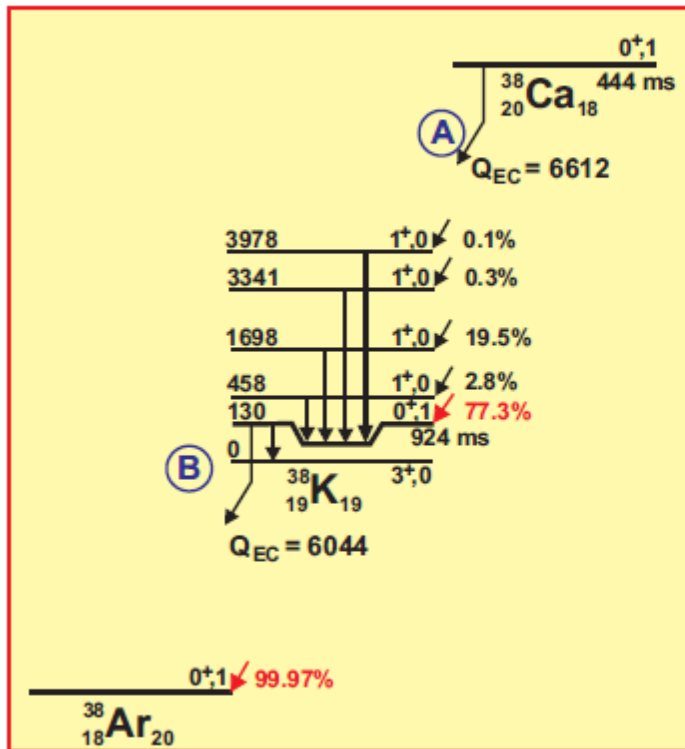
- If theory improves, broadly improve all experimental  $ft$  values

# Testing $\delta_c$ calculations with measurements of mirror superallowed transitions

$$f t = f t (1 + \delta'_R) [1 - (\delta_C - \delta_{NS})]$$

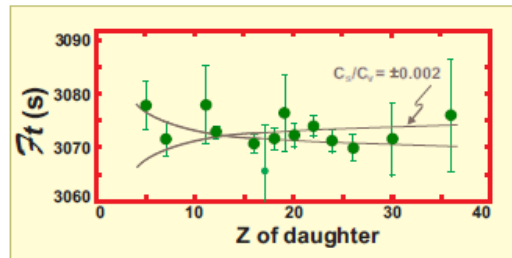
$$\frac{f t_A}{f t_B} = \frac{(1 + \delta'_R)^B [1 - (\delta_C^B - \delta_{NS}^B)]}{(1 + \delta'_R)^A [1 - (\delta_C^A - \delta_{NS}^A)]}$$

$$= 1 + (\delta'_R)^B - \delta'_R^A + (\delta_{NS}^B - \delta_{NS}^A) - (\delta_C^B - \delta_C^A)$$



# Further progress coming at low-energy

## Search for scalar currents

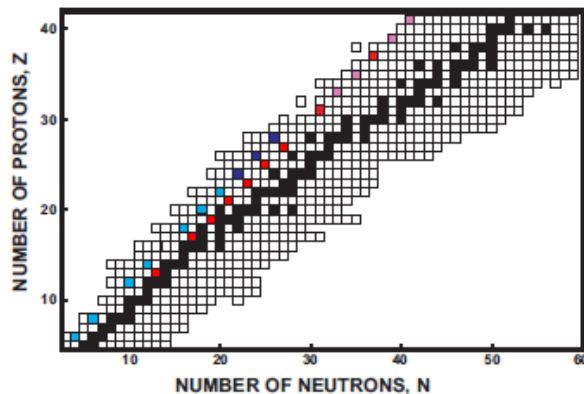


limit,  $C_S/C_V = 0.0014$  (13)

- Tighten uncertainties on the  $ft$  values for  $^{10}\text{C}$  and  $^{14}\text{O}$

Require (order of priority):  
 $^{10}\text{C}$  branching ratio  
 $^{14}\text{O}$  branching ratio  
 $^{14}\text{O}$   $Q_{\text{EC}}$  value  
 $^{10}\text{C}$  half-life

## Adding new transitions

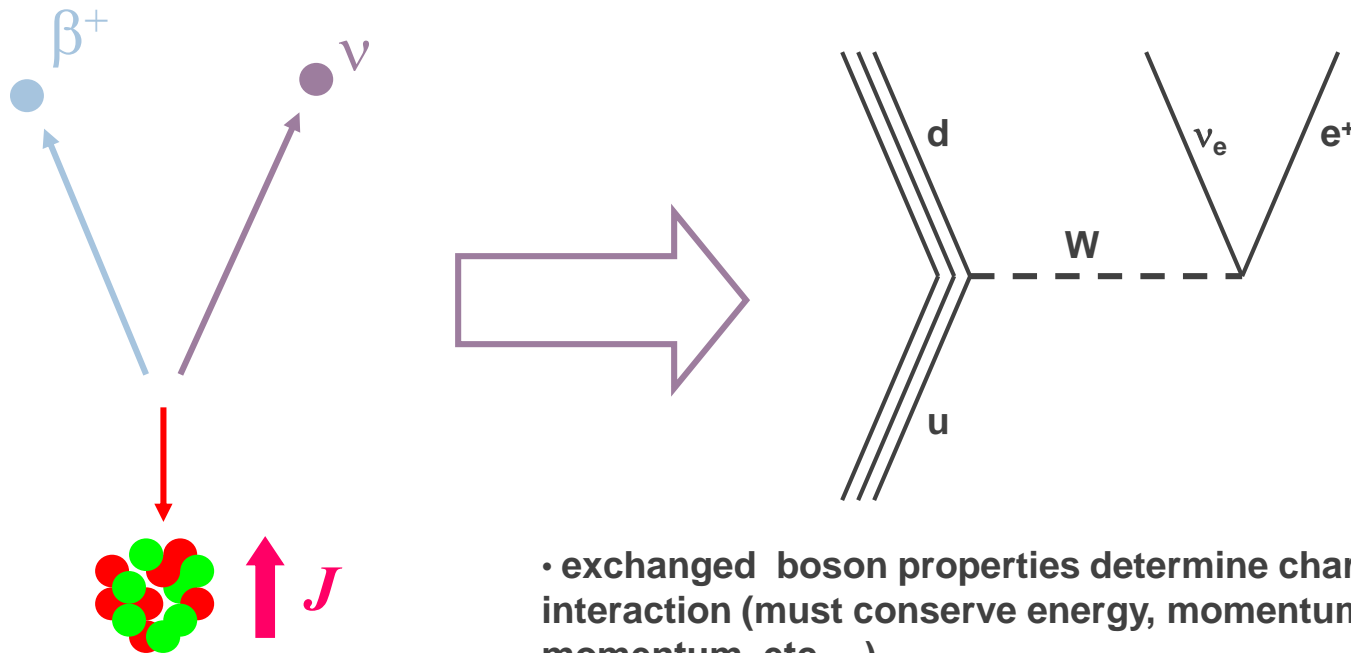


- Potential FRIB measurements
- Complete more pairs of mirror superallowed transitions:  $^{46}\text{Cr}$ ,  $^{50}\text{Fe}$ ,  $^{54}\text{Ni}$  ...

Tests  $\delta_C$ ,  $\delta_{\text{NS}}$  calculations in  $f_{7/2}$  shell nuclei

- Add new heavy  $T_z = 0$  superallowed emitters:  $^{66}\text{As}$ ,  $^{70}\text{Br}$ ,  $^{78}\text{Y}$  ...

# Correlations in nuclear $\beta$ -decay



- exchanged boson properties determine character of interaction (must conserve energy, momentum, angular momentum, etc ...)

e.g. **vector boson** in an allowed Fermi decay (no angular momentum for emitted particles) implies that the lepton spins must anti-align (since no spin change) ... i.e. **neutrino and positron (opposite chirality)** emitted preferentially in the same direction in Fermi decay

- interaction best studied in decay selected to limit other contributions

# Basic rules for getting form of $\Gamma$

Take neutron beta decay as an example  $n \rightarrow p + e^- + \bar{\nu}_e$

Available vectors:  $\vec{\sigma}_n, \vec{p}_n, \vec{\sigma}_p, \vec{p}_p, \vec{\sigma}_e, \vec{p}_e, \vec{\sigma}_\nu, \vec{p}_\nu$

$\Gamma$  is invariant under rotation so scalar or pseudoscalar and can therefore contain terms like

$\vec{\sigma}_n \bullet \vec{\sigma}_p$	$\vec{\sigma}_n \bullet \vec{p}_e$	← P-violating
$\vec{p}_p \bullet \vec{p}_e$	$\vec{\sigma}_n \bullet (\vec{p}_e \times \vec{p}_\nu)$	← T-violating
$\vec{p}_\nu \bullet \vec{p}_e$	....	

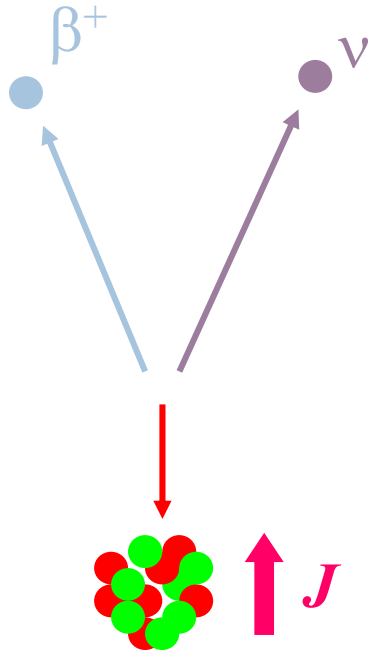
Using as independent vectors  $\vec{\sigma}_n, \vec{p}_e, \vec{p}_\nu$  :

$$\Gamma \propto 1 + \alpha(\vec{p}_e \bullet \vec{p}_\nu) + \vec{\sigma}_n \bullet [A \vec{p}_e + B \vec{p}_\nu + D \vec{p}_e \times \vec{p}_\nu]$$





# Angular correlations in $\beta$ -decay



Angular correlations between momentum and spin vectors of the particle emitted in beta-decay yield information about the nature of the interaction.

- Compare experimental values to SM predictions
- Put limits on terms “forbidden” by SM

$$dW = dW_o \varepsilon \left[ 1 + \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} \mathbf{a} + \frac{\Gamma m_e}{E_e} \mathbf{b} + \vec{J} \cdot \left( \frac{\vec{p}_e}{E_e} \mathbf{A} + \frac{\vec{p}_\nu}{E_\nu} \mathbf{B} + \frac{\vec{p}_e \times \vec{p}_\nu}{E_e E_\nu} \mathbf{D} \right) + \dots \right]$$

# Boson mass range that can be probed

If you observe a nuclear beta-decay, even in a table top experiment, you are already at 80.4 GeV/c<sup>2</sup>.

How high can you go?

$$\text{Coupling} \sim (M^2 - q^2)^{-1} \longrightarrow M^{-2} \text{ (as } q \rightarrow 0 \text{)}$$

$$\text{Observable} \sim (\text{Coupling})^2 \longrightarrow M^{-4} \text{ (as } q \rightarrow 0 \text{)}$$

$$1 \% \text{ expt} \longrightarrow (0.01)^{-4} M_W \sim 260 \text{ GeV}/c^2$$

$$0.1\% \text{ expt} \longrightarrow (0.001)^{-4} M_W \sim 460 \text{ GeV}/c^2$$

Sensitivity is also different than for other experiments, eg: non-manifest left-right model (assuming  $\zeta=0$  and  $\nu_R$  much lighter than  $M_R$ ), we get the following sensitivities:

nuclear  $\beta$ -decay

$$\left(\frac{g_R}{g_L}\right)^4 \left(\frac{V_{ud_R}}{V_{ud_L}}\right)^2 \left(\frac{M_L}{M_R}\right)^4$$

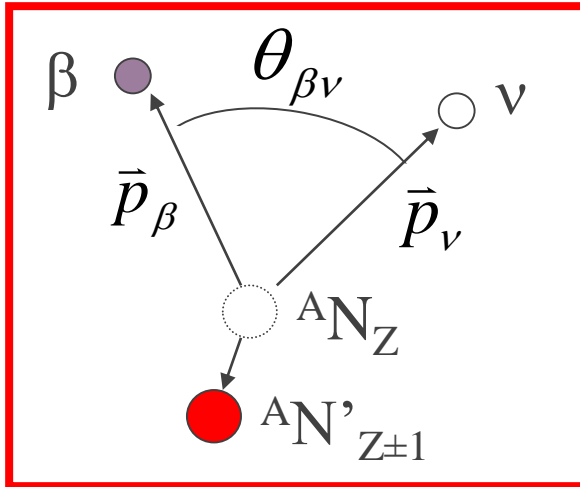
$\mu$  decay

$$\left(\frac{g_R}{g_L}\right)^4 \left(1 - \left(\frac{V_{ud_R}}{V_{ud_L}}\right)^2\right) \left(\frac{M_L}{M_R}\right)^4$$

pp collider

$$\left(\frac{g_R}{g_L}\right)^2 \left(\frac{V_{ud_R}}{V_{ud_L}}\right)^2 \text{function} \left(\frac{M_L}{M_R}\right)$$

# $\beta$ - $\nu$ Correlation



$$\Gamma \propto 1 + a_{\beta\nu} \frac{\vec{p}_\beta \cdot \vec{p}_\nu}{E_\beta E_\nu}$$

$$\propto 1 + a_{\beta\nu} \frac{v}{c} \cos \theta_{\beta\nu}$$

Experiments difficult – correlation inferred from recoil of nucleus

$a > 0$  leads to larger average recoil energies

Recoil energy  $\sim 100$  eV

- 1) requires acceleration of daughters
- 2) infer recoil from energy shift in subsequent particle decay

Continuous energy spectra in all particles

sensitive to detector thresholds and resolution

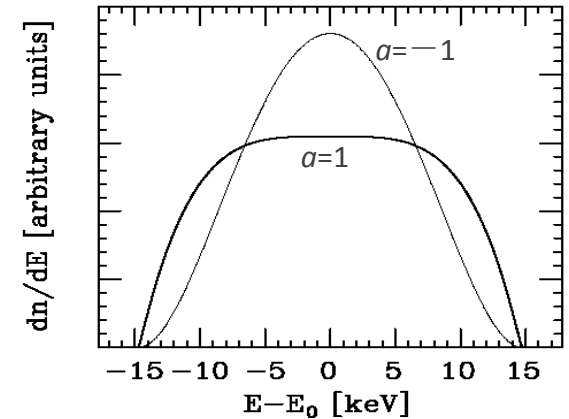
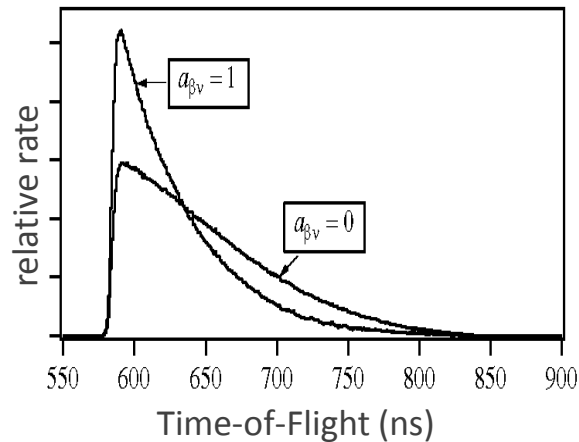
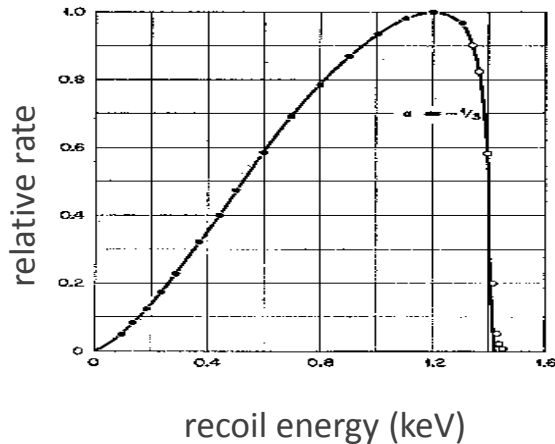
sensitive to approximations made about underlying physics

Correlation easily perturbed by molecular effects, scattering, etc...

