Coulomb form factors of E4 transitions in s–d shell nuclei are discussed taking into account core-polarization effects due to hexadecupole giant resonances. The calculation has been performed within the framework of microscopic theory and gives a remarkably good agreement with experimental form factors both in the absolute strength and the q-dependence.

Recent high-resolution electron scattering data make it possible to obtain model independent transition densities and current densities in many nuclei over a broad region of the mass table [1]. These data provide precise and interesting information which can be used to test the validities of various nuclear models.

A microscopic model has recently been proposed by the present authors [2] in order to study the core-polarization effect on the transition and current densities of single-particle configurations. This model is essentially made from two parts. At the first stage, we calculate the single-particle wave functions and giant resonances by using the self-consistent Hartree–Fock (H–F) + random phase approximation (RPA) theory. Then, we evaluate the core-polarization effect due to these giant resonances by the particle-vibration coupling model. Our model gives quite satisfactory results for describing the E2 core-polarization charges in the s–d shell nuclei in comparison with the empirical ones. Moreover, the Coulomb E2 form factor of the (1d7/2 → 2s1/2) transition in 39K is well reproduced by our calculation in both the q-dependence and the absolute cross section [2].

Higher multipole transitions might also give interesting and important information about the nuclear wave functions since they are more sensitive to the radial profiles of the wave functions than the lower multipole transitions, especially at the surface region [3]. So far, the E4 transition strengths and the Coulomb form factors have been discussed by using s–d shell-model wave functions with phenomenological effective charges [4,5] and also by using the deformed H–F wave functions [6,7]. We will discuss in this letter the core-polarization effect on the Coulomb E4 form factor in the s–d shell nuclei based on a hybrid microscopic theory which combines shell-model wave functions and highly-excited giant resonances by using the perturbation theory.

The reduced one-body matrix element for shell-model wave functions can be expressed as a linear combination of the single-particle matrix elements:

$$U_{\alpha\beta} = \sum_{\alpha,\beta} C_{J_f, J_i} (\alpha, \beta) (\alpha \parallel \hat{T}_\lambda \parallel \beta),$$

where $$J_f$$ and $$J_i$$ stand for the shell-model states and the $$C_{J_f, J_i}(\alpha, \beta)$$ are the structure factors (one-body transition densities). The particle-vibration coupling model [2,8] gives the modified single-particle matrix element,

$$\langle \alpha \parallel \hat{T}_\lambda \parallel \beta \rangle = \langle \alpha \parallel \hat{T}_\lambda \parallel \beta \rangle + \sum_{\omega} [2\omega \chi_(\epsilon^2_{\alpha \beta} - \chi^2_{\omega})] \times \langle \beta \times \omega \chi \alpha \parallel V_{ph} \parallel \beta \rangle \langle \omega \chi \parallel \hat{T}_\lambda \parallel 0 \rangle/(2\lambda + 1)^{1/2},$$

where $$\omega$$ and $$\epsilon_{\alpha \beta}$$ are the excitation energy of giant resonance and single particle energy difference, respectively. The particle-vibration coupling hamiltonian $$V_{ph}$$ is derived from the Skyrme-type interaction by replacing the velocity dependent terms by a Fermi-
momentum dependent $\delta$-interaction [2]. The modified transition matrix element for the shell-model wave function is now given by inserting $\langle \alpha || \tilde{T}_\lambda || \beta \rangle$ in eq. (1);

$$\langle \tilde{T}_{\lambda} || \tilde{T}_\mu \rangle = \sum_{\alpha, \beta} C_{\lambda, \mu} \langle \alpha || \tilde{T}_\lambda || \beta \rangle .$$  \hspace{1cm} (3)

This effect can be regarded as a polarization of the core protons by the valence protons and neutrons through the proton–proton and proton–neutron two-body interaction. The proton and neutron core-polarization charges are defined by,

$$\delta e_p = 1 - \frac{\langle \alpha || \tilde{T}_\lambda || \beta \rangle \pi / \langle \alpha || \tilde{T}_\lambda || \beta \rangle} .$$  \hspace{1cm} (4)

