



# U.S. Particle Accelerator School

Education in Beam Physics and Accelerator Technology

*Self-Consistent Simulations of Beam and Plasma Systems*

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## A3. Special Topics

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Lawrence Berkeley National Laboratory

# Outline

- Particle pushers
  - Relativistic Boris pusher
  - Lorentz invariant pusher
  - Application to the modeling of electron cloud instability
- Quasistatic method
  - Concept
  - Application to the modeling of electron cloud instability
- Optimal Lorentz boosted frame
  - Concept
  - Application to the modeling of electron cloud instability
  - Generalization
  - Application to the modeling of laser-plasma accelerators



# Relativistic Boris pusher

For the velocity component, the Boris pusher writes

$$u^{n+1} = u^n + \frac{q\Delta t}{m} \left( E^{n+1/2} + \frac{u^{n+1} + u^n}{2\gamma^{n+1/2}} \times B^{n+1/2} \right) \quad \text{with} \quad u = \gamma v$$

which decomposes into

**one acceleration** + **one rotation** + **one acceleration**



$$u^- = u^n + \frac{q\Delta t}{2m} E^{n+1/2} \quad \Rightarrow \quad u^+ - u^- = \frac{q\Delta t}{2m\gamma^{n+1/2}} (u^+ + u^-) \times B^{n+1/2} \quad \Rightarrow \quad u^{n+1} = u^+ + \frac{q\Delta t}{2m} E^{n+1/2}$$

with

$$\gamma^{n+1/2} = \sqrt{1 + \left( u^n + \frac{q\Delta t}{2m} E^{n+1/2} \right)^2 / c^2} = \sqrt{1 + \left( u^{n+1} - \frac{q\Delta t}{2m} E^{n+1/2} \right)^2 / c^2}$$



# Relativistic Boris pusher: problem with $E+v\times B\approx 0$

Assuming  $E$  and  $B$  such that  $E+v\times B=0$ :

$$\rightarrow u^{n+1} = u^n \quad \rightarrow \gamma^{n+1/2} = \gamma^n = \gamma^{n+1}$$

$$\rightarrow \gamma^{n+1/2} = \sqrt{1 + \left( u^n + \frac{q\Delta t}{2m} E^{n+1/2} \right)^2 / c^2} = \sqrt{1 + \left( u^n - \frac{q\Delta t}{2m} E^{n+1/2} \right)^2 / c^2}$$

$$\rightarrow E^{n+1/2} = -E^{n+1/2} = 0 \quad \rightarrow B^{n+1/2} = 0$$

meaning that pusher is consistent with  $(E+v\times B=0)$  only if  $E=B=0$ , and is thus inaccurate for e.g. ultra-relativistic beams.



# Lorentz invariant particle pusher

Replace Boris velocity pusher

- Velocity push: 
$$\mathbf{u}^{n+1} = \mathbf{u}^n + \frac{q\Delta t}{m} \left( \mathbf{E}^{n+1/2} + \frac{\mathbf{u}^{n+1} + \mathbf{u}^n}{2\gamma^{n+1/2}} \times \mathbf{B}^{n+1/2} \right) \quad \mathbf{u} = \gamma \mathbf{v}$$

with

- Velocity push: 
$$\mathbf{u}^{n+1} = \mathbf{u}^n + \frac{q\Delta t}{m} \left( \mathbf{E}^{n+1/2} + \frac{\mathbf{v}^{n+1} + \mathbf{v}^n}{2} \times \mathbf{B}^{n+1/2} \right)$$

Looks implicit but solvable analytically

$$\begin{cases} \gamma^{i+1} = \sqrt{\frac{\sigma + \sqrt{\sigma^2 + 4(\tau^2 + u^{*2})}}{2}} \\ \mathbf{u}^{i+1} = [\mathbf{u}' + (\mathbf{u}' \cdot \mathbf{t})\mathbf{t} + \mathbf{u}' \times \mathbf{t}] / (1 + t^2) \end{cases}$$

with 
$$\begin{cases} \mathbf{u}' = \mathbf{u}^i + \frac{q\Delta t}{m} \left( \mathbf{E}^{i+1/2} + \frac{\mathbf{v}^i}{2} \times \mathbf{B}^{i+1/2} \right) \\ \boldsymbol{\tau} = (q\Delta t / 2m) \mathbf{B}^{i+1/2} \\ u^* = \mathbf{u}' \cdot \boldsymbol{\tau} / c \\ \sigma = \gamma'^2 - \tau^2 \\ \gamma' = \sqrt{1 + u'^2 / c^2} \\ \mathbf{t} = \boldsymbol{\tau} / \gamma^{i+1} \end{cases}$$



# Lorentz invariant particle pusher: test w/ 1 particle

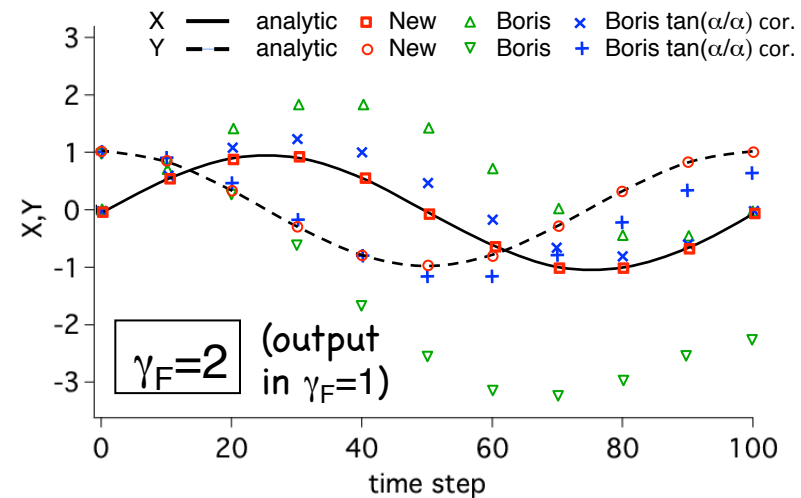
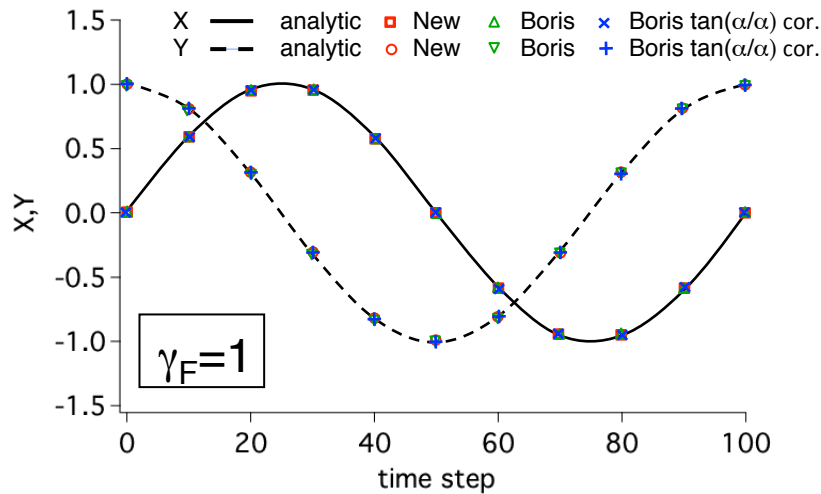
## Lab frame