# How to measure?

Avoid detecting neutrino, measure other observables related to  $\theta_{(\beta\nu)}$ .

Measure decay product energy directly	Measure TOF of recoil nucleus	Measure delayed particle emission
n, $^{19}\text{Ne}$ , $^{23}\text{Ne}$ , $^{35}\text{Ar}$ , $^{60}\text{Co}(\beta)$ $^6\text{He}$	n, $^6\text{He}$ , $^{38}\text{Mg}$ $^{21}\text{Na}$ , $^{37}\text{K}$ , $^{19}\text{Ne}$	$^8\text{Li}$ , $^{11}\text{Be}$ , $^{14}\text{O}$ $^{18}\text{Ne}$ , $^{20}\text{Na}$ , $^{32}\text{Ar}(p)$



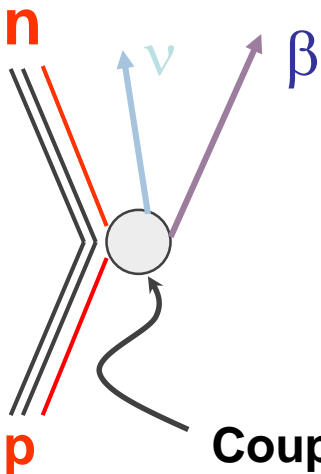
$0^+ \rightarrow 0^+$  delayed proton for  $a=1$  &  $a=-1$

C.H. Johnson *et al.* Phys. Rev. 132, 1149

N. D. Scielzo *et al.* PRL.93,102501, 2004

E.G. Adelberger *et al.* PRL.83, 1299 (1999).

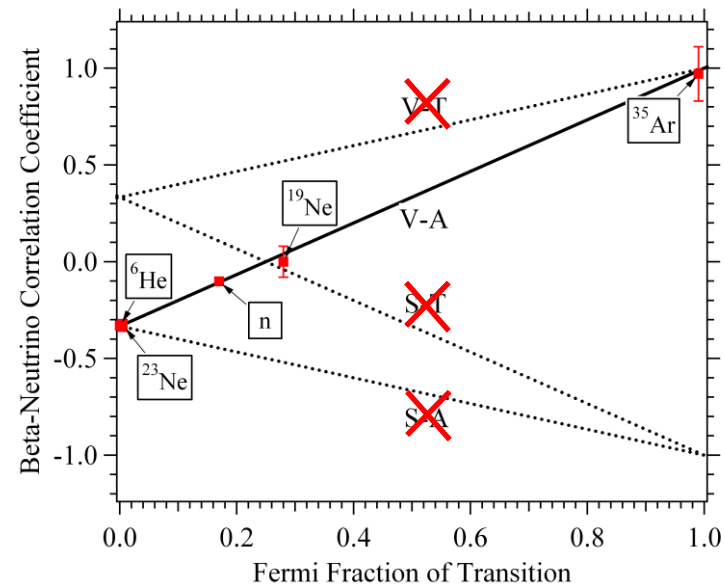
# Measurements of $\beta$ -decay angular correlations helped establish the V-A structure of the electroweak interaction



Coupling constants:  
 $C_V, C_A, \cancel{C_S}, \cancel{C_T}$

$$dW = dW_0 \left[ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{\Gamma m_e}{E_e} \right]$$

In the 1960's, the V-A structure of the weak interaction was determined by measurements of the beta-neutrino correlation,  $a$ , in noble gas nuclei



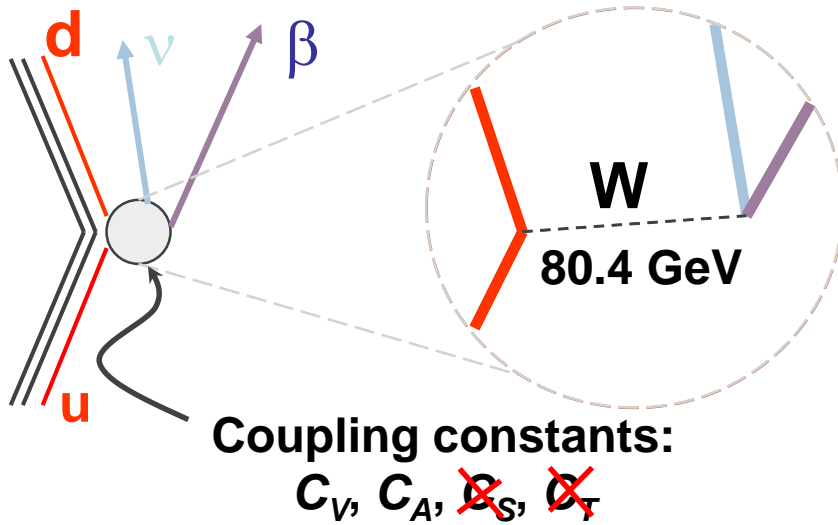
Fermi decay

$\rightarrow C_V, C_S$

Gamow-Teller decay

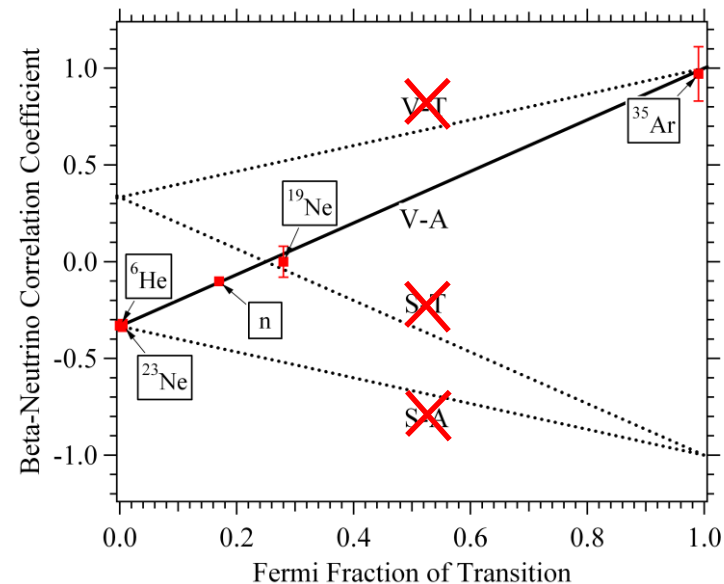
$\rightarrow C_A, C_T$

# Measurements of $\beta$ -decay angular correlations helped establish the V-A structure of the electroweak interaction



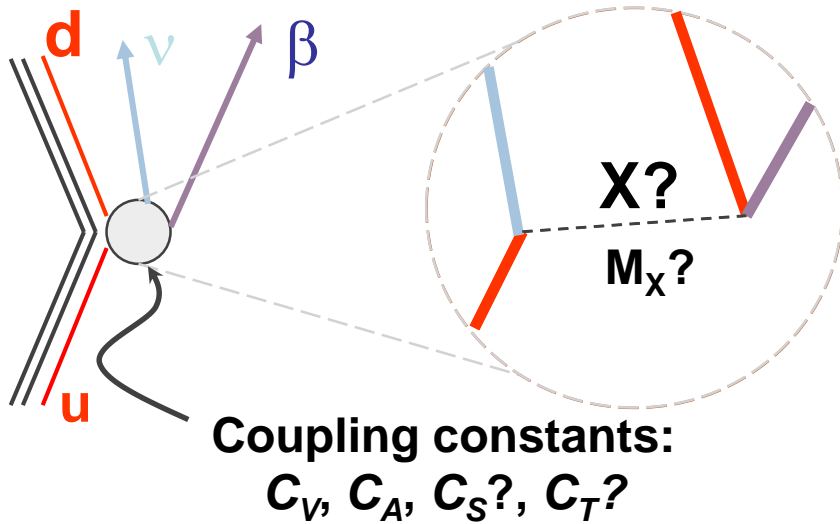
$$dW = dW_0 \left[ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{\Gamma m_e}{E_e} \right]$$

We now know  $\beta$  decay is mediated by the W boson with well known properties



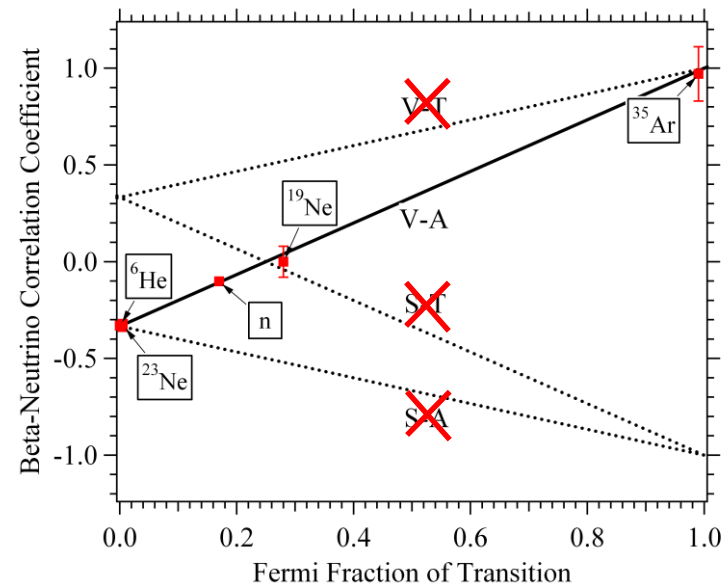
Fermi decay  $\rightarrow C_V$   
 Gamow-Teller decay  $\rightarrow C_A$

# Measurements of $\beta$ -decay angular correlations helped establish the V-A structure of the electroweak interaction



$$dW = dW_0 \left[ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{\Gamma m_e}{E_e} \right]$$

But perhaps there is more to discover if we look closely enough...



Fermi decay

$\rightarrow C_V, C_S?$

Gamow-Teller decay

$\rightarrow C_A, C_T?$

At higher precision, have to contend with recoil-order terms, which are sensitive to additional SM symmetries

Give rise to small  $E_e/M$  dependence to correlations

$$dW \sim \left[ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m_e}{E_e} \right]$$

Plus additional correlations:

$$g_{10}(E_e) = -\frac{E_e}{M} (1 + d \mp b_{wm}) \cdot \cos^2 \theta_{e\alpha}$$

$$\left( 1 - \frac{2E_0}{3M} (1 + d \pm b_{wm}) + \frac{2E_e}{3M} (5 \pm 2b_{wm}) \right)$$

- $b_{wm}$  (changes sign with  $\beta^\pm$ )
  - CVC hypothesis – weak vector coupling constant is not renormalized in the nucleus
- $d$  (independent of  $\beta^\pm$ )
  - Second-class currents – induced terms that do not obey same symmetries as strong interaction

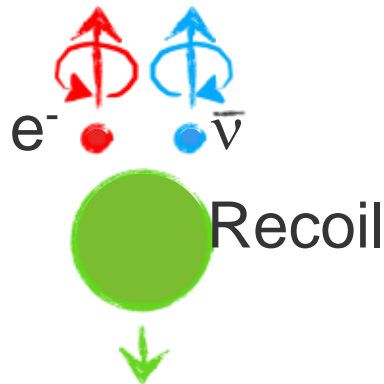
Need to study mirror nuclei to disentangle, e.g. both  ${}^8\text{Li}$  and  ${}^8\text{B}$



# Intuitive Picture

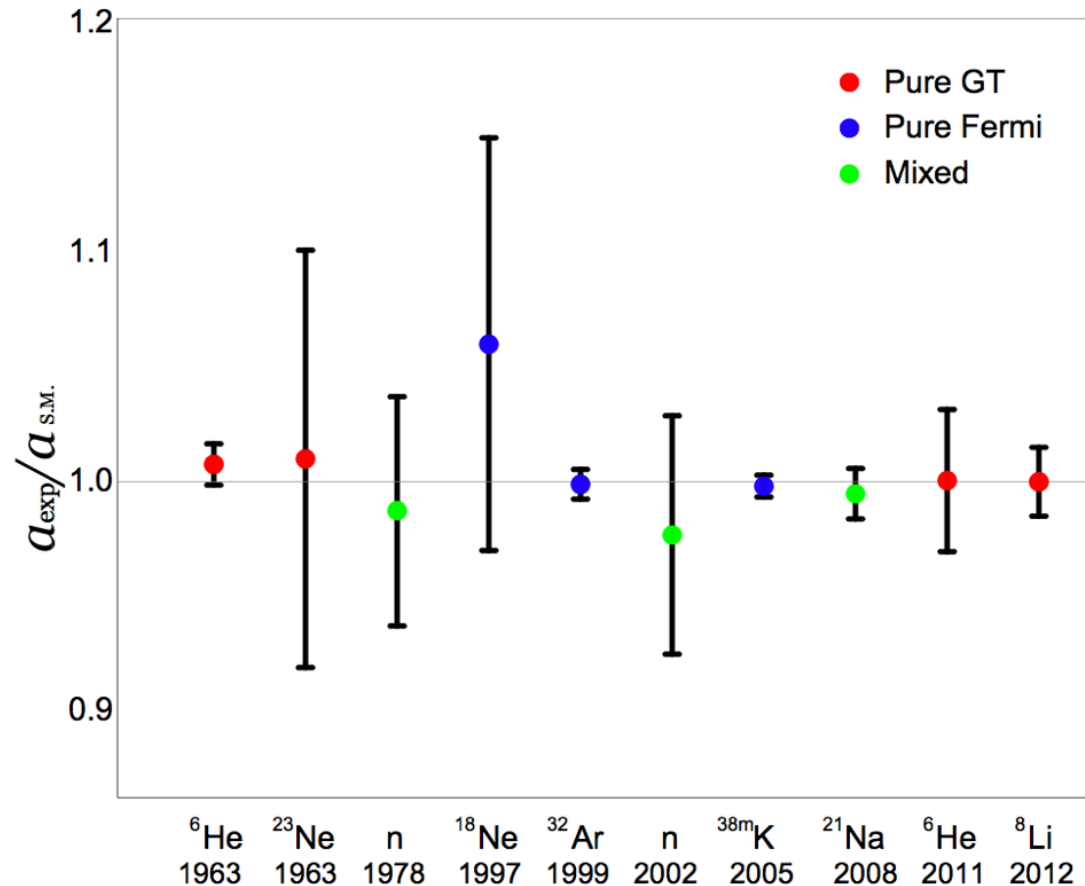
## Fermi Decay

- $\Delta J=0$  -> spins anti-align in singlet state
- $e^-$  &  $\bar{\nu}$  couple with opposite helicity
- $e^-$  &  $\bar{\nu}$  are preferentially emitted in same direction  $\rightarrow a=+1$