We performed the self-consistent H–F + RPA calculations assuming a $(s^4p^{12}(d_{5/2})^2)$-core using the Skyrme-interaction SGII [9]. This interaction gives the H–F RMS charge radius $(r^2)_{c}^{1/2} (H–F) = 3.107$ fm for $^{28}$Si which is quite close to the experimental value $(r^2)_{c}^{1/2} (\text{exp.}) = 3.125$ fm. The single-particle $B(E4)$-values calculated with harmonic oscillator and the H–F wave functions are listed in table 1. The oscillator length of the harmonic-oscillator wave functions is taken to be $b = 1.819$ fm which provides the RMS charge radius $(r^2)_{c}^{1/2} (\text{HO}) = 3.107$ fm. While the RMS radii of the harmonic-oscillator and H–F ground state wave functions are the same, there is a significant difference (up to 20%) in the $B(E4)$-values (see table 1). The H–F wave functions give $(ld_{3/2}^2 l^4d_{5/2}) (H–F) = 244$ fm$^4$ and $(ld_{5/2}^4 l^4d_{5/2}) (H–F) = 170$ fm$^4$, while the harmonic oscillator wave function gives $(ld^4 l^4d) (\text{HO}) = 172$ fm$^4$.

The RPA responses for the isoscalar (IS) and isovector (IV) hexadecupole operators $T_{IS} = 1/2 \sum_i d^{4}_{Y4m}(r_i)$ and $T_{IV} = 1/2 \sum_i d^{4}_{Y4m}(r_i) r_i$. The solid curve shows the IS response, while the dashed one corresponds to the IV response.

Functions give $(ld_{3/2}^2 l^4d_{3/2}) (H–F) = 244$ fm$^4$ and $(ld_{5/2}^4 l^4d_{5/2}) (H–F) = 170$ fm$^4$, while the harmonic oscillator wave function gives $(ld^4 l^4d) (\text{HO}) = 172$ fm$^4$.

The RPA responses for the isoscalar (IS) and isovector (IV) hexadecupole operators are shown in fig. 1 for $^{28}$Si-core. We can see few strong resonances in the IS response at around $E_x = 20–25$ MeV where 23% of the total strength is existing, while the IV response spreads out in a broad energy region $E_x = 20–60$ MeV without any strong peaks. The IS strength distributions between $E_x = 20–56$ MeV are divided into seven energy regions for the calculation of the core-polarization effects. The transition strength in this energy region exhausts 83% of the energy-weighted sum rule value. The IV response is divided into six energy regions between $E_x = 22–62$ MeV where we found 78% of the IV energy-weighted sum rule value with the enhancement factor $\kappa = 0.09$. Remaining transition strengths are in the high energy tail above $E_x = 60$ MeV. In each energy region, the radial shape of the transition density has about the same shape. The calculated IS and IV core-polarization charges are given in table 1. The averaged $\delta e^{\text{IS}}$ and $\delta e^{\text{IV}}$ are 0.45 and 0.15, respectively. Thus, the

<table>
<thead>
<tr>
<th>$\alpha$</th>
<th>$\beta$</th>
<th>HO $(\times 10^3)$</th>
<th>H–F $(\times 10^3)$</th>
<th>IS</th>
<th>IV</th>
</tr>
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<tr>
<td>ld_{5/2}</td>
<td>ld_{5/2}</td>
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<td>0.133</td>
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<td>0.465</td>
<td>0.153</td>
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<tr>
<td>ld_{3/2}</td>
<td>ld_{3/2}</td>
<td>4.06 5.03</td>
<td>0.412</td>
<td>0.133</td>
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</table>