particle cycling in constant B field



## Boosted frame $\gamma=2$

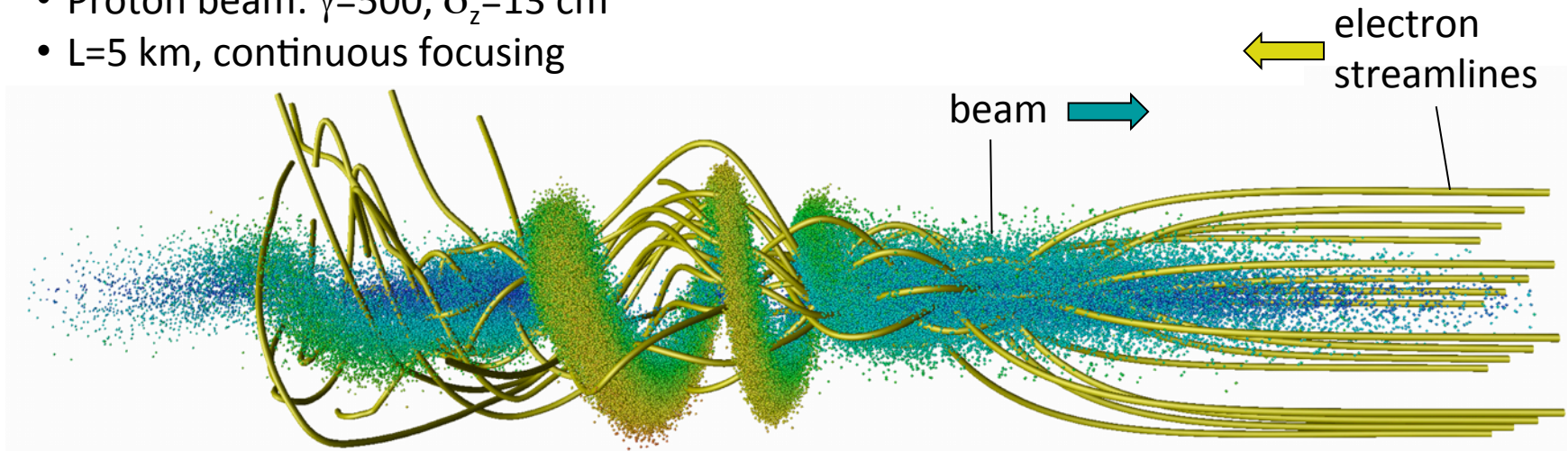
ExB drift adds to gyration



# Application to modeling of two-stream instability

Calculation of e-cloud induced instability of a proton bunch

- Proton beam:  $\gamma=500$ ,  $\sigma_z=13$  cm
- L=5 km, continuous focusing

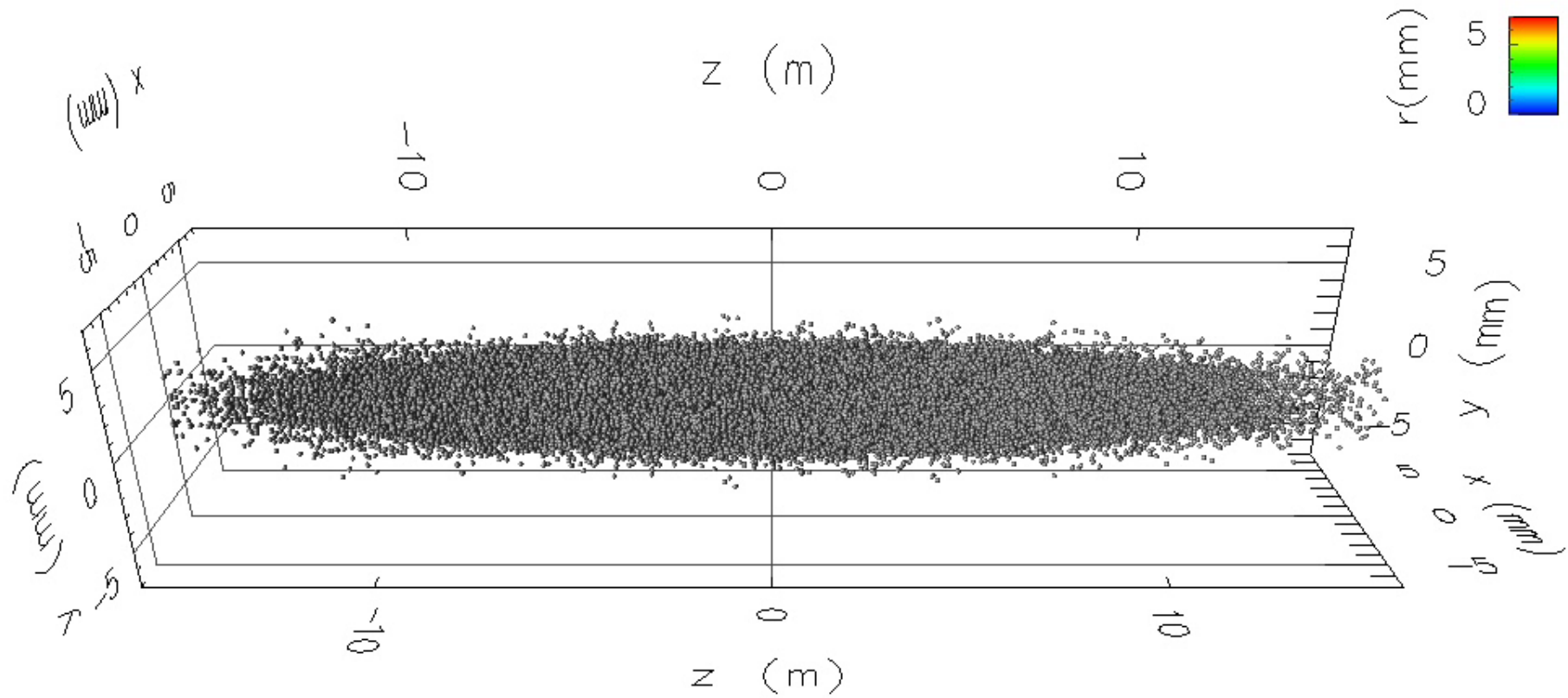


Beam was lost after a few betatron oscillations with Boris pusher.

**Accurate result was obtained with new pusher.**

# Application to modeling of two-stream instability

WARP-3D



Zlab = 0m





# Outline

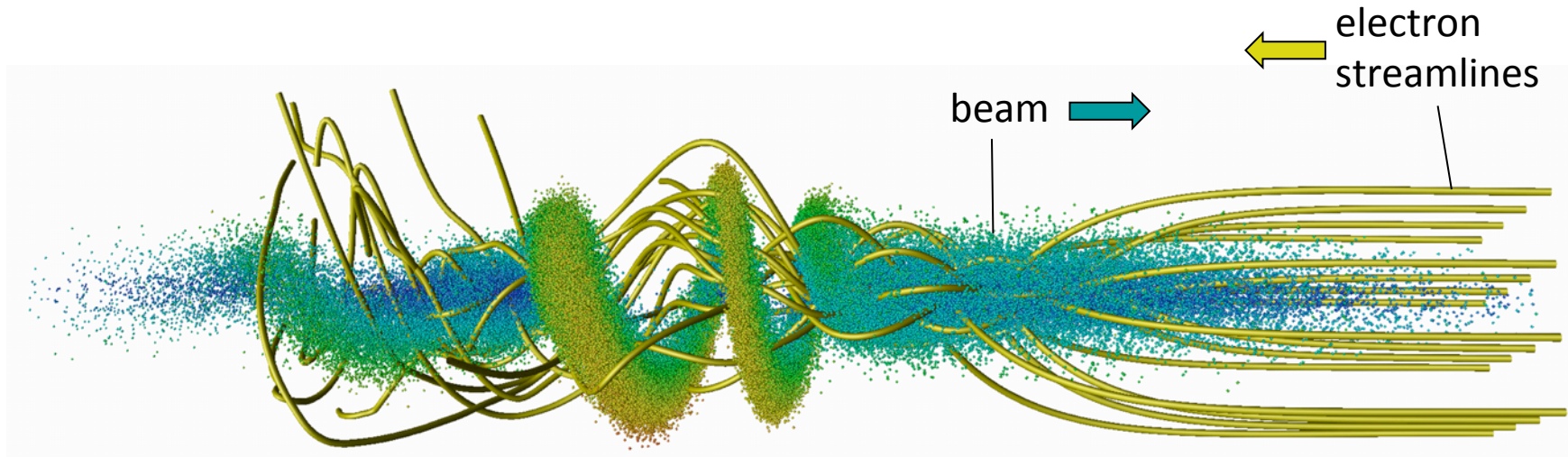
- Particle pushers
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# Modeling of two-stream instability is expensive

Need to follow short ( $\sigma_z=13$  cm) and stiff ( $\gamma=500$ ) proton beam for 5 km:

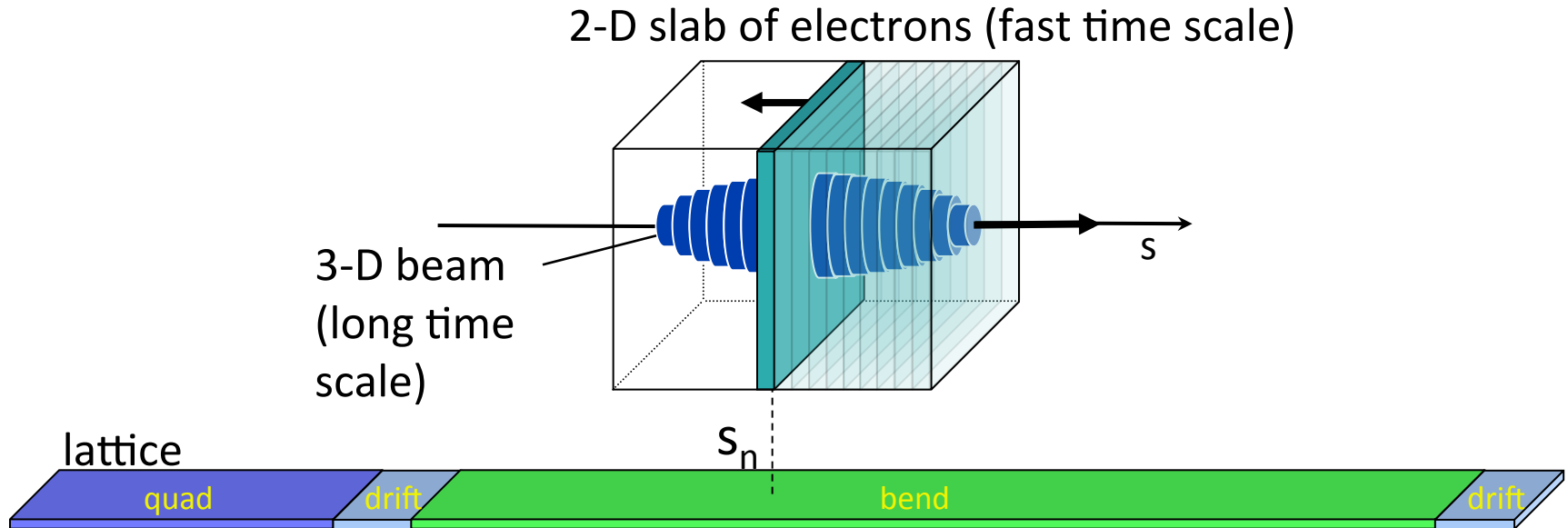
- mobile background electrons react in fraction of beam  $\rightarrow$  small time steps