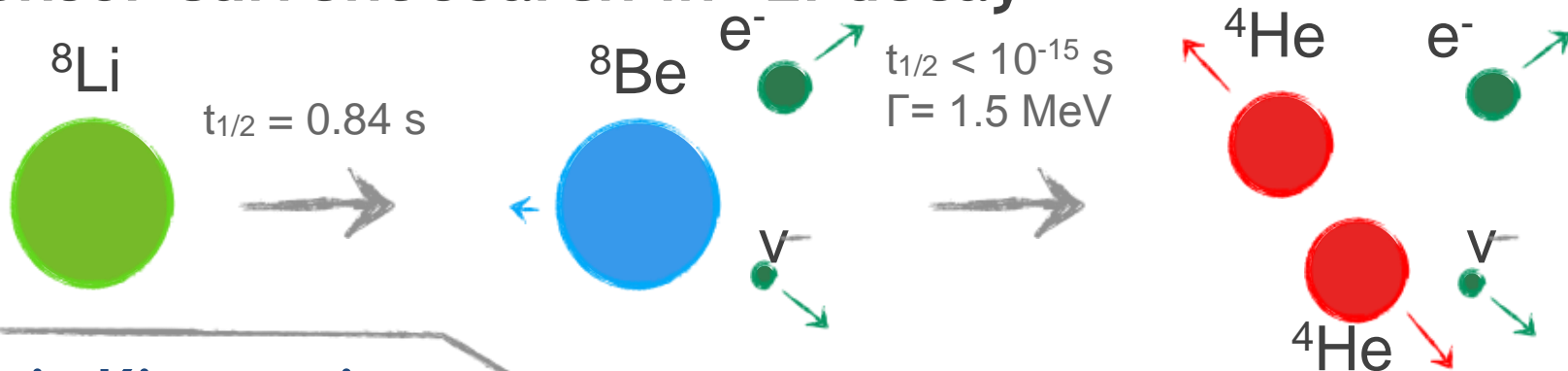


## GT Decay

- $\Delta J=0, \pm 1$  -> spins are in triplet
- same reasoning  $\rightarrow a=-1/3$

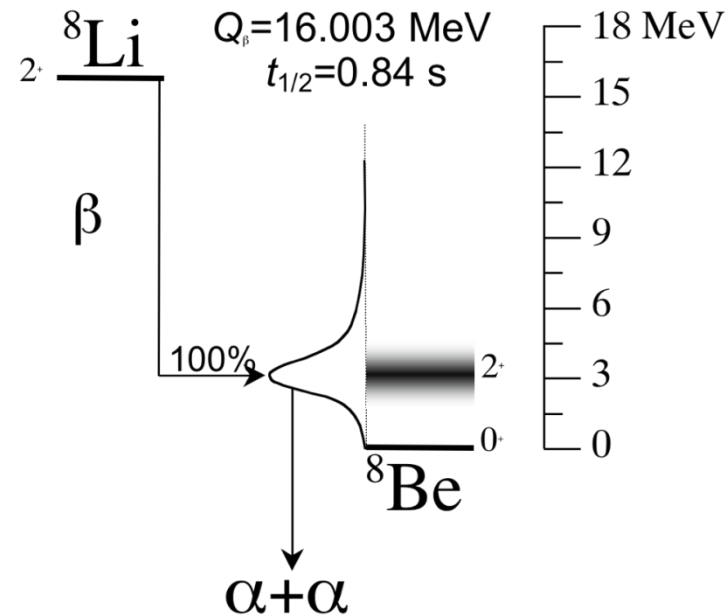


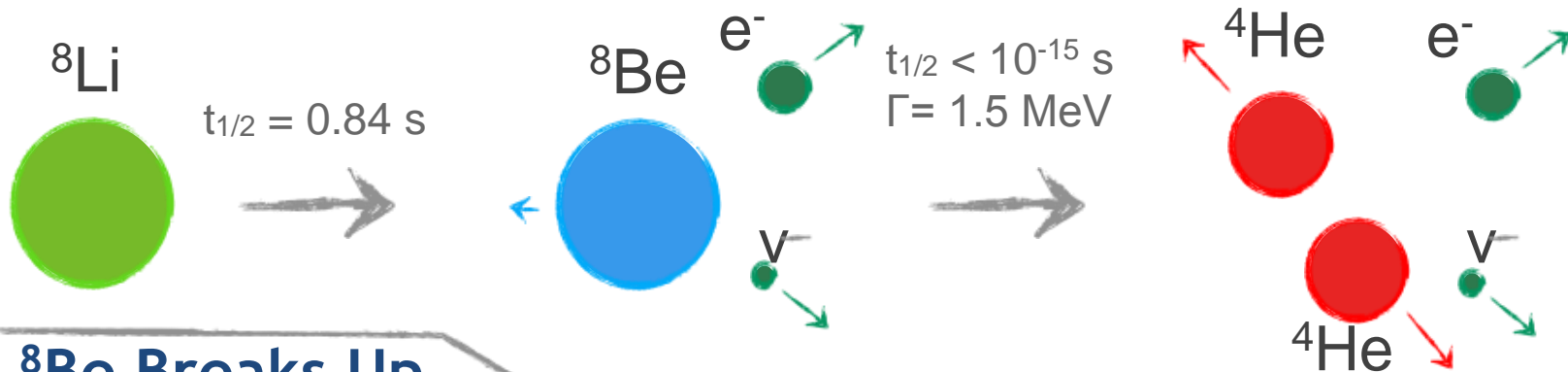
# Tensor current search in $^8\text{Li}$ decay



## Basic Kinematics

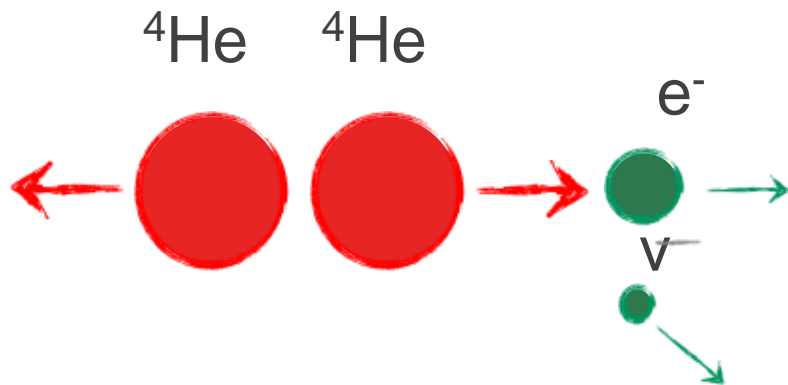
- “Pure” GT decay
- Large Q value and light nucleus gives large recoil energy up to 12 keV
- Immediate breakup to two  $\alpha$ 's gives coincident constraint providing strong background suppression
- $\alpha$ 's are emitted back to back in  $^8\text{Be}$  rest frame leading to large shifts in lab frame

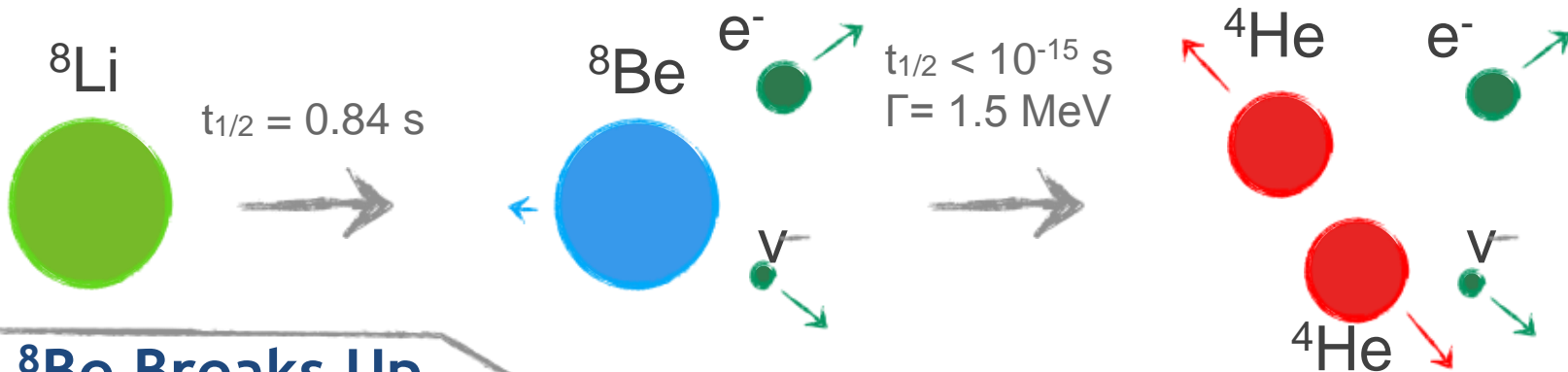




**${}^8\text{Be}$  Breaks Up  
Parallel to  $e^-$**

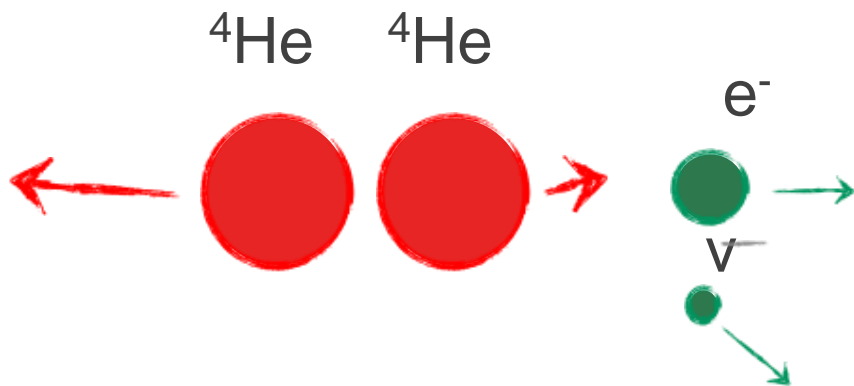
Reference Frame of  
Recoiling  ${}^8\text{Be}$



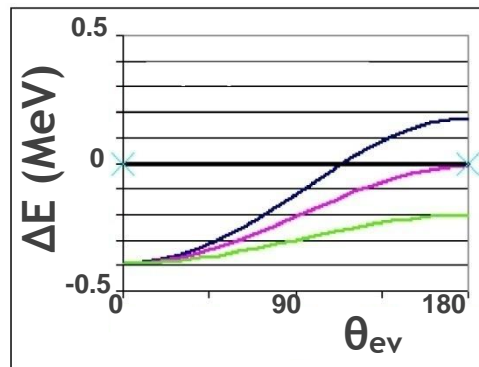


## ${}^8\text{Be}$ Breaks Up Parallel to $e^-$

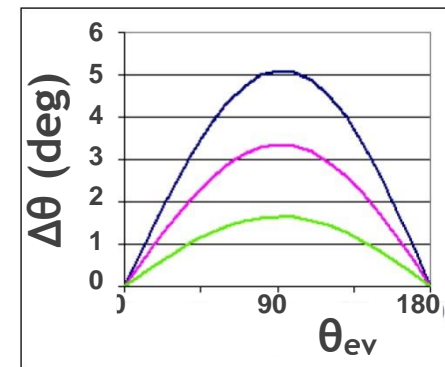
### Reference Frame of Laboratory



Energy Difference of  $\alpha$ 's

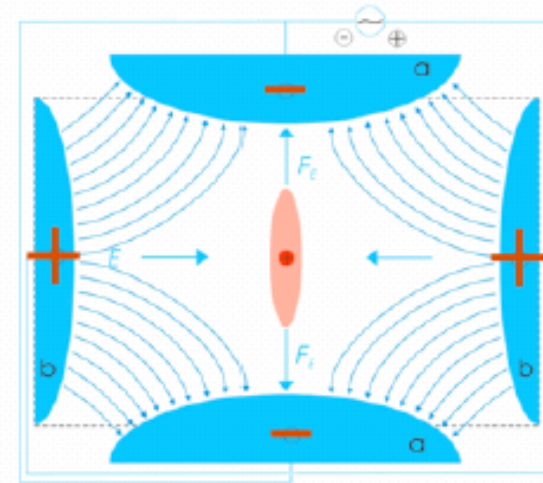
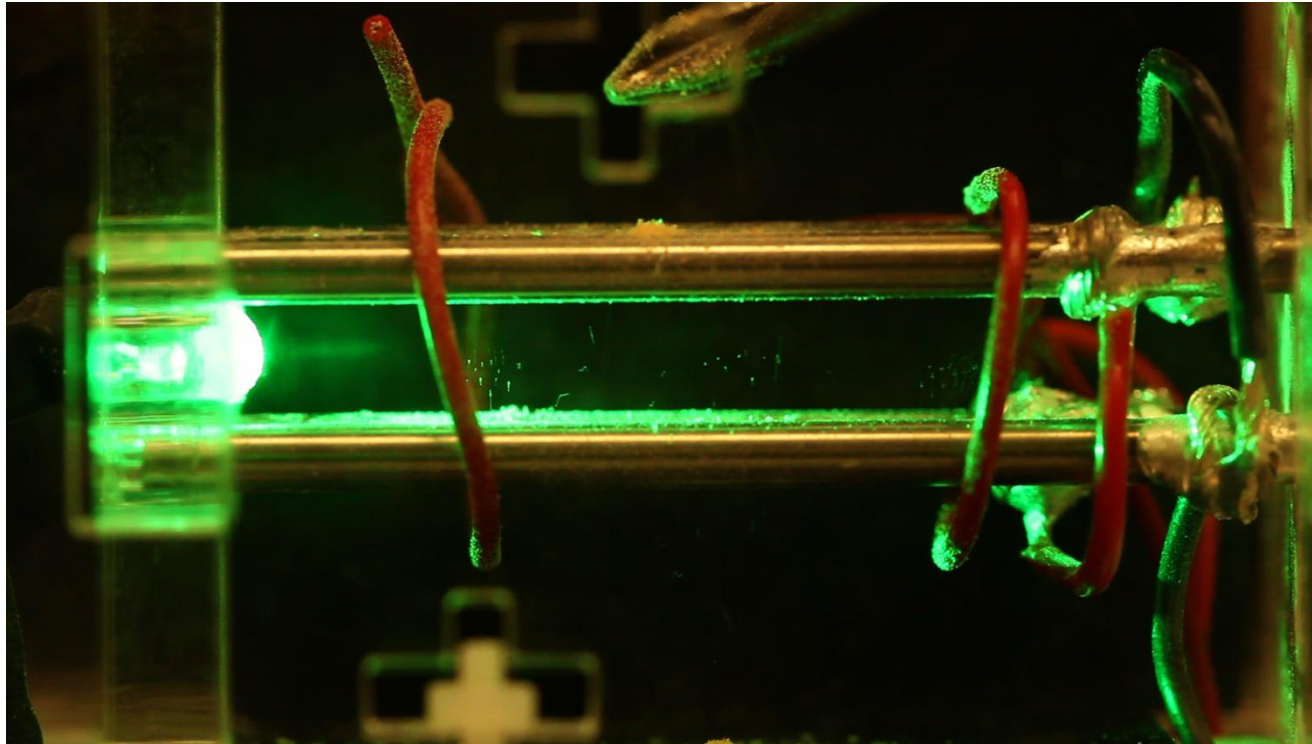


Angle Between  $\alpha$ 's



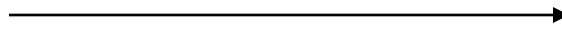
- $e^-$  takes 1/4 of decay energy
- $e^-$  takes 1/2 of decay energy
- $e^-$  takes 3/4 of decay energy