We performed the self-consistent H–F + RPA calculations assuming a $(s^4p^{12}(d_{5/2})^2)$-core using the Skyrme-interaction SGII [9]. This interaction gives the H–F RMS charge radius $(r^2)_{c}^{1/2} (H–F) = 3.107$ fm for $^{28}$Si which is quite close to the experimental value $(r^2)_{c}^{1/2} (\text{exp.}) = 3.125$ fm. The single-particle $B(E4)$-values calculated with harmonic oscillator and the H–F wave functions are listed in table 1. The oscillator length of the harmonic-oscillator wave functions is taken to be $b = 1.819$ fm which provides the RMS charge radius $(r^2)_{c}^{1/2} (\text{HO}) = 3.107$ fm. While the RMS radii of the harmonic-oscillator and H–F ground state wave functions are the same, there is a significant difference (up to 20%) in the $B(E4)$-values (see table 1). (The H–F wave
Table 2
Excitation energies and B(E4)-values for E4 transitions in s–d shell nuclei. The calculated B(E4)-values are obtained by the s–d shell model wave functions with and without core-polarization effects. The experimental B(E4)-values are taken from ref. [10] (24Mg), ref. [13] (26Mg), ref. [14] (26Al) and ref. [15] (26Si).

<table>
<thead>
<tr>
<th>Nucleus</th>
<th>#</th>
<th>Energy</th>
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<tr>
<td></td>
<td></td>
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<tr>
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<td>4.58</td>
<td>4.51</td>
</tr>
<tr>
<td>26Si</td>
<td>1</td>
<td>4.66</td>
<td>4.62</td>
</tr>
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</table>

Proton and neutron core polarization charges are given by δe_p = δe_1S + δe_1V = 0.60 and δe_n = δe_1S − δe_1V = 0.30. The particle-vibration coupling model with the separable interactions [3,8] gives δe_1S = 0.50 and δe_1V = 0.32 (for a 40Ca-core). The larger IV core-polarization charge stems from the strong IV separable interaction. On the other hand, the E4 polarization charges are about 30% smaller in the first-order perturbation theory calculations using phenomenological gaussian-type interactions [3] or using the bare G-matrices [11]. This might be due to the fact that our particle-vibration coupling model takes into account the higher-order RPA perturbation terms in the calculations [12].

The shell-model calculations have been performed in the full s–d shell model space with the empirical two-body matrix elements of Wildenthal [4]. The excitation energies and B(E4)-values of collective states are listed in table 2. The shell-model wave functions give a quite satisfactory agreement in the excitation

![Fig. 2. Transition densities and Coulomb form factors for the (0° → 4+1) transition in 28Si.](Image)
energies in comparison with experimental data. Nevertheless, the calculated B(E4)-values are typically several times smaller than the empirical ones. We have calculated the B(E4)-values including core-polarization charges. The strong hexadecupole transitions in $^{24}\text{Mg}$, $^{27}\text{Al}$ and $^{28}\text{Si}$ are dominated by the isoscalar part and the B(E4)-values are enhanced by the core-polarization effects by a factor of 3.5. The agreement between experiment and the shell model calculations with core polarization is remarkably good. In $^{26}\text{Mg}$, the ($0^+ \rightarrow 4^+_{1}$) transition has some isovector component and the enhancement factor is relatively small. The fourth $4^+$ state has a large proton (1d$_{5/2}$ $\rightarrow$ 1d$_{3/2}$) amplitude and hence the enhancement is also smaller for B(E4)-value.

We show the transition density and the Coulomb form factor for the ($0^+ \rightarrow 4^+_{1}$) transition in $^{28}\text{Si}$ in fig. 2. The form factor is increased by a factor of 2.5 at the maximum around $q = 1.4$ fm$^{-1}$, however, there is not much increase in the high q-region above

![Coulomb E4 form factors](image)

Fig. 3. Coulomb E4 form factors for the ($0^+ \rightarrow 4^+_{1}$) transition in $^{24}\text{Mg}$, the (5/2$^+ \rightarrow 11/2^+$) transition in $^{27}\text{Al}$ and the ($0^+ \rightarrow 4^+_{1}$) and ($0^+ \rightarrow 4^+_{2}$) transitions in $^{26}\text{Mg}$. The solid and dashed curves correspond to the results with and without core-polarization effects, respectively. The data are taken from: ref. [18] (triangles), ref. [10] (circles) and ref. [10] (squares) for $^{24}\text{Mg}$, ref. [14] for $^{27}\text{Al}$ and ref. [13] for $^{26}\text{Mg}$. 