Two solutions:

- separate treatment of slow (beam) and fast (electrons) components  $\rightarrow$  quasistatic approx.
- solve in a Lorentz boosted frame which matches beam & electrons time scales

# Quasistatic approximation



1. 2-D slab of electrons is stepped backward (with small time steps) through the beam field and its self-field (solving 2-D Poisson at each step),
2. 2-D electron fields are stacked in a 3-D array and added to beam self-field,
3. 3-D field is used to kick the 3-D beam,
4. 3-D beam is pushed to next station with large time steps,
5. Solve Poisson for 3-D beam self-field.

repeat

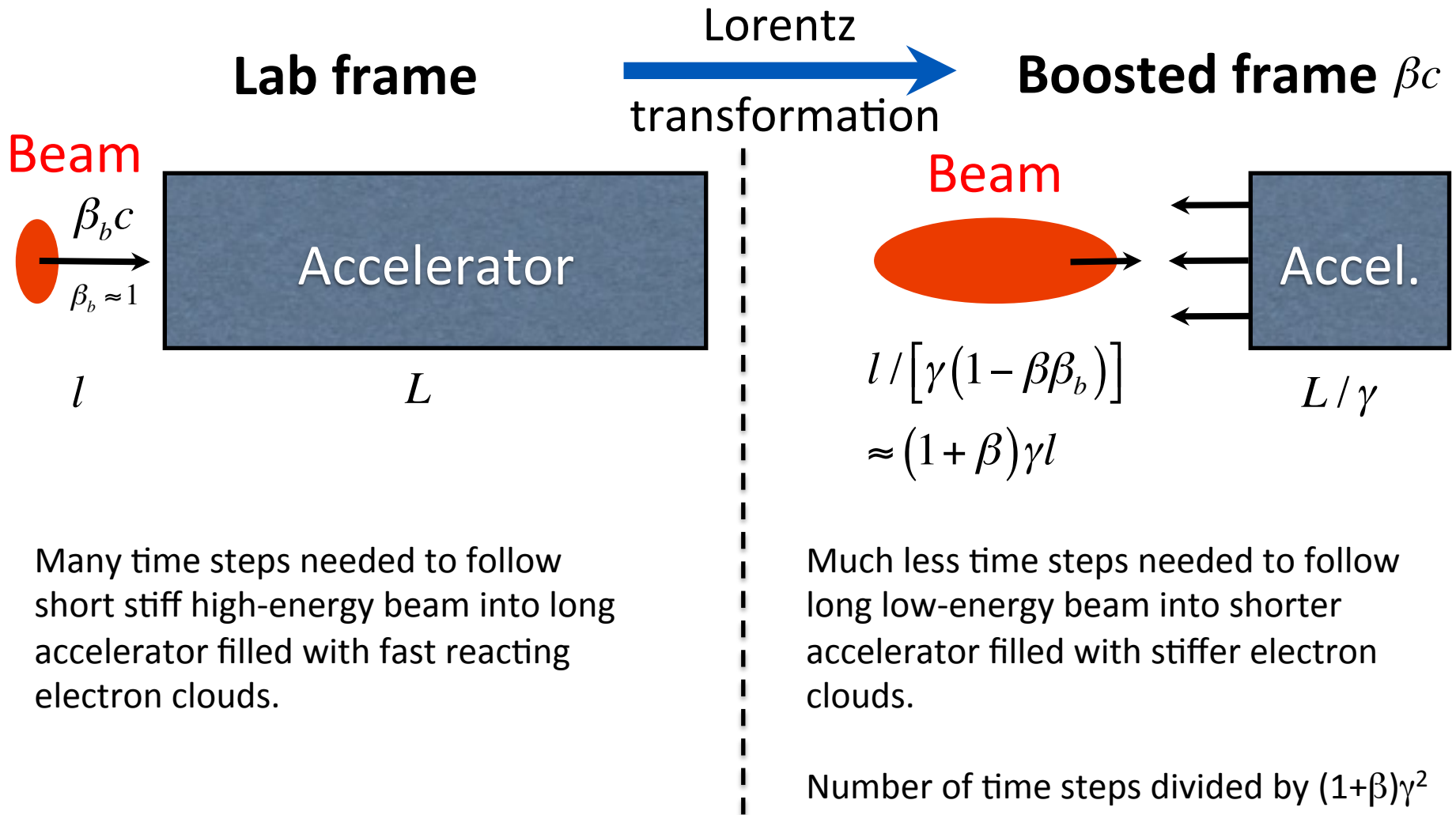


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# Optimal Lorentz boosted frame



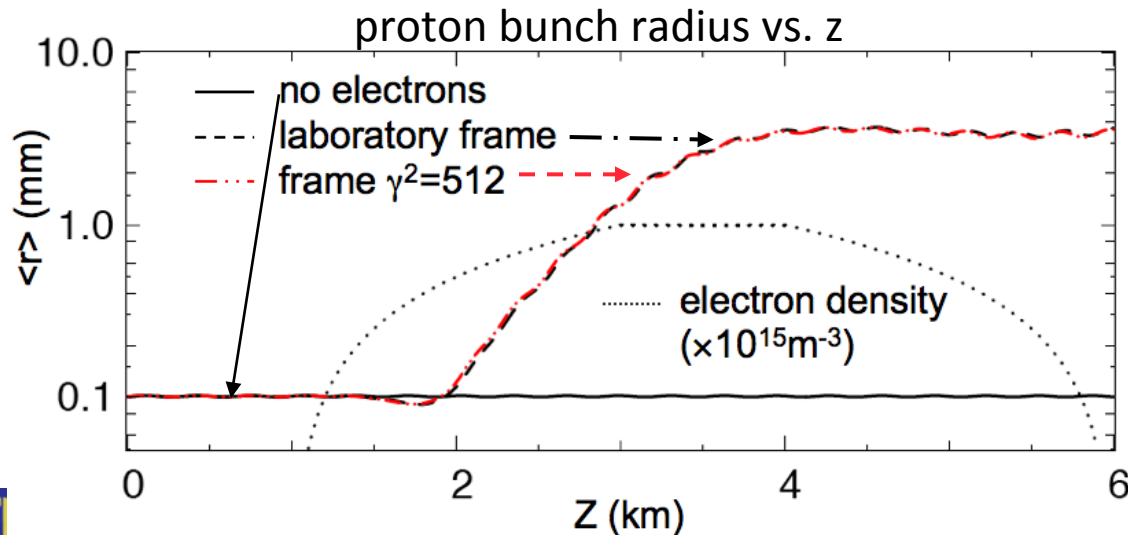
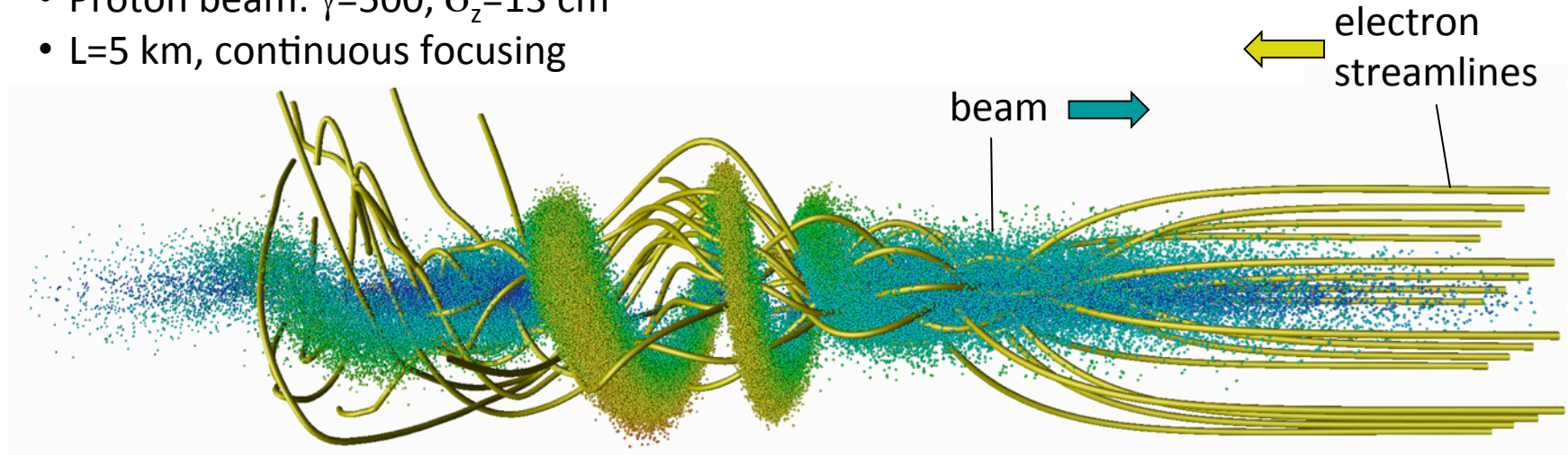
**With high  $\gamma$ , orders of magnitude speedups are possible.**