# Paul Traps

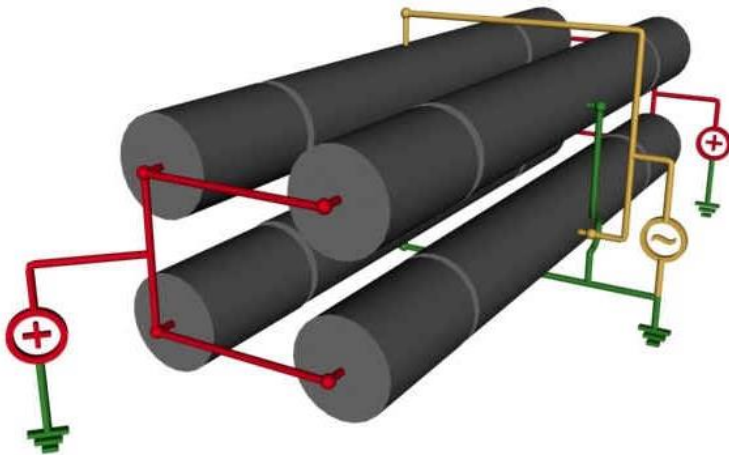


# Trap Geometry

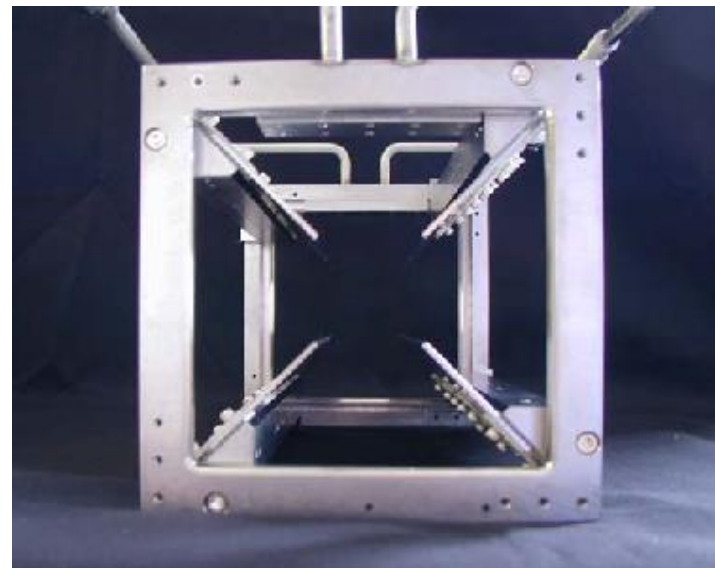
To get open geometry



- Standard Linear Paul Trap:  
Hyperbolic electrodes



- Planar Linear Ion Trap

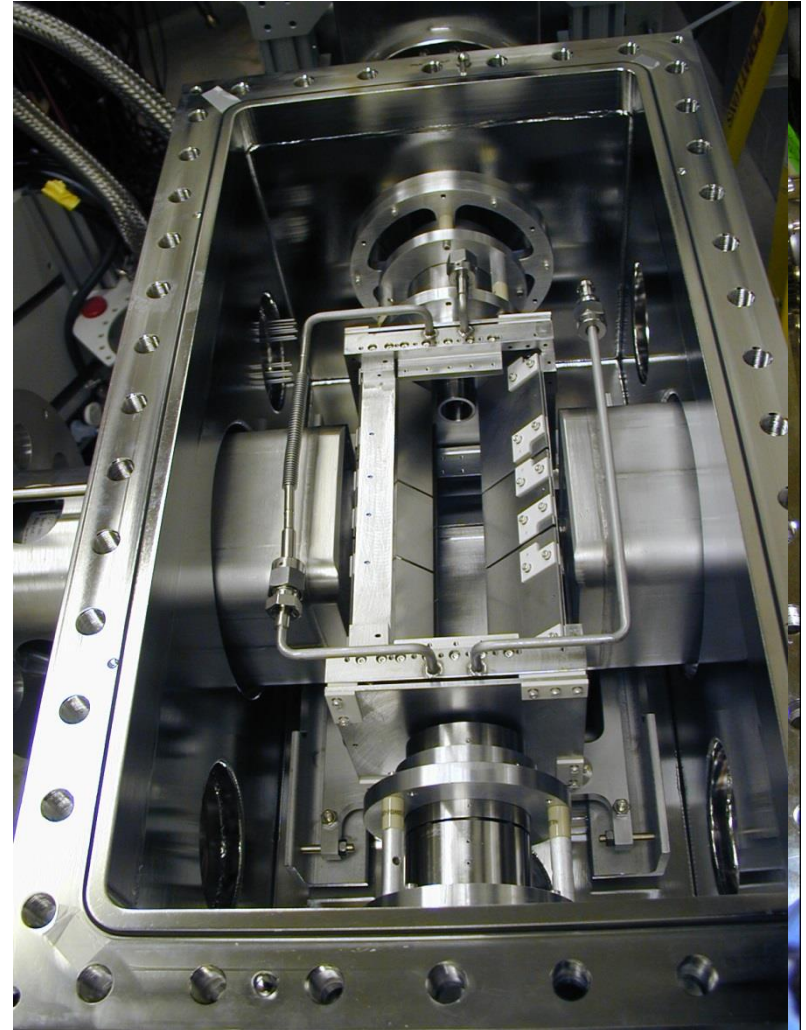


# The Beta-decay Paul Trap



- Confine up to  $\sim 10^6$  ions at once
- Hold for  $> 200$  sec
- Accessible half-life  $> 50$  ms
- Confine in  $\sim 1\text{-mm}^3$  volume
- DC fields of  $\sim 100\text{V}$
- RF fields of  $200\text{-}1000\text{ V}_{pp}$  at  $0.2\text{-}1.3\text{ MHz}$
- He buffer gas cools ions to  $\sim 0.1\text{ eV}$

N.D. Scielzo *et al.*, NIM A **681**, 94 (2012)



# Correlations determined from $\beta$ - $\alpha$ - $\alpha$ coincidences

Double-sided silicon strip detectors (DSSDs) used to determine:

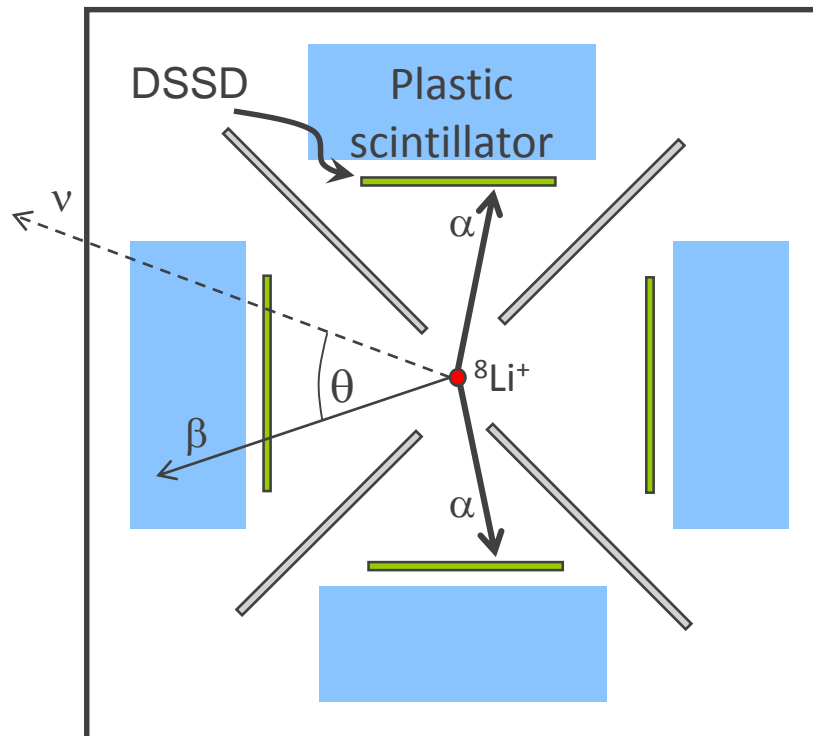
- $\beta$  momentum direction
- Momentum and energy for both  $\alpha$  particles

This is sufficient to fully reconstruct the decay kinematics

- $E_{\alpha 1} + E_{\alpha 2}$  :  ${}^8\text{Be}$  excitation energy
- $p_{\alpha 1} + p_{\alpha 2}$  : nuclear recoil

Additional measurement of  $\beta$  energy not required... but overconstrains the kinematics

trapped ions surrounded by DSSDs and plastic scintillators



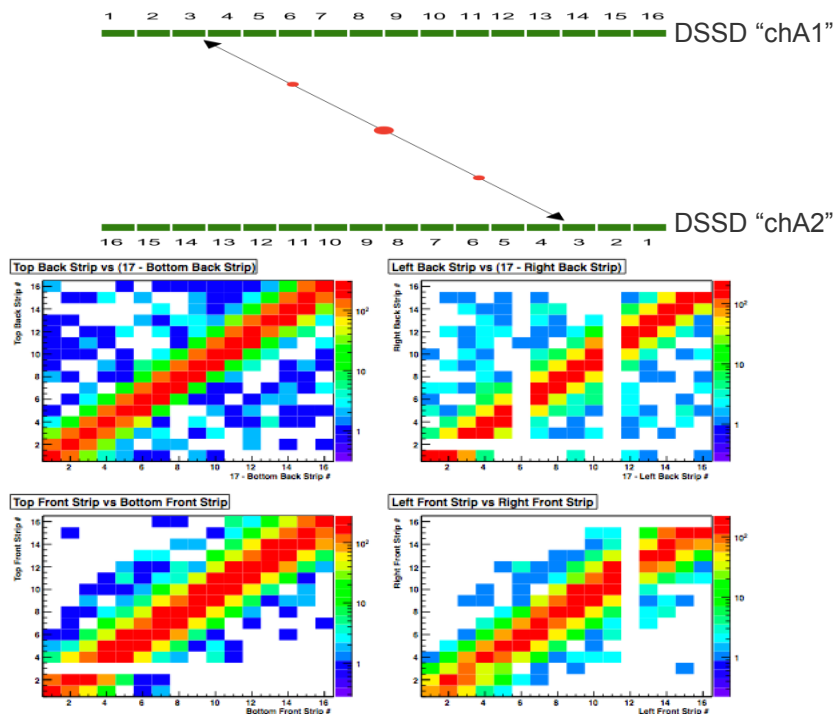
DSSDs:

- 1-mm thick, 64x64-mm<sup>2</sup> detector, 2-mm strips
- ~100-nm dead layer
- Continuous *in situ* calibration using  ${}^{148}\text{Gd}$  and  ${}^{244}\text{Cm}$   $\alpha$  sources



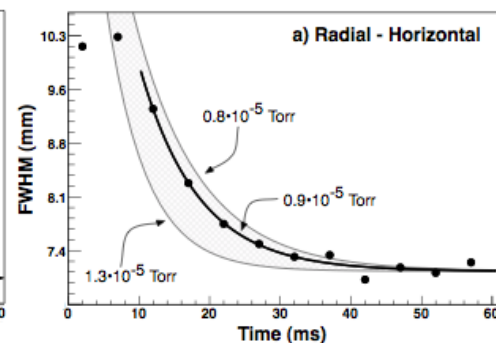
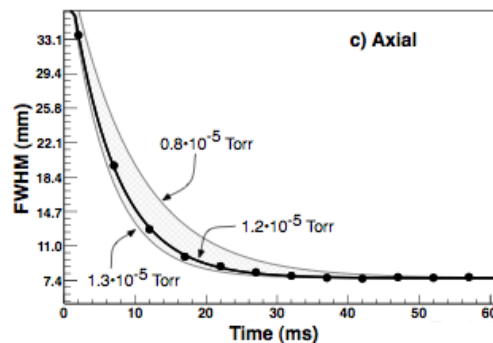
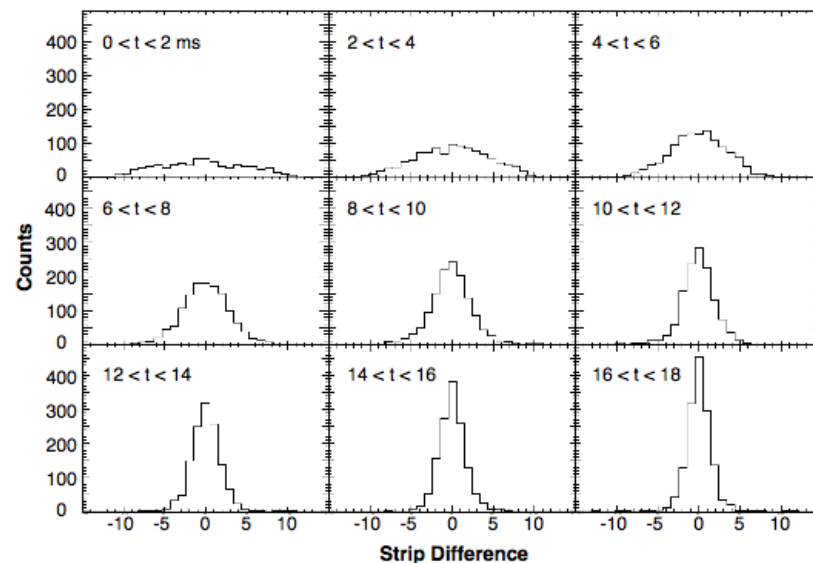
# First Results with $\beta$ - $\alpha$ - $\alpha$ Coincidences

## Imaging the ion cloud with back-to-back $\alpha$ 's



- Slight broadening due to resolution of strips & angular broadening of  $\alpha$ 's from recoil
- Image is consistent with ion cloud  $\sim 1 \text{ mm}^3$
- Finer strips will provide more precise imaging

- Looking at the strip differences over time we can watch the ion cloud cool



# Analysis depends on high-fidelity simulations...

## Beta-decay Event Generator

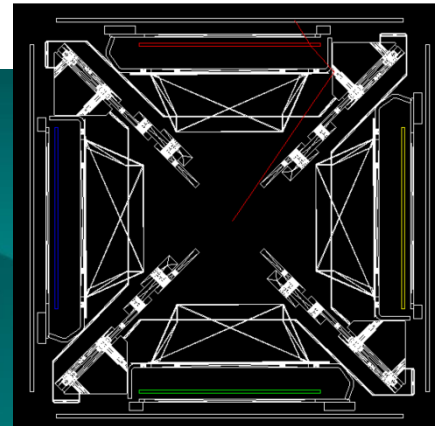
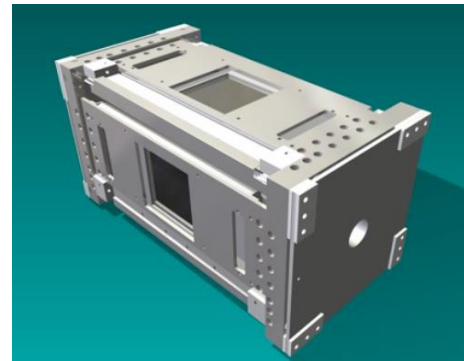
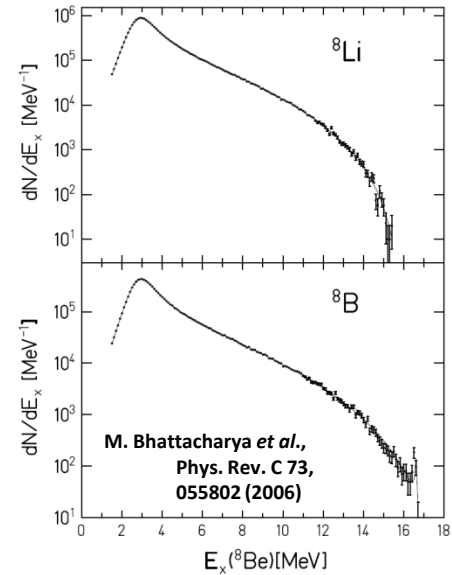
- Beta-decay phase space and angular correlations
- Final-state distribution of  ${}^8\text{Be}^*$
- Recoil-order terms and radiative corrections
- Ion cloud distribution

## Apparatus Simulation

- Propagate  $\beta$  particles through experimental geometry using GEANT4
- DSSD detector response (resolution and dead layer) incorporated