250
2.0 fm\(^{-1}\). This change is attributed to the enhancement of the transition density at the surface region. In fig. 3, we show the Coulomb form factors for other strong E4 transitions in the vicinity of \(^{28}\)Si. The enhancement factor due to the core-polarization around the peak is almost the same (2.5 times) in every case except for the \((0^+_1 \rightarrow 4^+_1)\) transition in \(^{26}\)Mg. The agreement of the calculation with the experimental data is in general remarkably good. (The experimental data for the \((5/2^+ \rightarrow 11/2^+)\) transition in \(^{27}\)Al was obtained at 90\(^{\circ}\). There is some transverse M3 component, but it is negligible compared to the longitudinal part.) The \((0^+_1 \rightarrow 4^+_1)\) transition in \(^{26}\)Mg has a relatively large isovector component and the enhancement factor around the peak is smaller than for the other cases.

It has been previously noted [4,5] that s-d shell model-space requires an empirical isoscalar effective charge which is slightly larger for the E4 operator, \(\delta e^{1S(E4)} = 0.5\), than for E2, \(\delta e^{1S(E2)} = 0.35\). This feature comes out of our microscopic calculations with the Skyrme interaction SG11 showing the average core polarization charges \(\delta e^{1S(E4)} = 0.43\) and \(\delta e^{1S(E2)} = 0.34\). In ref. [4], the E4 form factors for the lowest states in all stable even–even s-d shell nuclei were calculated based on the assumption of an orbital independent E4 effective charge and a Tassie-type transition density for the core-polarization. The present result based on the microscopic theory justifies these assumptions; our calculated orbital dependence is small (see table 1) and the shape of the transition density for the core-polarization contribution to the low-lying states as well as for the giant resonance is strongly surface peaked and close to the Tassie model shape.

We noted in ref. [2] that a Skyrme-type interaction without momentum-dependent terms gives two times larger core polarization charge for E2 transition than that of the SGII interaction while both interactions give reasonable saturation properties. This is due to the fact that the Skyrme force without momentum dependence gives very low excitation energy for the isoscalar E2 giant resonances. This argument also holds for the E4 case. Thus, we should remark that the momentum-dependent parts of the Skyrme interaction reflecting the finite range part of the two-body force are crucial in order to obtain the proper amount of core polarization charge.

Attempts have been made to explain the E4 transitions in the s–d shell on the basis of deformed H–F calculations [6,7]. It is interesting to compare \(B(E4)\)-values for the transitions to the lowest 4\(^{\pm}\) states in the even–even \(N = Z\) nuclei \((^{26}\)Ne, \(^{24}\)Mg, \(^{28}\)Si, \(^{32}\)S and \(^{36}\)Ar) which are (5.2, 1.8, 0.29, 0.018 and 0.030) (in units of \(10^3 e^2 f m^8\)) for the SI Skyrme interaction and (22.5, 4.6, 16.9, 0.78 and 0.96) for the SII interaction in the deformed H–F calculations. The shell-model calculations with effective charges [4] give (41.9, 0.24, 27.7, 49.9 and 38.2) for these E4 transitions in these units. The experimental values extracted from electron scattering for the first three cases are (38 ± 8, 2.0 ± 0.3, 27 ± 5) [5]. Thus, these empirical data favor our s–d shell model calculation with effective charges.

In the cases where there is a well-defined intrinsic state, we might expect the shell-model and deformed H–F model to give a similar result. However, for cases where the deformed potential energy surface has shallow minima (not necessarily one, but also two or three minima), there is no simple way to calculate the deformed wave function. We might expect the shell model to give a more reliable result for these cases. We note that the two calculations give very different predictions for the E4 transition strength to the lowest 4\(^{+}\) state in \(^{32}\)S and \(^{36}\)Ar. It would be interesting to have electron scattering data for these transitions to study the validity of models.

We would like to thank B.H. Wildenthal for fruitful discussions. This work is supported by National Science Foundation grant no. 83-12245.

References