# Application to modeling of two-stream instability

Calculation of e-cloud induced instability of a proton bunch

- Proton beam:  $\gamma=500$ ,  $\sigma_z=13$  cm
- $L=5$  km, continuous focusing



CPU time (on 8 cores in 2006):

- lab frame: **>2 weeks**
- frame with  $\gamma^2=512$ : **<30 min**

**Speedup x1000**

# Generalization of optimal boosted frame approach

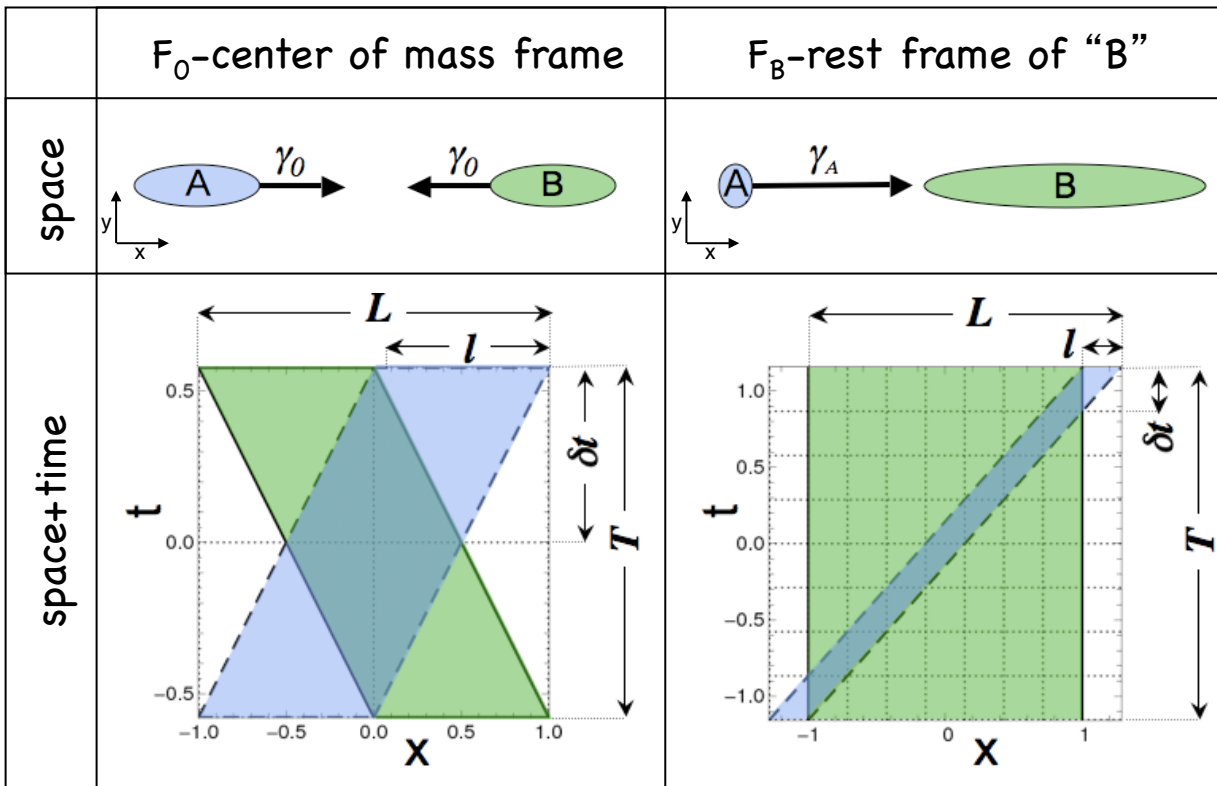
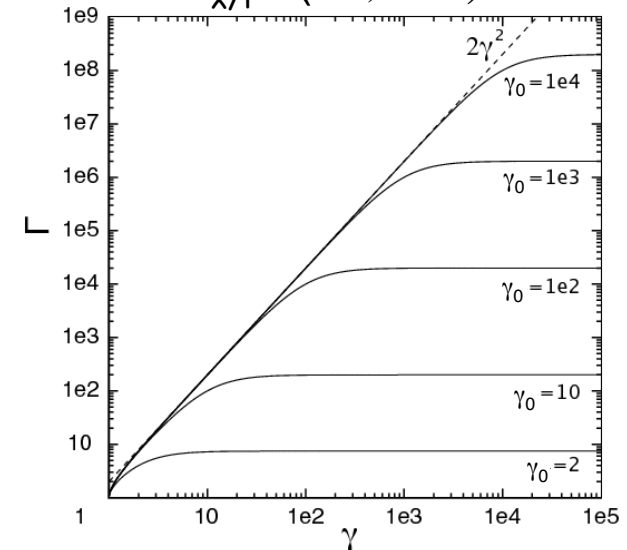
## General formulation:

crossing of 2 relativistic objects

Range of  
space/time scales

$$\Gamma_{x/t} \propto \gamma^2$$

$$\Gamma_{x/t} = (L/l, T/\delta t)^*$$



**The range of space and time scales is not a Lorentz invariant and scales as  $\gamma^2$  for the crossing of two relativistic objects (matter or photons).**

Applicable to study of electron cloud effects, plasma accelerators, free electron lasers, etc.



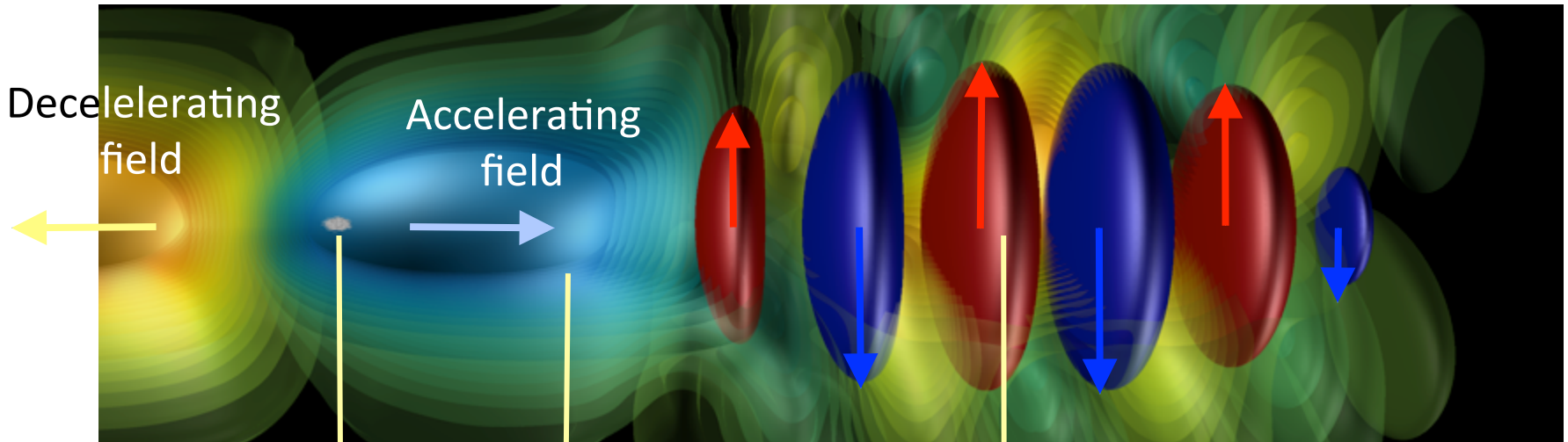
# Laser plasma accelerators “surf” electrons on plasma waves for acceleration on ultra short distances