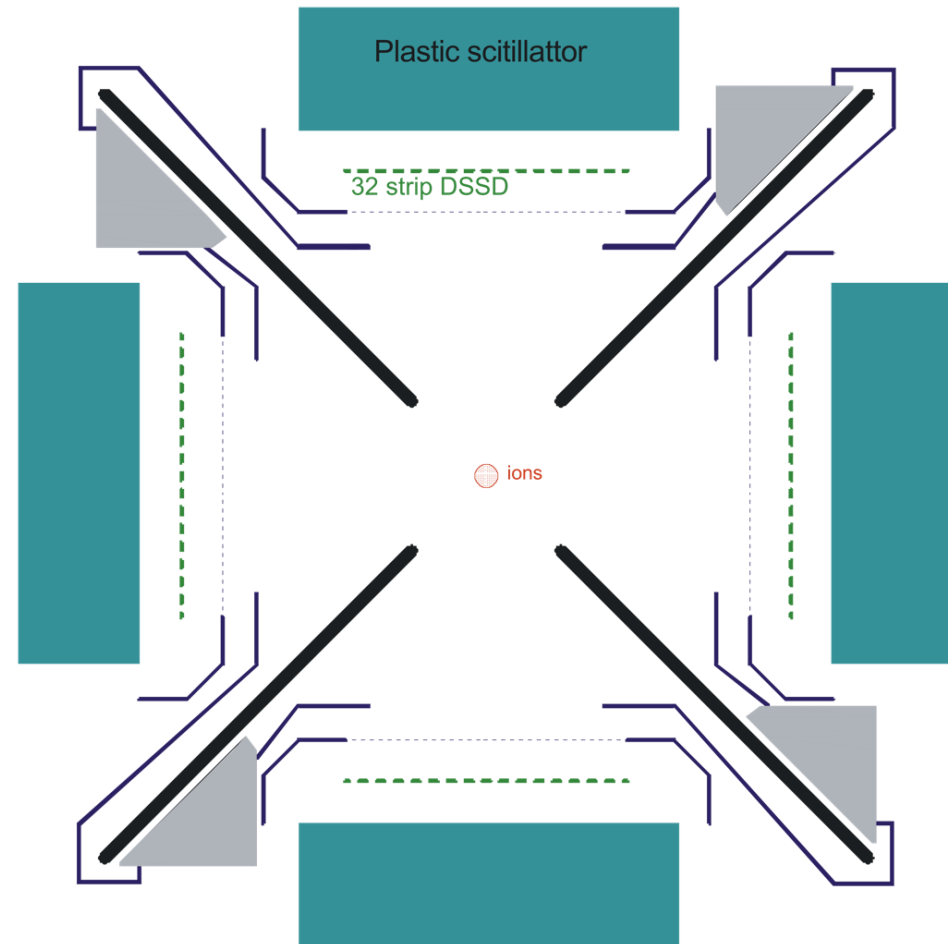
$$\begin{aligned}
 d^3\Gamma = & F_{\mp}(Z, E) \frac{G_v^2 \cos^2\theta_c}{2(2\pi)^6} (E_0 - E)^2 p E dE d\Omega_e d\Omega_n \\
 & \times \left( g_1(E) + g_2(E) \frac{\mathbf{p} \cdot \hat{k}}{E} + g_3(E) \left[ \left( \frac{\mathbf{p} \cdot \hat{k}}{E} \right)^2 - \frac{1}{3} \frac{p^2}{E^2} \right] \right. \\
 & + \delta_1(E, v^*, \tau_{J', J''}(L)) \frac{\hat{n} \cdot \mathbf{p}}{E} \\
 & + \delta_2(E, v^*, \tau_{J', J''}(L)) \hat{n} \cdot \frac{\mathbf{p} \cdot \mathbf{p}}{E E} \hat{k} \\
 & + \delta_3(E, v^*, \tau_{J', J''}(L)) \hat{n} \cdot \hat{k} \\
 & + \delta_4(E, v^*, \tau_{J', J''}(L)) \hat{n} \cdot \hat{k} \frac{\mathbf{p} \cdot \hat{k}}{E} \\
 & + \frac{1}{\sqrt{6}} \tau_{J', J''}(L) T^{(2)}(\hat{n}) : \left\{ g_{10}(E) [\mathbf{p}/E, \mathbf{p}/E] \right. \\
 & + g_{11}(E) [\mathbf{p}/E, \mathbf{p}/E] \frac{\mathbf{p} \cdot \hat{k}}{E} + g_{11}(E) [\mathbf{p}/E, \hat{k}] \\
 & + g_{12}(E) [\mathbf{p}/E, \hat{k}] \frac{\mathbf{p} \cdot \hat{k}}{E} + g_{13}(E) [\hat{k}, \hat{k}] \\
 & + g_{16}(E) [\hat{k}, \hat{k}] \frac{\mathbf{p} \cdot \hat{k}}{E} + g_{16}(E) \left[ \frac{\mathbf{p}}{E}, \frac{\mathbf{p}}{E} \times \hat{k} \right] \\
 & + g_{17}(E) \left[ \hat{k}, \frac{\mathbf{p}}{E} \times \hat{k} \right] \left. \right\} \\
 & + \delta_8(E, v^*, \tau_{J', J''}(L)) T^{(3)}(\hat{n}) : [\mathbf{p}/E, \mathbf{p}/E, \hat{k}] \\
 & + \delta_9(E, v^*, \tau_{J', J''}(L)) T^{(3)}(\hat{n}) : [\mathbf{p}/E, \hat{k}, \hat{k}] \\
 & + \frac{1}{\sqrt{6}} \omega_{J', J''}(L) T^{(4)}(\hat{n}) : \left\{ g_{25}(E) [\mathbf{p}/E, \mathbf{p}/E, \mathbf{p}/E, \hat{k}] \right. \\
 & + g_{26}(E) [\mathbf{p}/E, \mathbf{p}/E, \hat{k}, \hat{k}] \\
 & \left. + g_{27}(E) [\mathbf{p}/E, \hat{k}, \hat{k}, \hat{k}] \right\} . \tag{53}
 \end{aligned}$$

B.R. Holstein, RMP 46, 789 (1974)



# 2011 Detector System

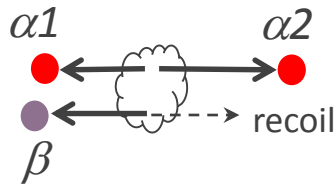
- 64 mm X 64 mm - 32 X 32 strips
  - ~30% solid angle
  - 1 mm thick
  - 100 nm Dead Layer
  - Angular resolution - 3 degrees
  - 10% efficiency for  $\beta$ - $\alpha$ - $\alpha$
- Scintillating Plastic for  $\beta$ 's
- Modified RF shielding
- Modified electrode design
  - Solid electrode design blocks source alphas at large angles
  - RF is only applied at tip of electrode to reduce pickup on DSSD
- There are 8 degrees of freedom
- We measure 9 - system is over constrained!



# $^8\text{Li}$ : 1<sup>st</sup> improved limits on $|C_T/C_A|^2$ from $\beta$ decay in 50 years

M.G. Sternberg *et al.*, PRL **115**, 182501 (2015)

Most sensitive measure of correlation from  $E_\alpha$  difference when  $\beta$  parallel to  $\alpha$

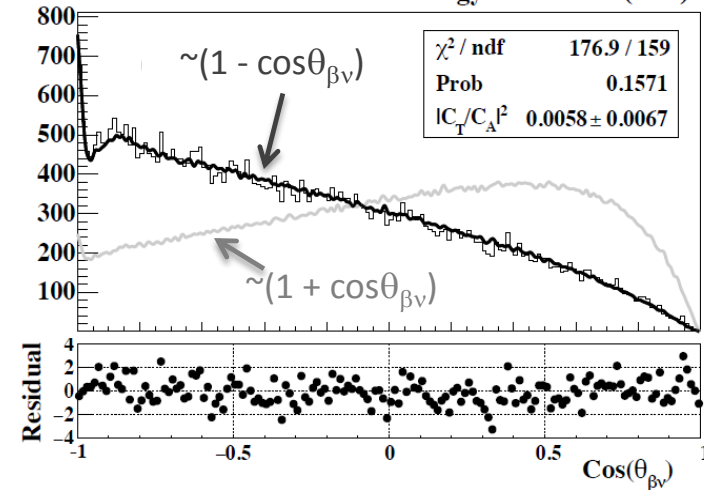
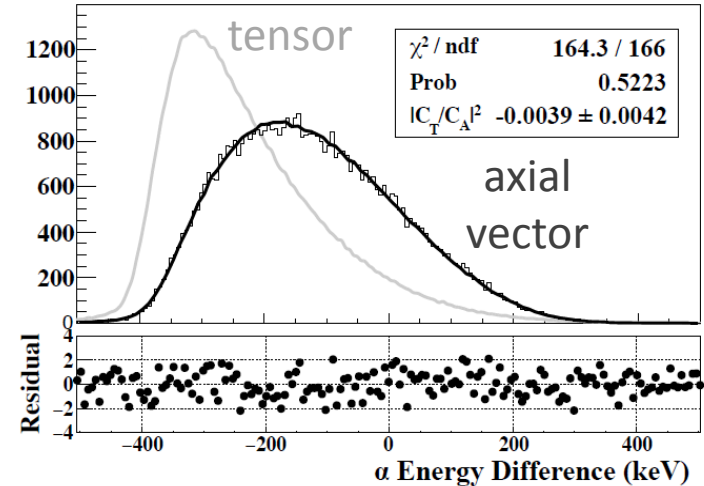


**Axial vector:** leptons preferentially emitted in opposite directions  
 $\rightarrow$  smaller recoil, smaller  $\Delta E_\alpha$

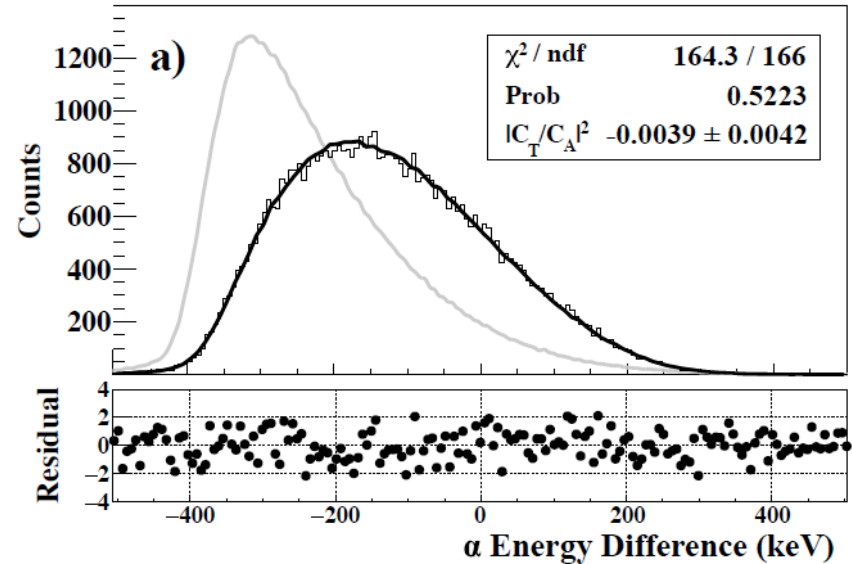
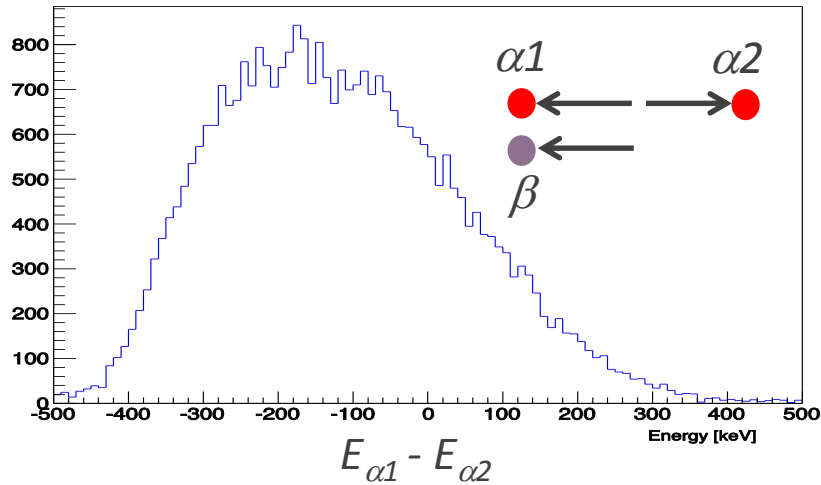
**Tensor:** leptons preferentially emitted in same direction  
 $\rightarrow$  larger recoil, larger  $\Delta E_\alpha$

$$a_{\beta\nu} = -0.3342 \pm 0.0026_{stat} \pm 0.0029_{syst}$$

$$|C_T/C_A|^2 = -0.0013 \pm 0.0038_{stat} \pm 0.0043_{syst}$$



# $\beta$ - $\alpha$ - $\alpha$ coincidences from ${}^8\text{B}$ decay under analysis



$\beta$ -decay angular correlations comparable to earlier  ${}^8\text{Li}$

Precise determination of  ${}^8\text{B}$  solar neutrino spectrum

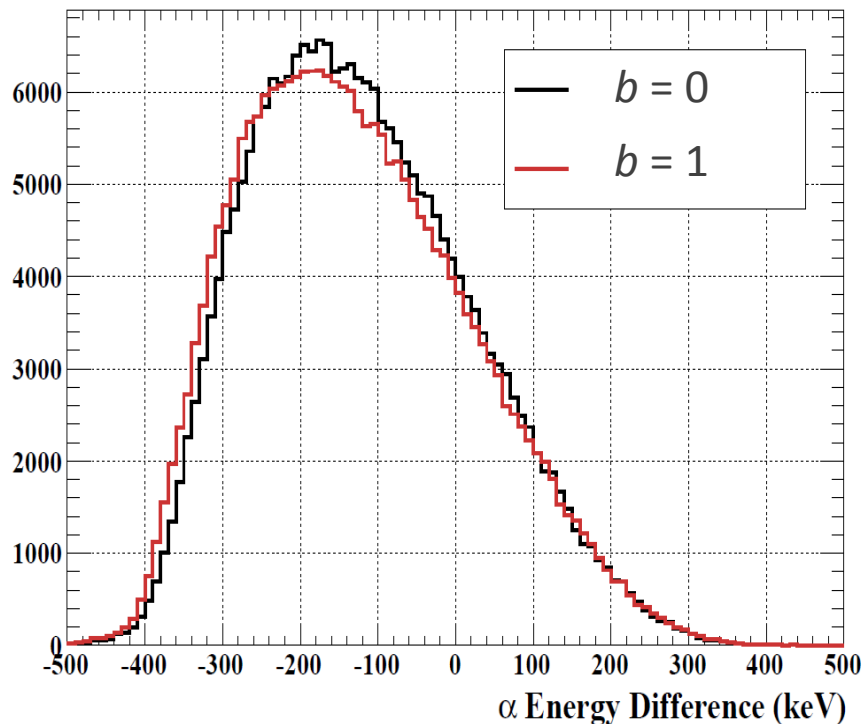
Comparison to  ${}^8\text{Li}$  data  $\rightarrow$  recoil-order terms (small  $E_e$  dependence)

- Weak magnetism (change sign with  $\beta^\pm$ )
- Second-class currents (independent of  $\beta^\pm$ )

Fully-reconstructed  ${}^8\text{Li}/{}^8\text{B}$  decays allows determination of recoil-order terms in several correlations

Also on the horizon...