surfer

wake

boat



e- beam

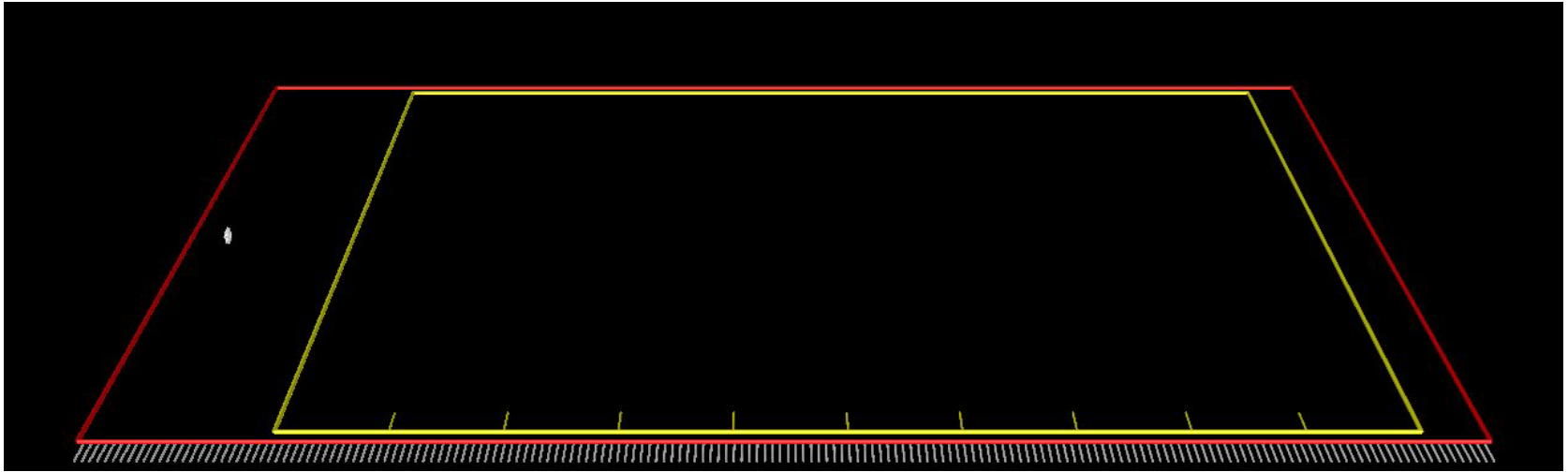
wake

laser





# Modeling from first principle is very challenging



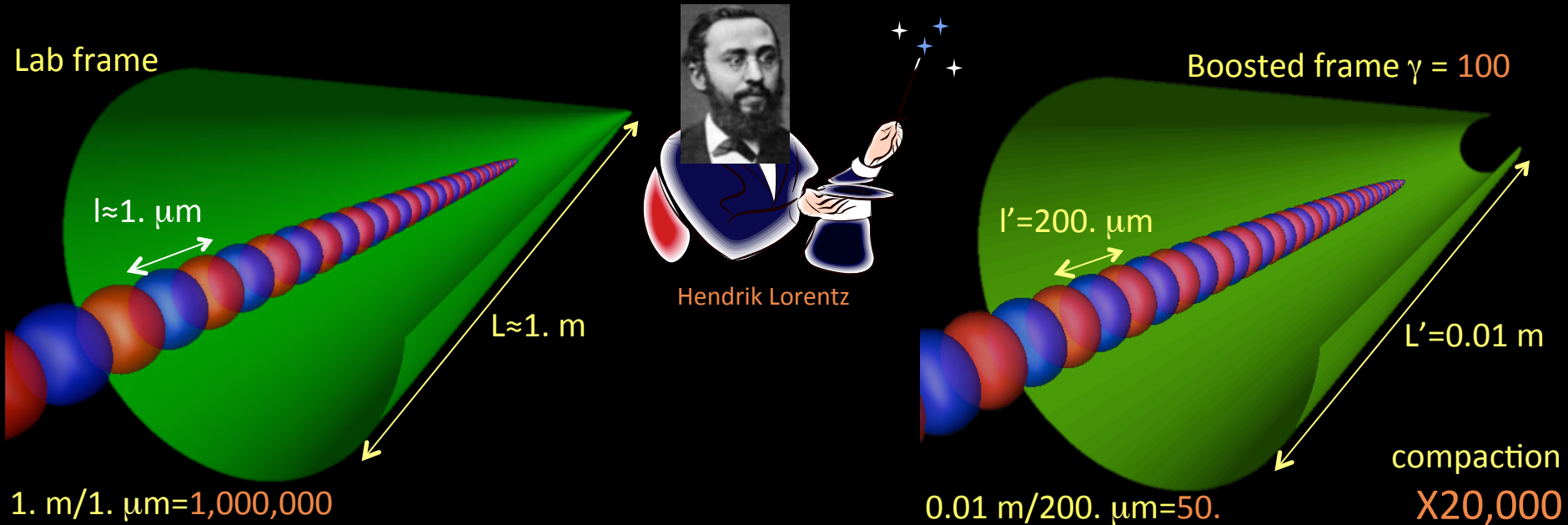
For a 10 GeV scale stage:

~**1 $\mu$ m** wavelength laser propagates into ~**1m** plasma

→ millions of time steps needed

(similar to modeling **5m** boat crossing ~**5000** km Atlantic Ocean)

# Optimal boosted frame enables large speedup

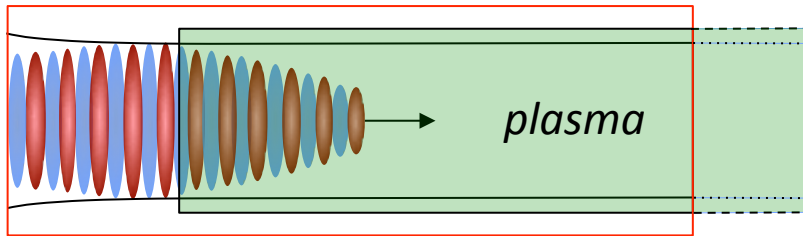


Alternate or complementary solutions: quasistatic, laser envelope, azimuthal Fourier decomposition (“Circ”), ...

# Laser injection through moving plane solves initialization issue in LBF

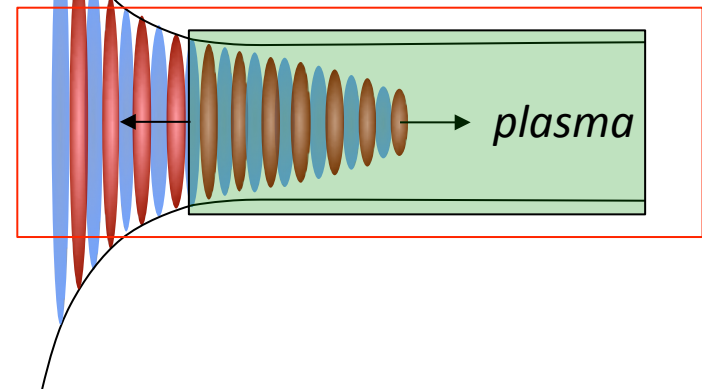
## Lab frame

Standard laser injection from left boundary or all at once

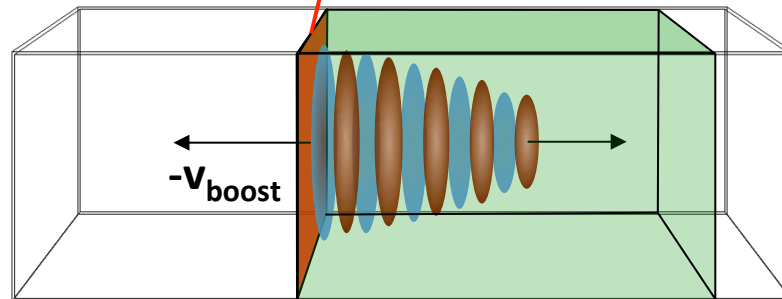


## Boosted frame

Shorter Rayleigh length  $L_R/\gamma_{\text{boost}}$  prevents standard laser injection