## A search for the Fierz interference term



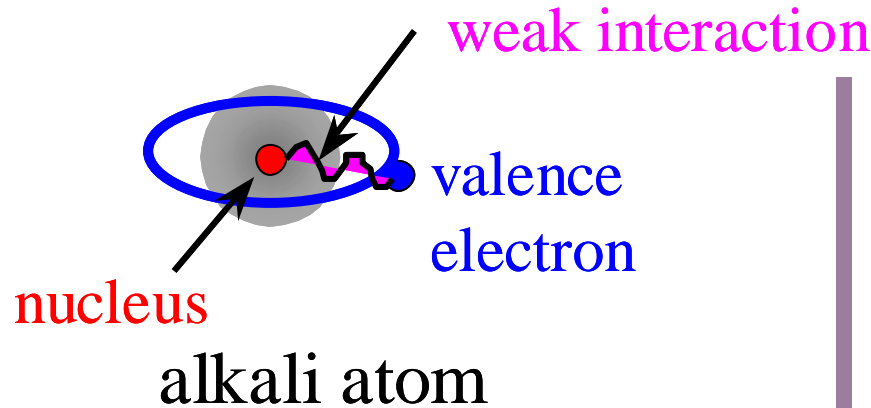
$$dW \sim \left[ 1 + a \frac{\vec{p}_e \cdot \vec{p}_\nu}{E_e E_\nu} + b \frac{m_e}{E_e} \right]$$

$\sim 0.001$        $\sim 0.01$

Sensitivity to Fierz interference term is  $\sim 10\times$  less than for tensor contribution

Systematics need to be carefully investigated

# The weak interaction between the outer electron and the nucleus is predicted by the Standard Model



$$|nS\rangle' = |nS\rangle + \delta|nP\rangle$$

$$H_W = \frac{G_F}{\sqrt{8}} Q_W \rho_{nuc}(r) \gamma_5$$

$$\delta^M = b \left[ -V + \underbrace{\Sigma (1 - \gamma_5 \theta^M)}_{\sim 0} \right]$$

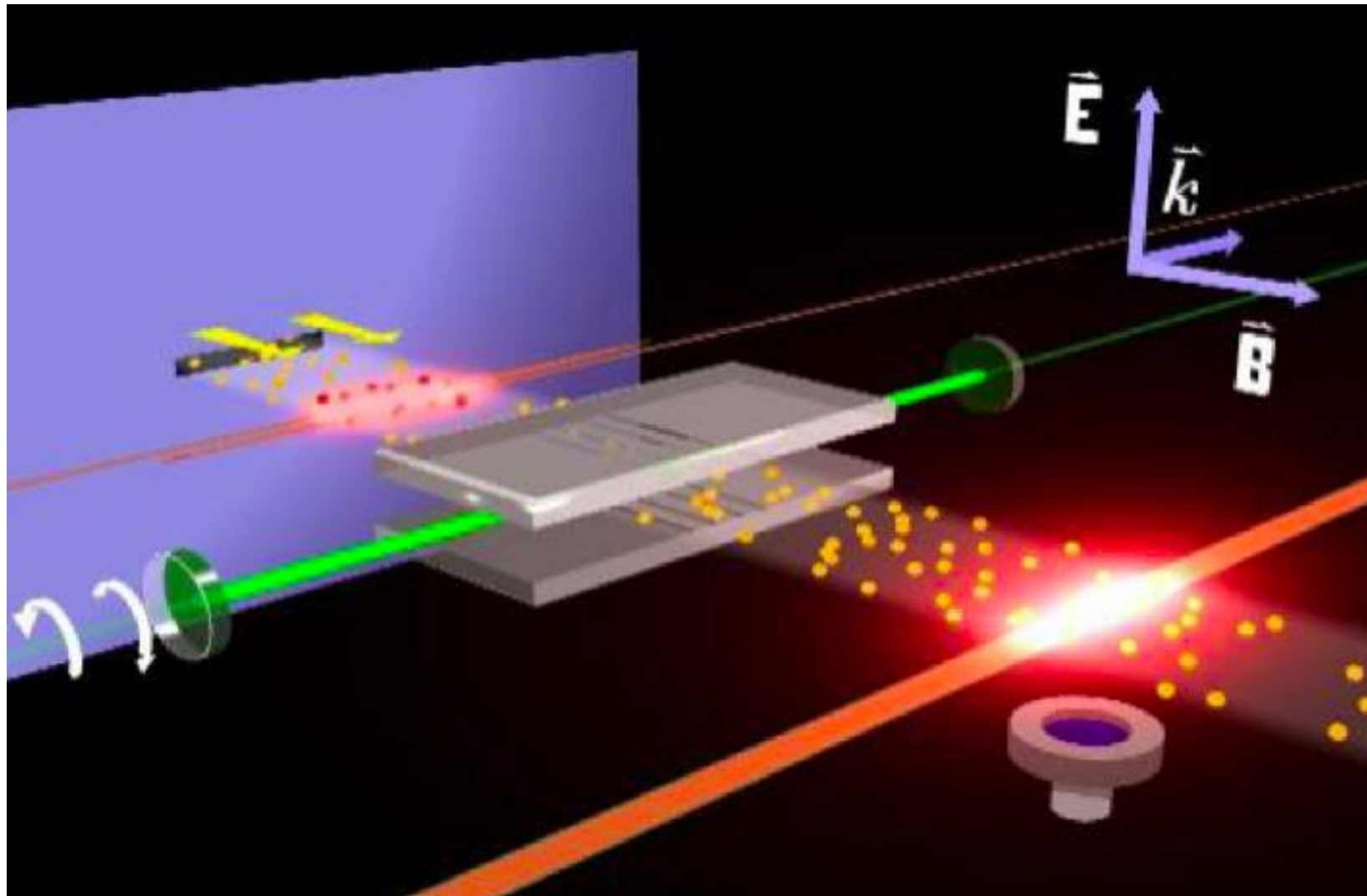
$$Q_W(\text{Experiment})^* = -72.06 \pm 0.28_{\text{exp}} \pm 0.34_{\text{th}}$$

$$Q_W(\text{Standard Model}) = -73.20$$

\*PRL 82 (1999) 2484

# The Boulder $Cs$ PNC Experiment

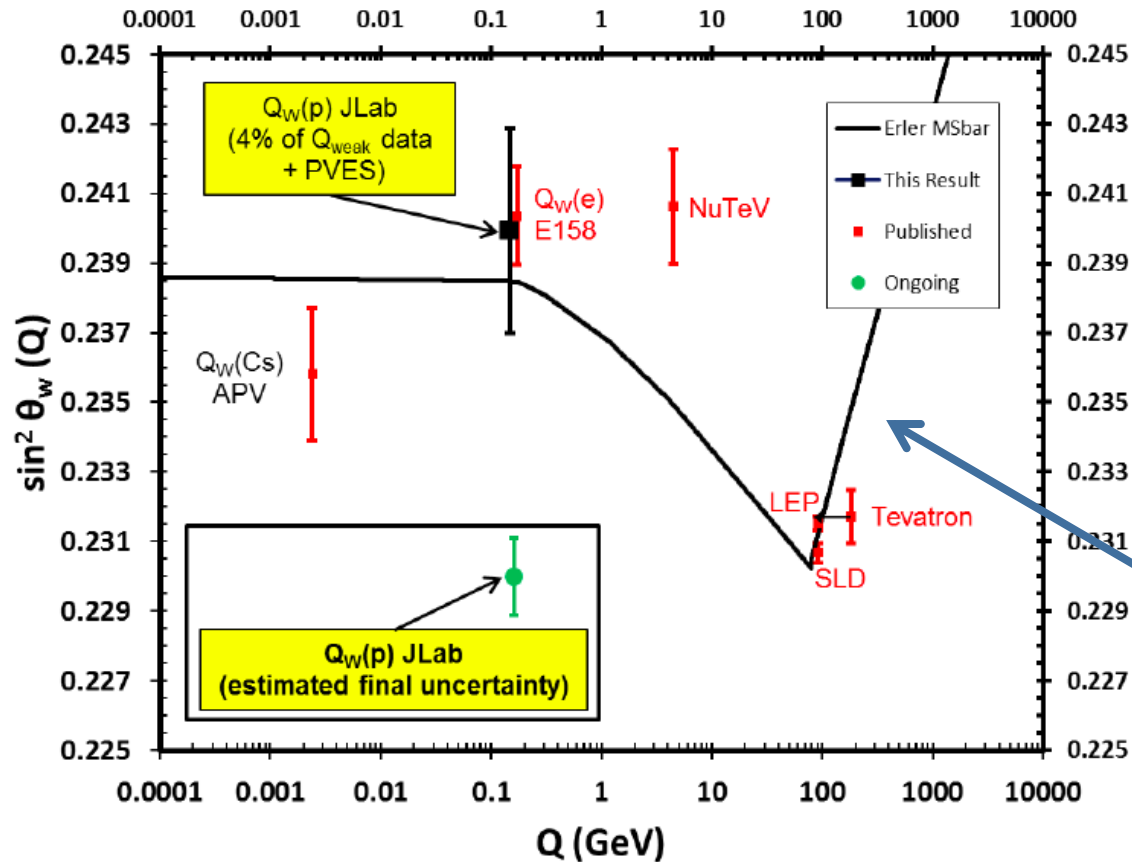
1982-1999



- P-odd, T-even correlation:  $\sigma \cdot [E \times B]$
- 5 reversals to distinguish PNC from systematics



# Parity non-conservation in atoms



An example of the so called running of fundamental constants.

Prediction of the Standard Electro-Weak Theory

A credible path to necessary improvements on parity non-conservation in atoms requires an intense source of Fr atoms:

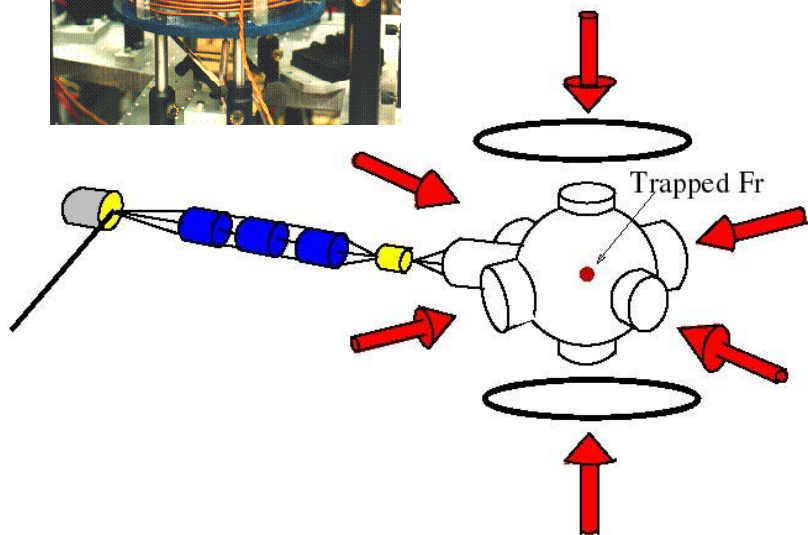
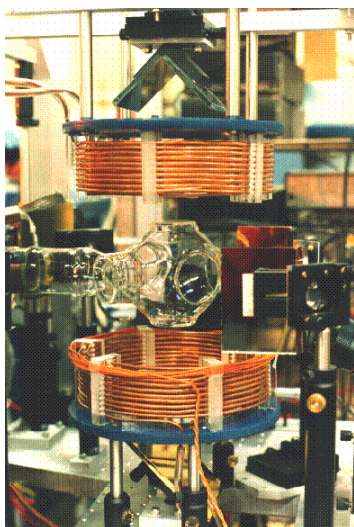
- X18 larger signal
- atomic structure independent
- neutron distribution ....  $^{208}\text{Pb}$  at JLAB and hyperfine anomaly

# Francium-a new laboratory for fundamental symmetry tests

87	Fr [223]	Fr200	Fr201	Fr202	Fr203	Fr204	Fr205	Fr206	Fr207	Fr208	Fr209	Fr210	Fr211	Fr212
		48 ms (9/2-) EC,α	0.34 s (9/2-) EC,α	0.55 s (9/2-) EC,α	1.7 s (3+) EC,α *	3.85 s (9/2-) EC,α	15.9 s (5+) EC,α *	14.8 s 9/2- EC,α	59.1 s 7+ EC,α	50.0 s 9/2- EC,α	3.18 m 6+ EC,α	3.18 m 9/2- EC,α	20.0 m 5+ EC,α	

Fr213	Fr214	Fr215	Fr216	Fr217	Fr218	Fr219	Fr220	Fr221	Fr222
34.6 s 9/2- EC,α	5.0 ms (1-) α *	86 ns 9/2- α	0.70 μs (1-) EC,α	22 μs 9/2- α	1.0 ms 1- α *	20 ms 9/2- α	27.4 s 4+ β,α	4.9 m 5/2- α	14.2 m 2- β

Fr223	Fr224	Fr225	Fr226	Fr227	Fr228	Fr229	Fr230	Fr231	Fr232
21.8 m 3/2(-) β,α	3.33 m 1- β	4.0 m 3/2- β	49 s 1- β	2.47 m 1/2+ β	38 s 2- β	50 s β	19.1 s β	17.5 s β	5 s β



A systematic study of Parity non-conservation for different isotopes in Francium would lead to greatly enhanced sensitivity to new physics, complementing high energy physics studies. To study a wide range of isotopes requires ISAC or FRIB.

A possible experimental approach:

1. **Capture** Fr atoms in a MOT
2. **Accumulate** and cool in the MOT
3. **Transfer** to a second trap (purely optical)
4. Establish a “**coordinate system**” by dc electric field, dc magnetic field, k vector of the exciting laser
5. **Excite** 7S to 8S using a build up cavity and detect using the 7S to 7P transition.
6. **Reverse** the **coordinate** axis.
7. **Change isotope.**

From Luis Orozco Les Houche 2000

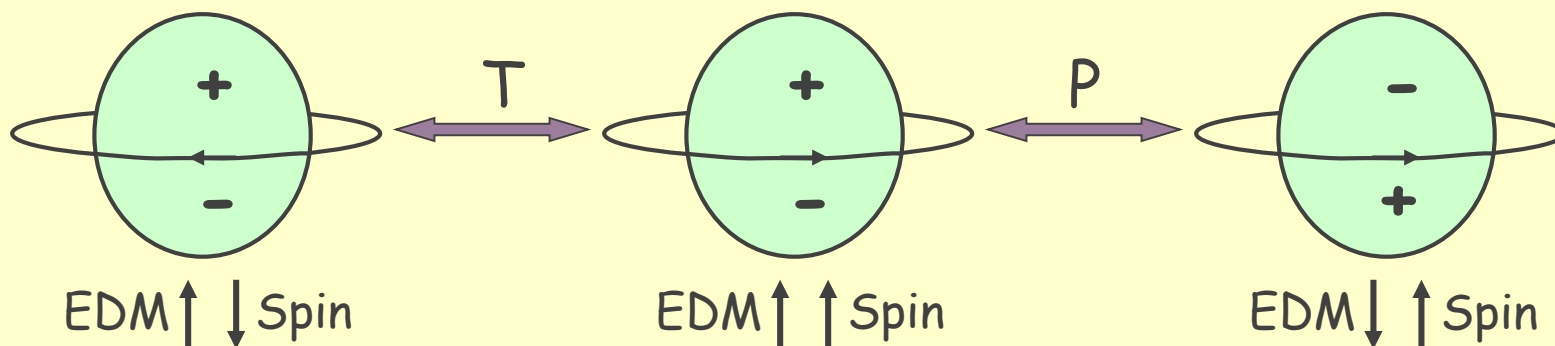


# CP Violation

- Astronomical Observations indicate that the universe is dominated by matter.
- The CP violation observed in K and B decays is not sufficient to explain the observed asymmetry of matter.
- There are many proposed theoretical models with CP violation.
- The most stringent test for models is the Electric Dipole Moments they predict for the neutron, the electron and for nuclei.
- RIB facilities allow these measurements of particular nuclei in which the moments are greatly enhanced.

# Search for an electric dipole moment and physics beyond the standard model

A permanent EDM violates both time-reversal symmetry and parity



**To understand the origin of the symmetry violations, you need many experiments!**

Neutron

Quark EDM

Diamagnetic Atoms  
(Hg, Xe, Ra, Rn)

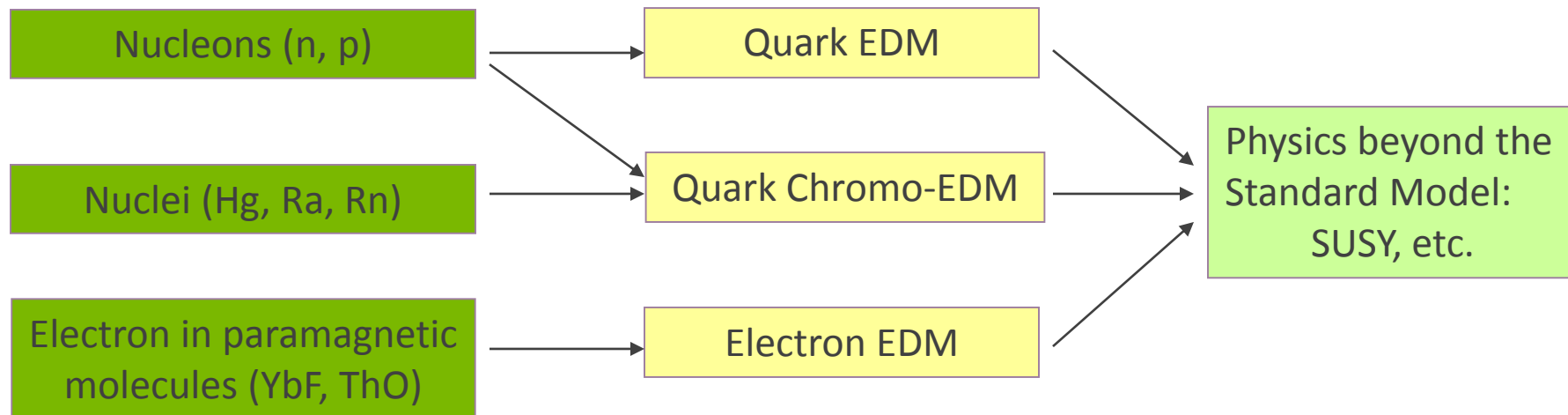
Quark Chromo-EDM

Physics beyond  
the Standard  
Model:  
SUSY, Strings ...

Paramagnetic Atoms (Tl, Fr)  
Molecules (PbO)

Electron EDM

# EDM Searches in Three Sectors



Sector	Exp Limit (e-cm)	Method	Standard Model
Electron	$9 \times 10^{-29}$	ThO in a beam	$10^{-38}$
Neutron	$3 \times 10^{-26}$	UCN in a bottle	$10^{-31}$
$^{199}\text{Hg}$	$3 \times 10^{-29}$	Hg atoms in a cell	$10^{-33}$

M. Ramsey-Musolf (2009)

Neutron EDM; Ramsey,  
Dress et al. Phys. Rep. 43, 410 (1978).

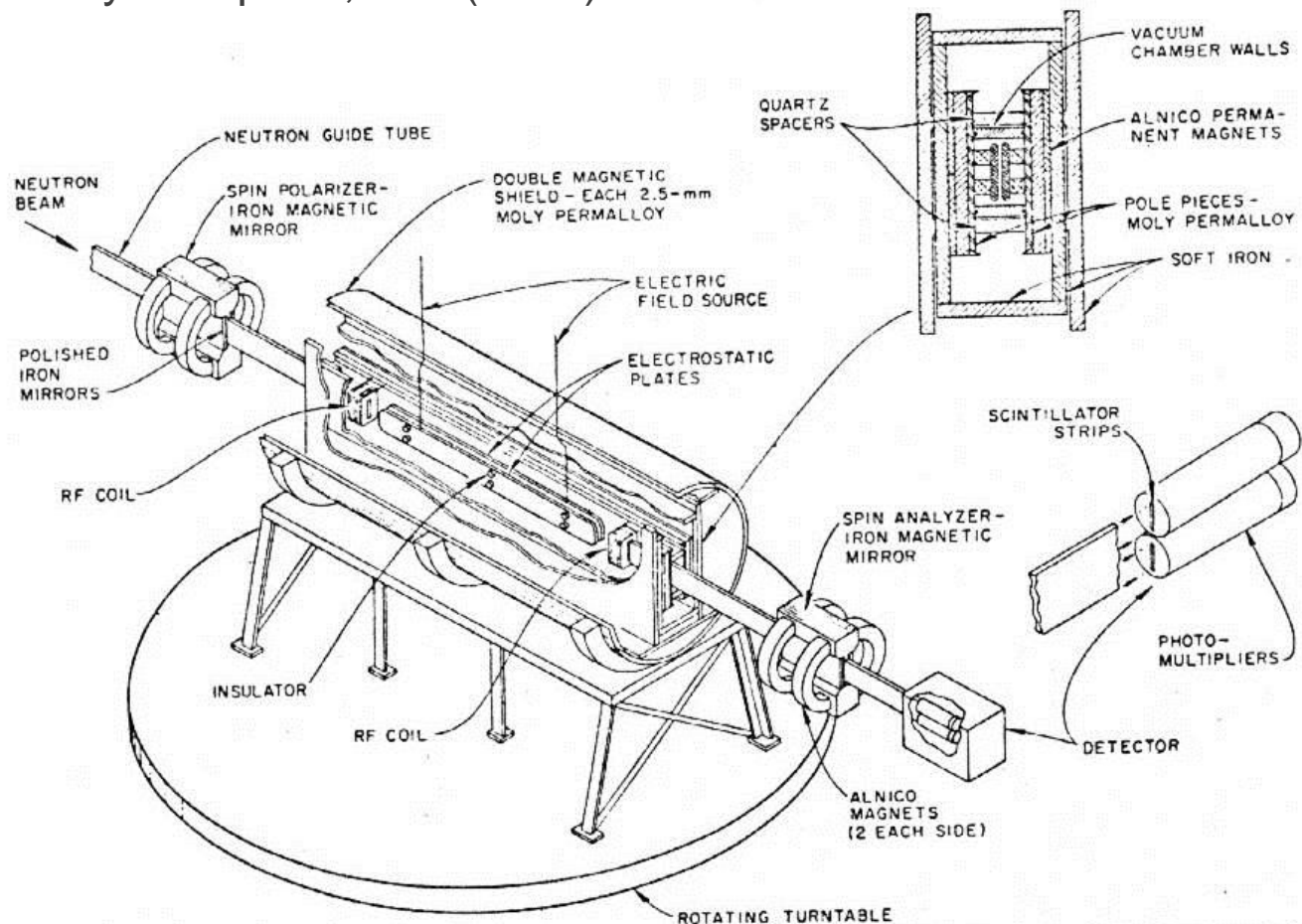
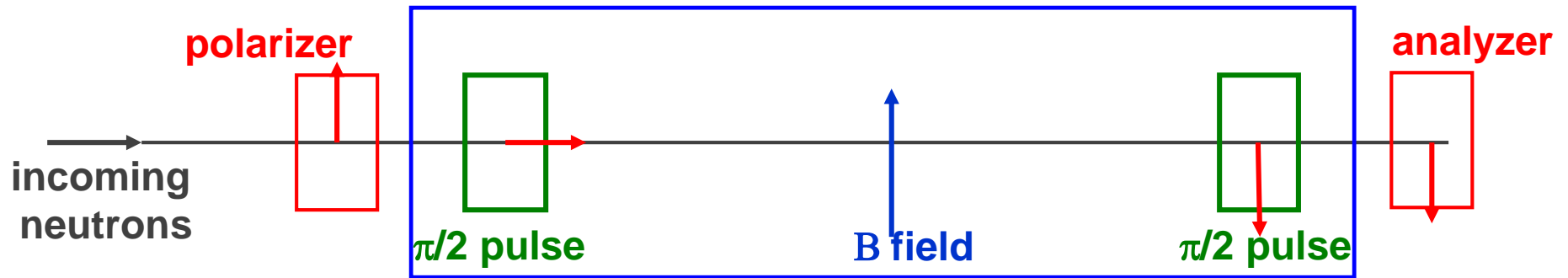


FIGURE 1 Pictorial representation of the neutron-beam spectrometer. The insert, upper right, is a cross-sectional view through the midpoint of the apparatus indicating various materials used in construction. The gap between the magnetic poles is 9 cm.

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Neutron EDM; Ramsey,  
Dress et al. Phys. Rep. 43, 410 (1978).

Simplified version





Neutron EDM; Ramsey,  
Dress et al. Phys. Rep. 43, 410 (1978).

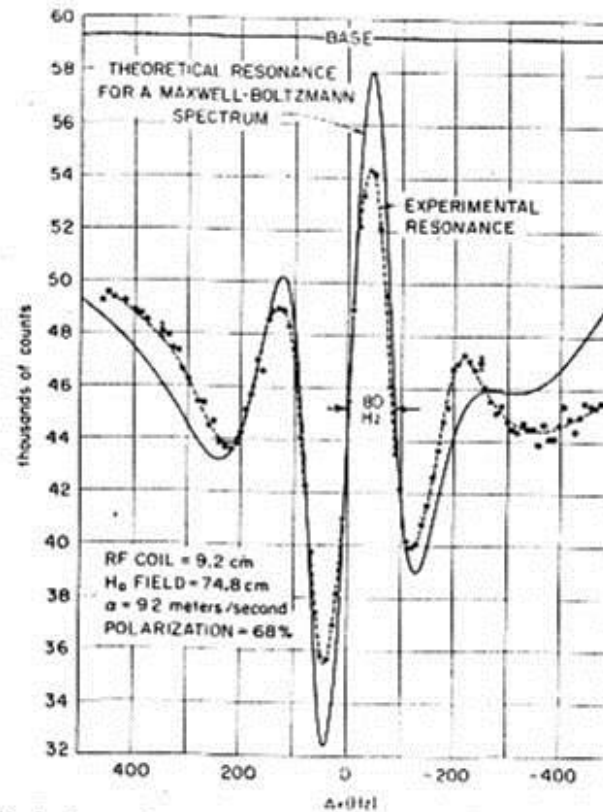
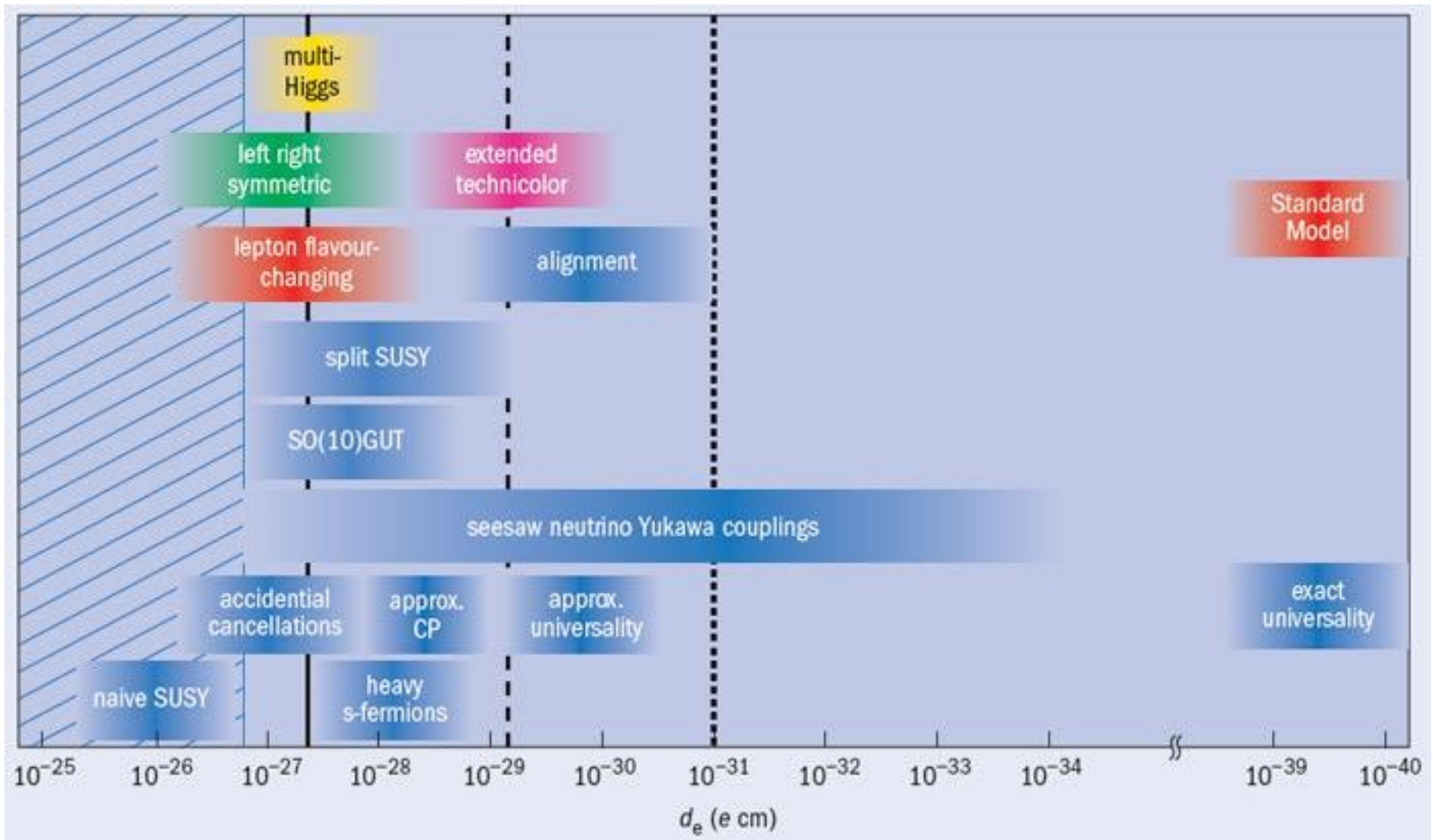


FIGURE 2 Typical magnetic resonance with the neutron-beam apparatus for a phase shift of  $\pi/2$  between the two oscillatory fields. The calculated transition probability for a Maxwell-Boltzmann distribution characterized by a temperature of 1 K is shown by the solid curve. The departure of the experimental curve from theory when far from resonance is to be expected from the known departure of the beam velocity from a Maxwell-Boltzmann distribution.