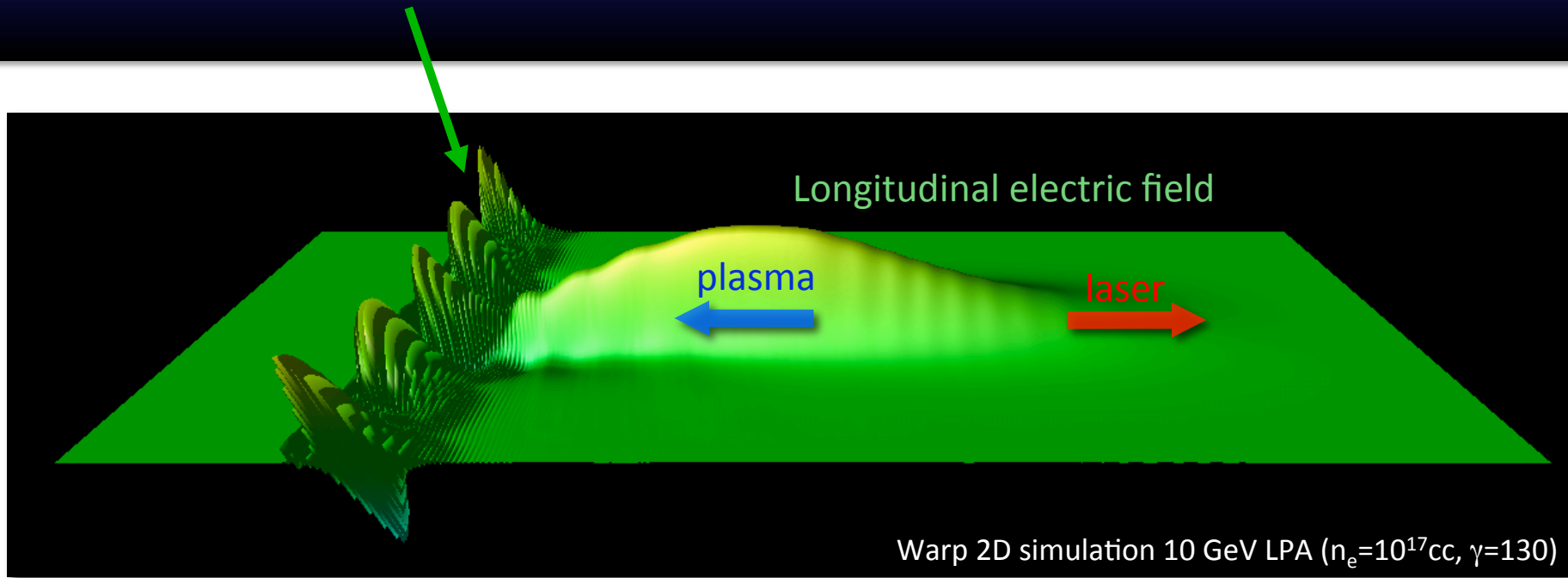


**Solution:** injection through a **moving planar antenna** in front of plasma\*



- Laser injected using macroparticles using Esirkepov current deposition ==> verifies Gauss' Law.
- For high  $\gamma_{\text{boost}}$ , backward radiation is blue shifted and unresolved.

# Short wavelength instability observed at entrance of plasma for large $\gamma$ ( $\geq 100$ )



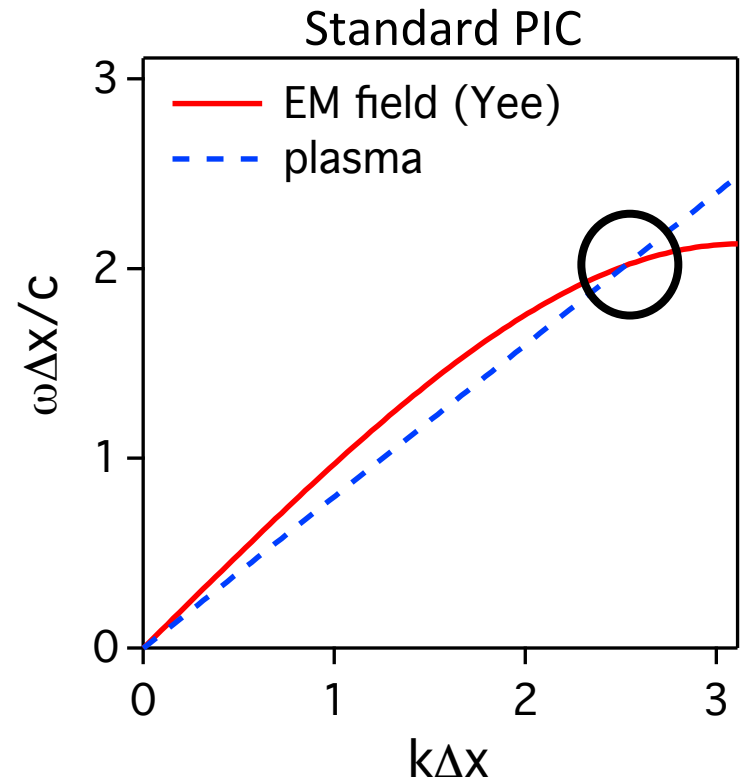
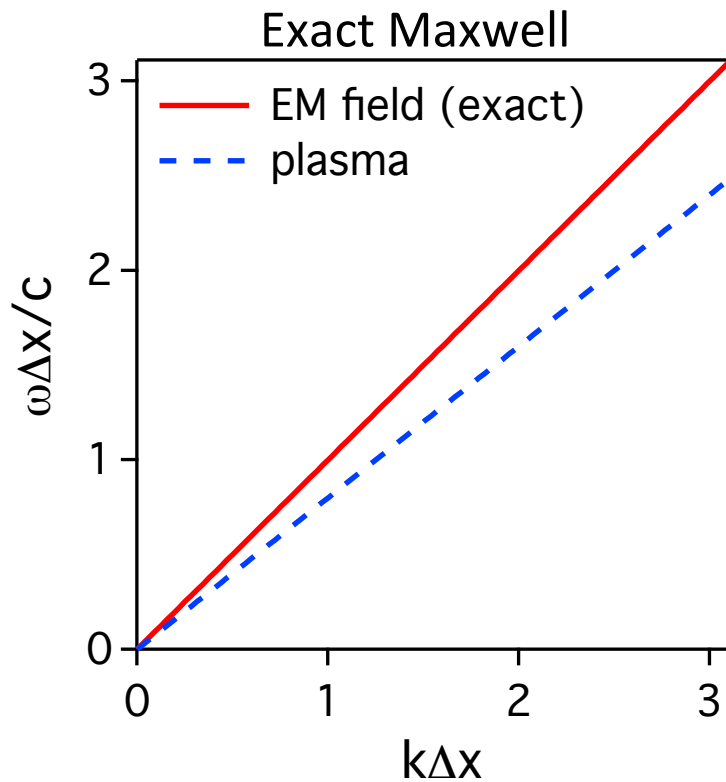
Is it numerical Cherenkov instability?

BTW, what is “numerical Cherenkov instability”?

# Relativistic plasmas PIC subject to “numerical Cherenkov”

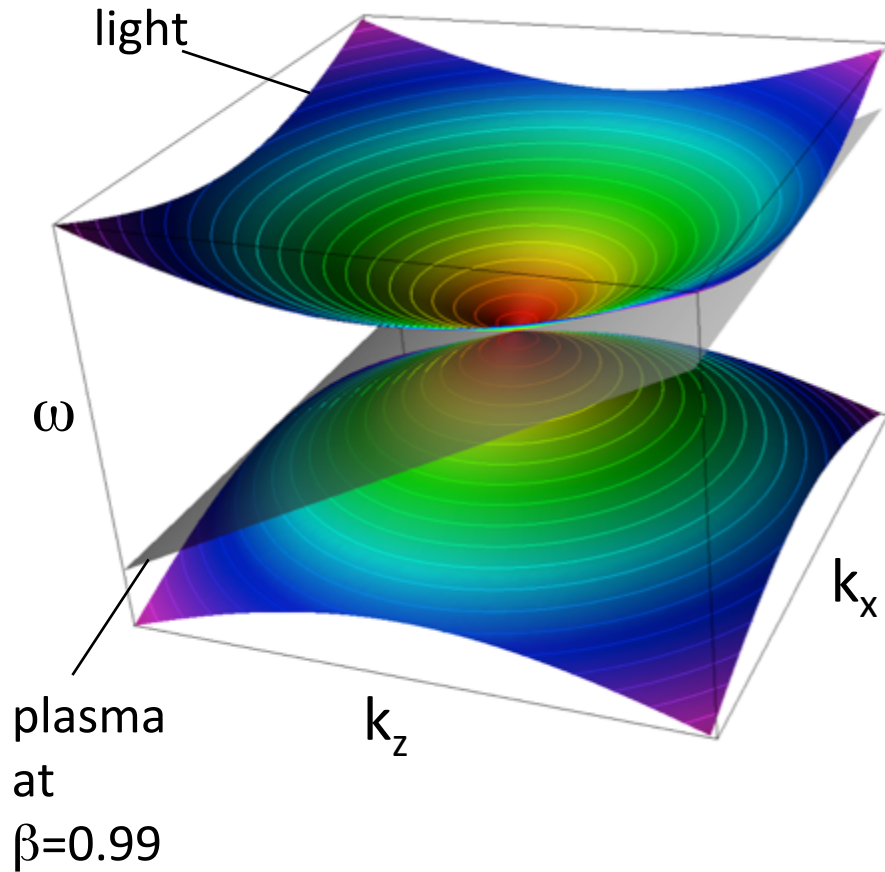
B. B. Godfrey, “Numerical Cherenkov instabilities in electromagnetic particle codes”,  
*J. Comput. Phys.* **15** (1974)

Numerical dispersion leads to crossing of EM field and plasma modes  $\rightarrow$  instability.