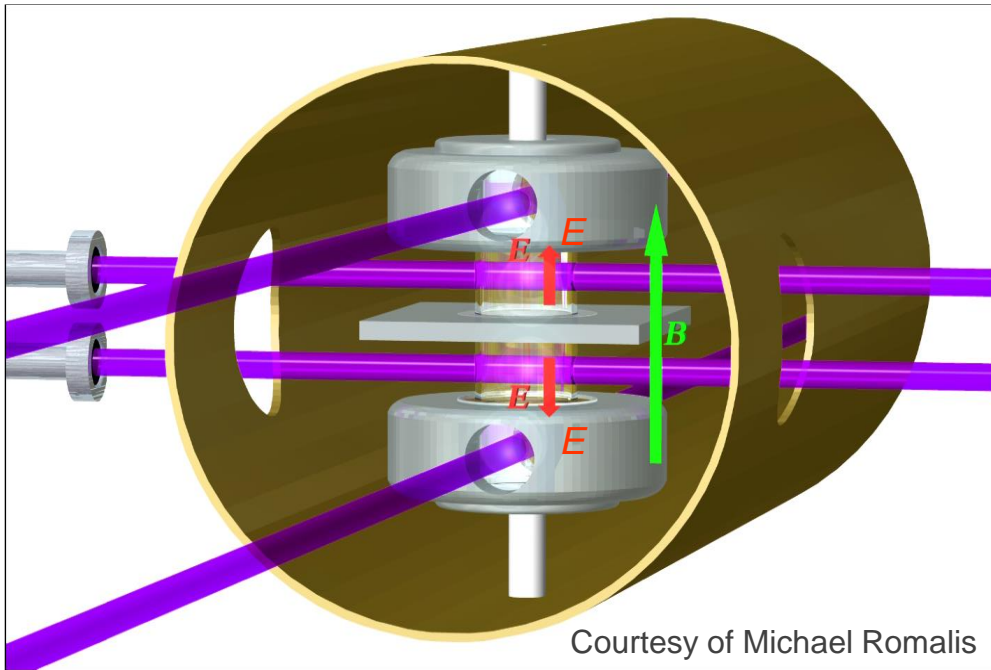
# EDM ... the Standard Model extension slayer



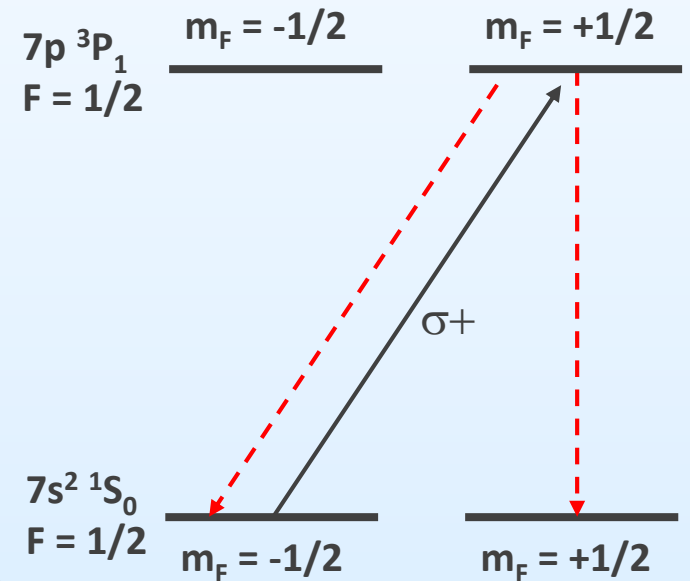
# The Seattle EDM Measurement

$^{199}\text{Hg}$

stable, high Z, groundstate  $^1S_0$ ,  $I = 1/2$ , high vapor pressure

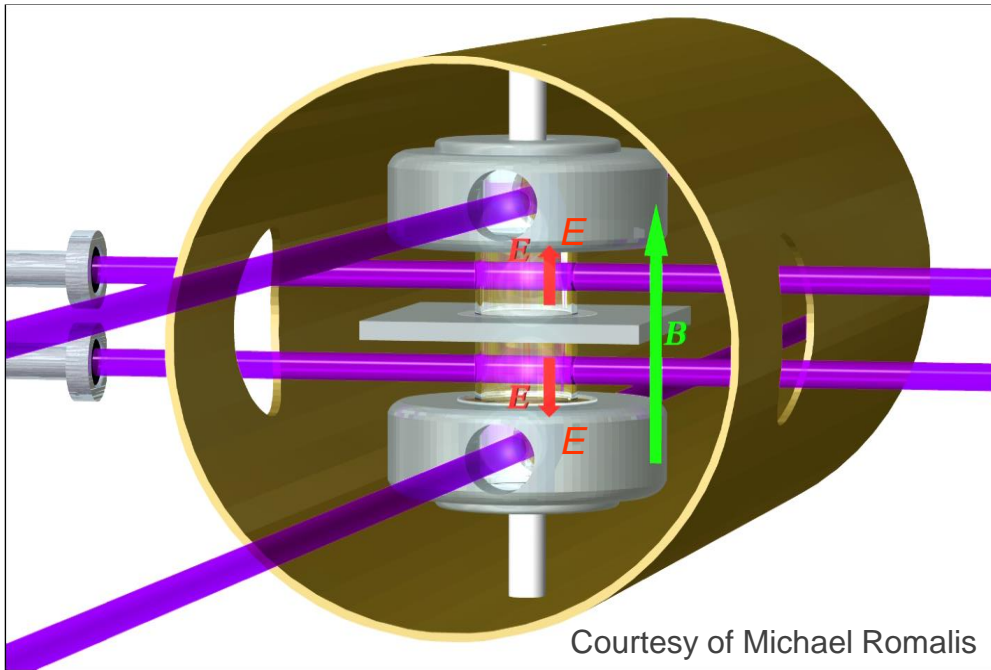


## Optical Pumping



# The Seattle EDM Measurement

$^{199}\text{Hg}$  stable, high Z, groundstate  $^1\text{S}_0$ ,  $I = \frac{1}{2}$ , high vapor pressure



Courtesy of Michael Romalis

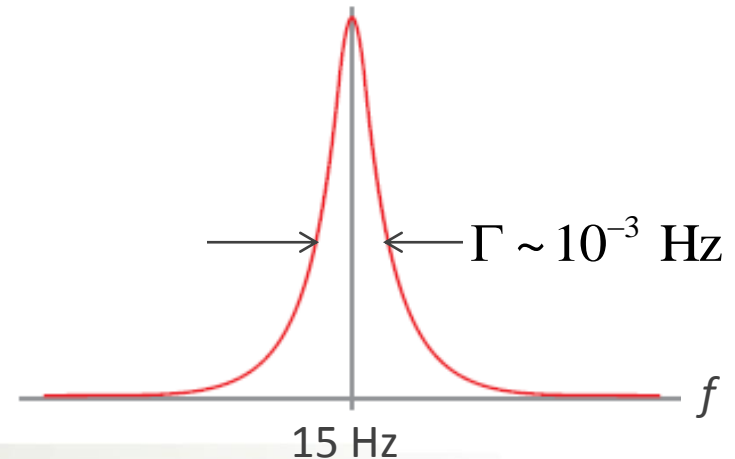
$$f_+ = \frac{2\mu B + 2dE}{h} \approx 15 \text{ Hz}$$

$$f_- = \frac{2\mu B - 2dE}{h} \approx 15 \text{ Hz}$$

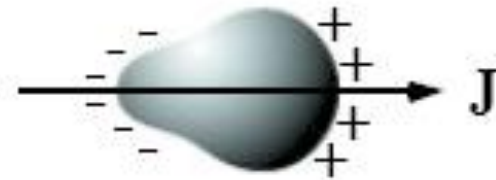
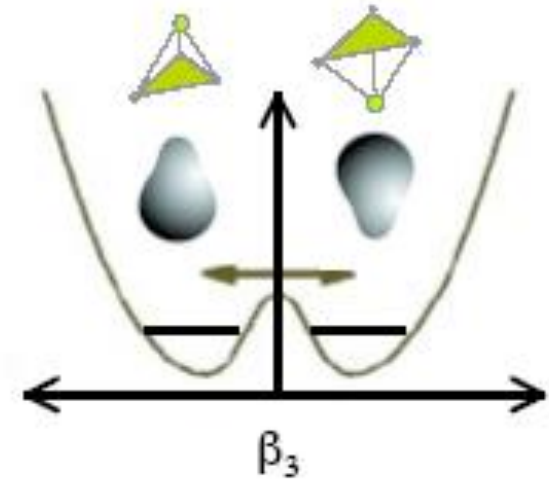
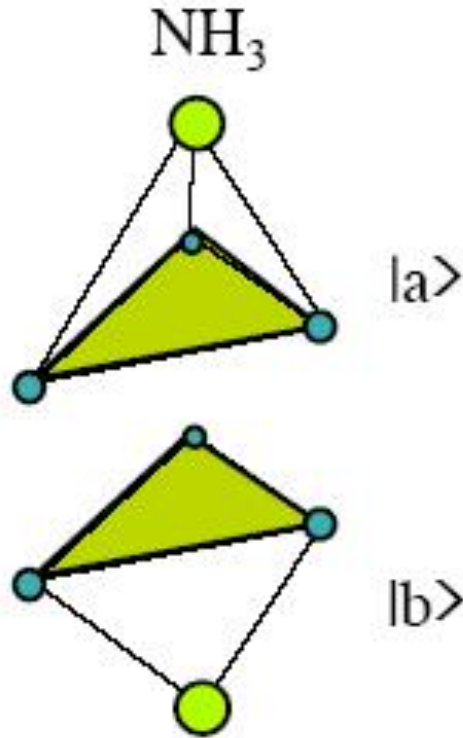
$$|f_+ - f_-| < 1 \times 10^{-10} \text{ Hz}$$

## Limits and Sensitivities

- Current:  $< 3 \times 10^{-29}$  e-cm  
-- Griffith *et al.*, PRL (2009)
- Next 5 years:  $3 \times 10^{-30}$  e-cm
- Beyond 2020:  $6 \times 10^{-31}$  e-cm



# Enhancement of nuclear EDM sensitivity in octupole deformed parity doublets

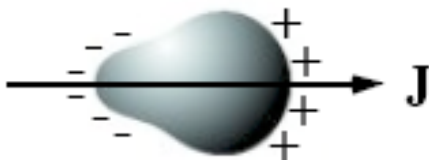


$$|\psi_{\pm}\rangle = \frac{1}{\sqrt{2}} (|a\rangle \pm |b\rangle)$$

$$S \sim \frac{\langle +|\eta r^3 \cos \theta|-\rangle}{E_+ - E_-} \sim \frac{\eta \beta_2 \beta_3^2 Z A^{2/3} r_0^3}{E_+ - E_-}$$

# Nuclei with Octupole Deformation/Vibration

(Haxton & Henley; Auerbach, Flambaum, Spevak; Engel, Hayes & Friar, etc.)

$$S \sim \frac{\langle +|\eta r^3 \cos \theta|-\rangle}{E_+ - E_-} \sim \frac{\eta \beta_2 \beta_3^2 Z A^{2/3} r_0^3}{E_+ - E_-}$$


	223Rn	223Ra	225Ra	223Fr	225Ac	229Pa	199Hg	129Xe
$t_{1/2}$	23.2 m	11.4 d	14.9 d	22 m	10.0 d	1.5 d		
I	7/2	3/2	1/2	3/2	3/2	5/2	1/2	1/2
$\Delta E_{th}^*$ (keV)	37†	170	47	75	49	5		
$\Delta E_{exp}$ (keV)	—	50.2	55.2	160.5	40.1	0.22		
$10^5 S$ (efm <sup>3</sup> )	1000	400	300	500	900	12000	-1.4	1.75
$10^{28} d_A$ (e-cm)	<b>3300</b>	<b>3300</b>	<b>2500</b>	<b>2800</b>			-5.6	0.8

Ref: Dzuba PRA66, 012111 (2002) - Uncertainties of 50%

\*Based on Woods-Saxon Potential

† Nilsson Potential Prediction is 137 keV

NOTES:

Octupole Enhancements

Engel et al. agree with Flambaum et al.

Even octupole vibrations enhance S (Engel, Flambaum & Zelevinsky)

$^{225}\text{Ra}$ :

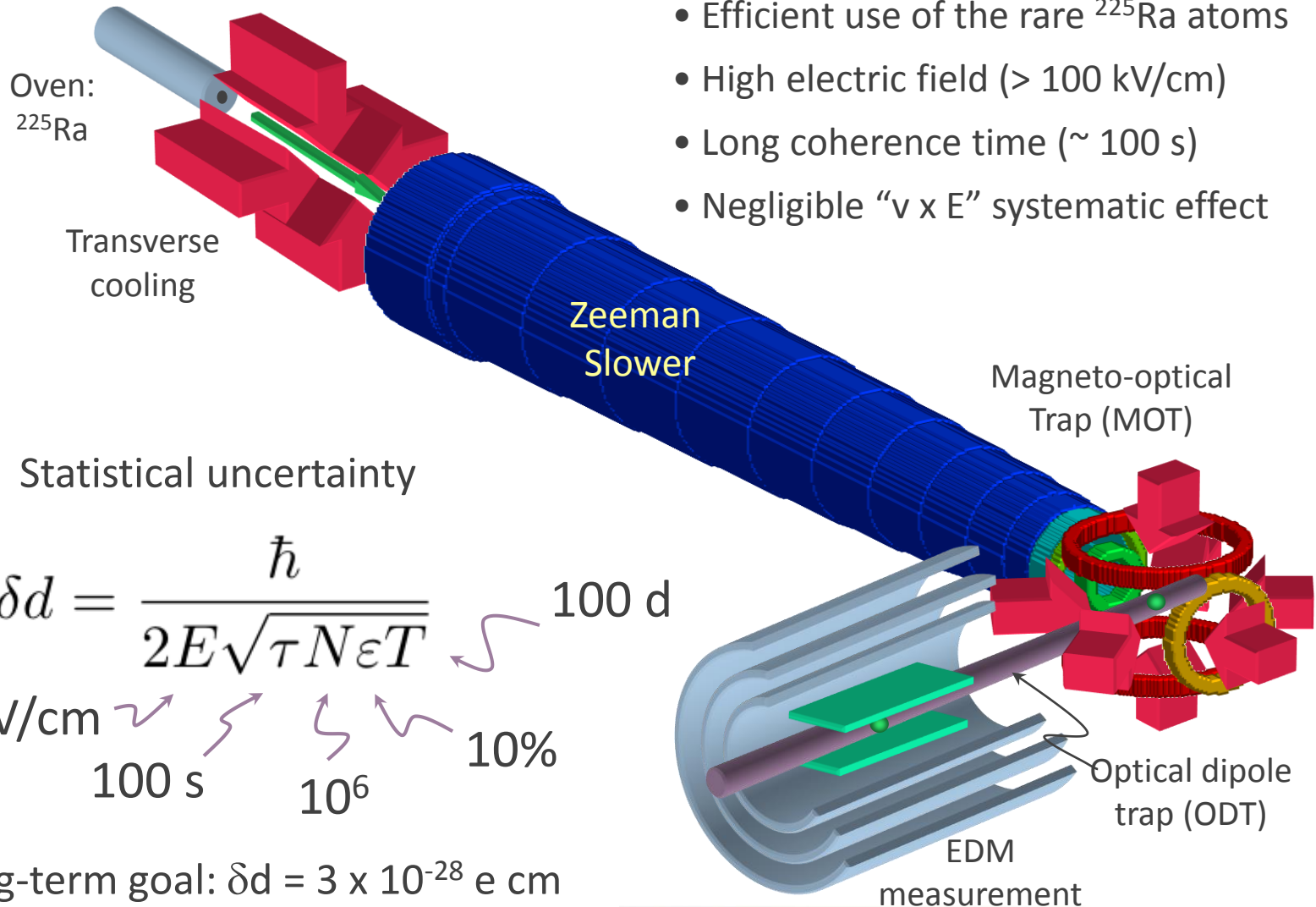
$I = 1/2$

$t_{1/2} = 15 \text{ d}$

# EDM measurement on $^{225}\text{Ra}$ in a trap

Collaboration of Argonne, Kentucky, Michigan State

- Efficient use of the rare  $^{225}\text{Ra}$  atoms
- High electric field ( $> 100 \text{ kV/cm}$ )
- Long coherence time ( $\sim 100 \text{ s}$ )
- Negligible " $\mathbf{v} \times \mathbf{E}$ " systematic effect



Long-term goal:  $\delta d = 3 \times 10^{-28} \text{ e cm}$

# Nuclear EDM outlook

- First results
  - 2015:  $5 \times 10^{-22}$  e cm
  - 2016:  $1.4 \times 10^{-23}$  e cm
- 2015-2018, **blue upgrade** – more efficient trap;
- Five-year goal (before FRIB):  $10^{-26}$  e cm;
- 2020 and beyond (at FRIB):  $3 \times 10^{-28}$  e cm;
- Far future: search for EDM in diatomic molecules
  - Effective E field is enhanced by a factor of  $10^3$ ;
  - Reach the Standard Model value of  $10^{-30}$  e cm.





# Low energy fundamental symmetry tests

- a number of opportunities, exploiting special features of low energy systems and symmetries of the interactions, are being pursued

→ A few presented here ... many more ongoing

- most require pure sources of isotopes with specific intrinsic properties and beam preparation

→ High intensity rare isotopes

- these studies are competitive with and complementary to higher energy experiments
- new technologies enhance our capabilities to pursue these opportunities but to take full advantage of them we need **FRIB**.