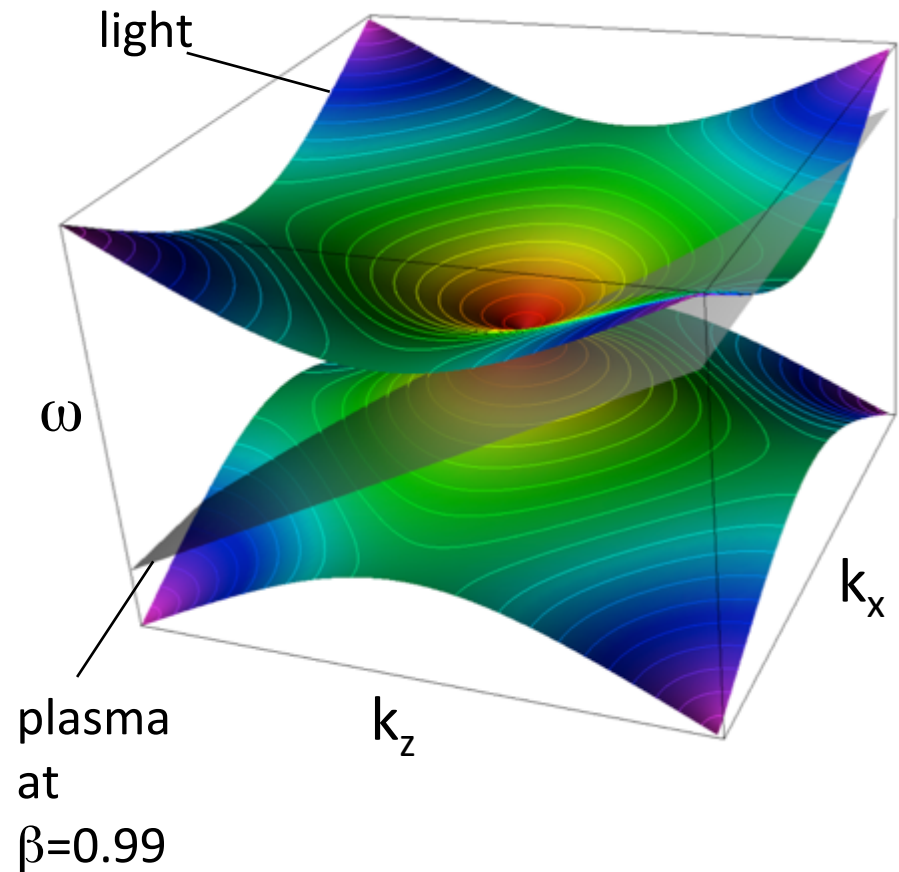


# Space/time discretization aliases $\rightarrow$ more crossings in 2/3-D

Exact Maxwell

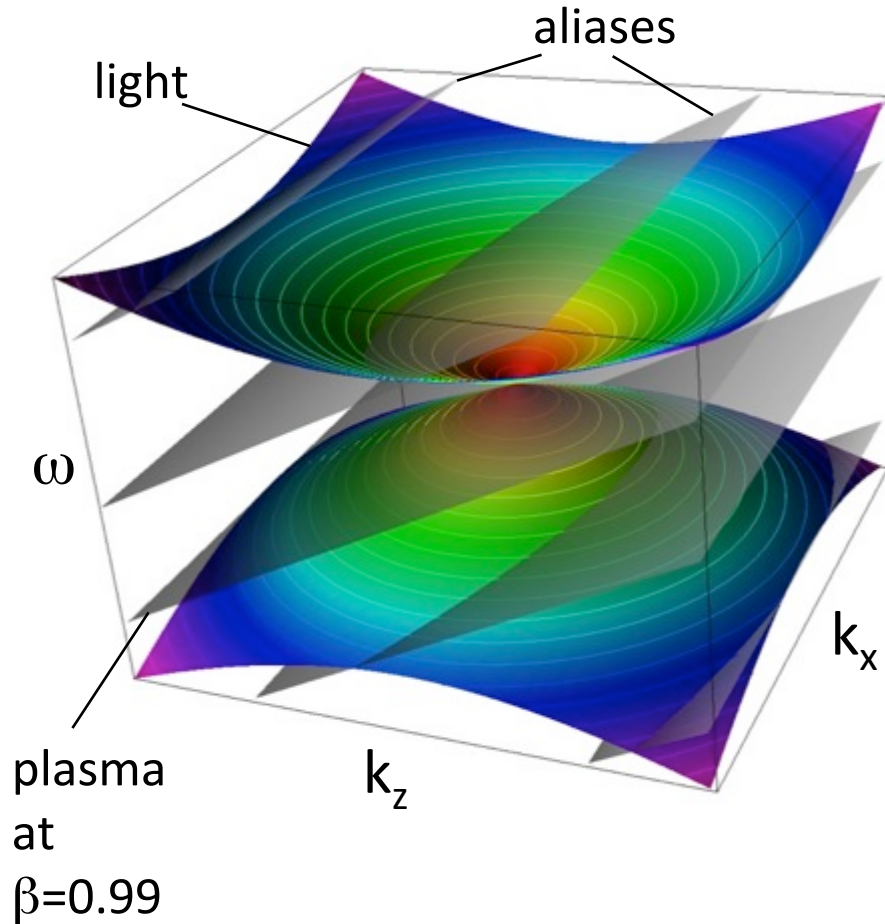


Standard PIC

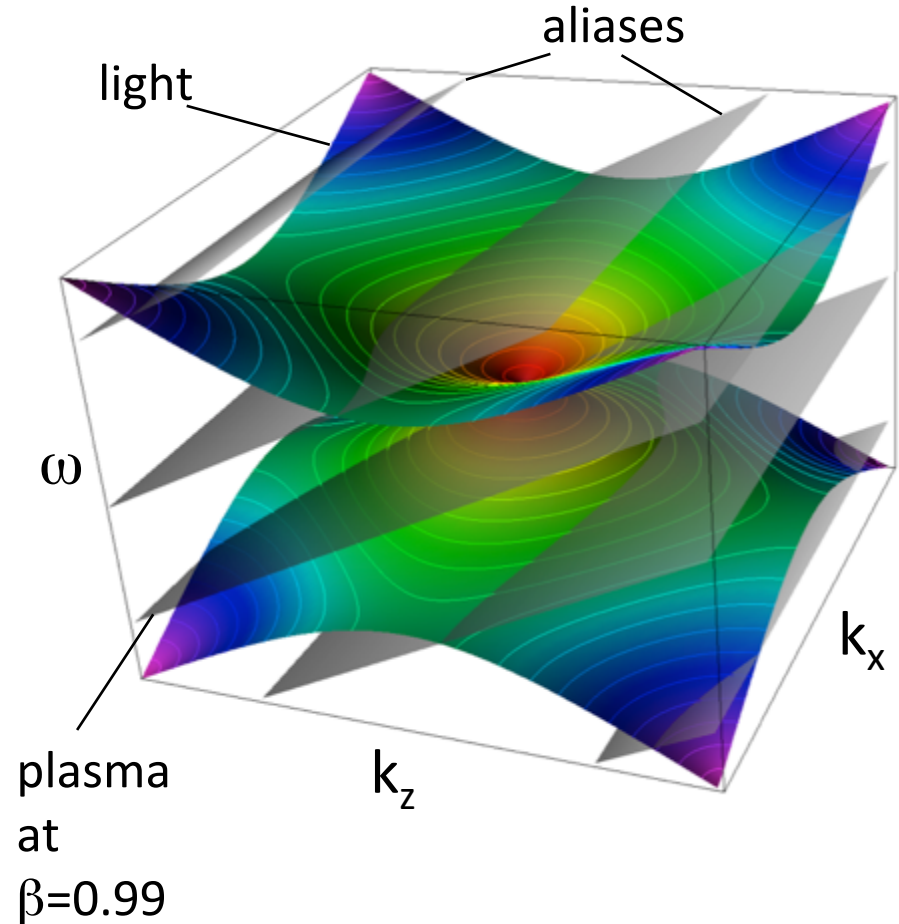


# Space/time discretization aliases $\rightarrow$ more crossings in 2/3-D

Exact Maxwell



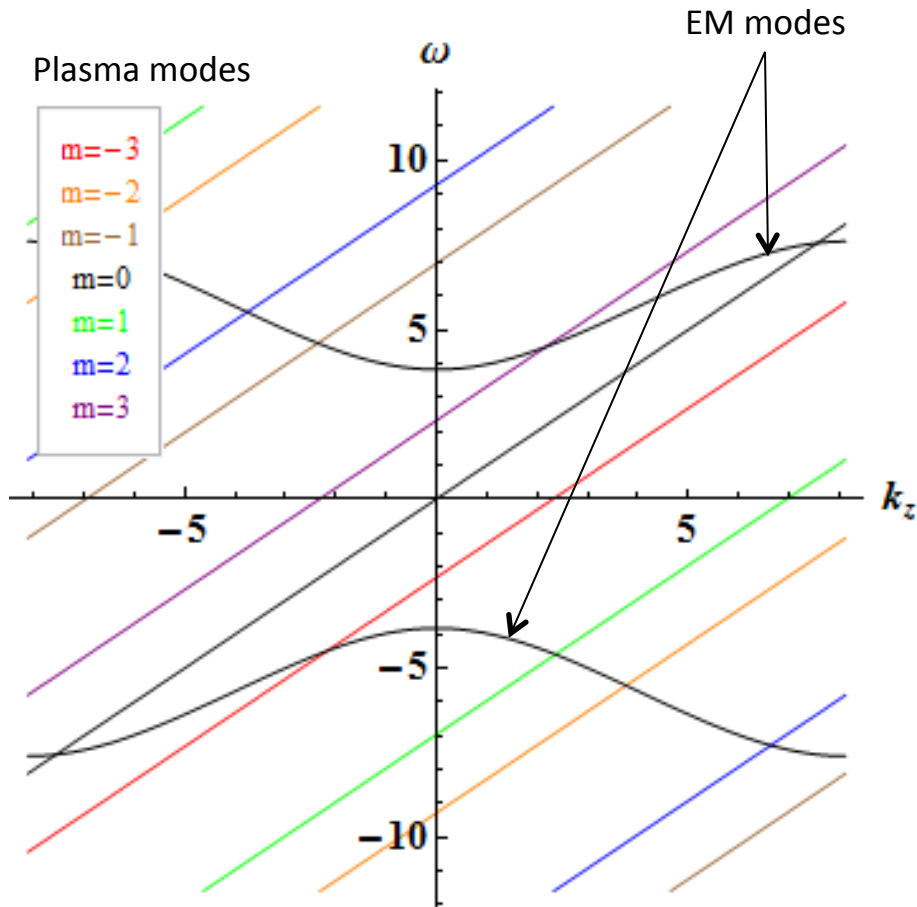
Standard PIC



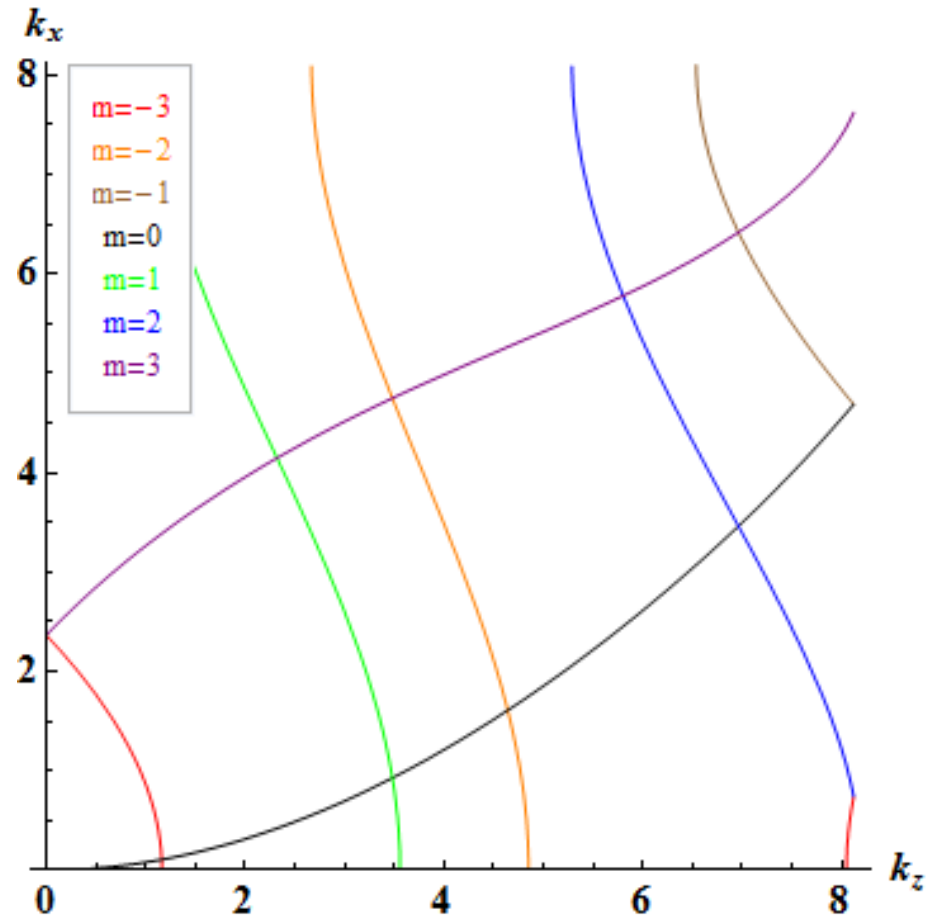
**Analysis calls for full PIC numerical dispersion relation**

# Maps of unstable modes

Normal modes  
at  $k_x=0.5\pi/\Delta x$  for  $c\Delta t=0.7\Delta z$



Projection of normal  
modes intersection





# Numerical dispersion relation of full-PIC algorithm

**2-D relation (Fourier space):**

$$\begin{pmatrix} \xi_{z,z} + [\omega] & \xi_{z,x} & \xi_{z,y} + [k_x] \\ \xi_{x,z} & \xi_{x,x} + [\omega] & \xi_{x,y} - [k_z] \\ [k_x] & -[k_z] & [\omega] \end{pmatrix} \begin{pmatrix} E_z \\ E_x \\ B_y \end{pmatrix} = 0.$$

$$[\omega] = \sin\left(\omega \frac{\Delta t}{2}\right) / \left(\frac{\Delta t}{2}\right) \quad [k_z] = k_z \sin\left(k \frac{\Delta t}{2}\right) / \left(k \frac{\Delta t}{2}\right) \quad [k_x] = k_x \sin\left(k \frac{\Delta t}{2}\right) / \left(k \frac{\Delta t}{2}\right)$$

$$S^J = \left[ \sin\left(k'_z \frac{\Delta Z}{2}\right) / \left(k'_z \frac{\Delta Z}{2}\right) \right]^{\ell_z+1} \left[ \sin\left(k'_x \frac{\Delta X}{2}\right) / \left(k'_x \frac{\Delta X}{2}\right) \right]^{\ell_x+1},$$

$$S^{E_z} = \left[ \sin\left(k'_z \frac{\Delta Z}{2}\right) / \left(k'_z \frac{\Delta Z}{2}\right) \right]^{\ell_z} \left[ \sin\left(k'_x \frac{\Delta X}{2}\right) / \left(k'_x \frac{\Delta X}{2}\right) \right]^{\ell_x+1} (-1)^{m_z},$$

$$S^{E_x} = \left[ \sin\left(k'_z \frac{\Delta Z}{2}\right) / \left(k'_z \frac{\Delta Z}{2}\right) \right]^{\ell_z+1} \left[ \sin\left(k'_x \frac{\Delta X}{2}\right) / \left(k'_x \frac{\Delta X}{2}\right) \right]^{\ell_x} (-1)^{m_x},$$

$$S^{B_y} = \cos\left(\omega \frac{\Delta t}{2}\right) \left[ \sin\left(k'_z \frac{\Delta Z}{2}\right) / \left(k'_z \frac{\Delta Z}{2}\right) \right]^{\ell_z} \left[ \sin\left(k'_x \frac{\Delta X}{2}\right) / \left(k'_x \frac{\Delta X}{2}\right) \right]^{\ell_x} (-1)^{m_z+m_x}.$$

\*B. B. Godfrey, J. L. Vay, I. Haber, J. Comp. Phys. 248 (2013)



# Numerical dispersion relation of full-PIC algorithm (II)

$$\xi_{z,z} = -n\gamma^{-2} \sum_m S^J S^{E_z} \csc^2 \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] \\ (kk'_z \Delta t + \zeta_z k_x^2 \sin(k\Delta t)) \Delta t [\omega] k'_z / 4k^3 k_z,$$

$$\xi_{z,x} = -n \sum_m S^J S^{E_x} \csc \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] \eta_z k'_x / 2k^3 k_z,$$

$$\xi_{z,y} = nv \sum_m S^J S^{B_y} \csc \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] \eta_z k'_x / 2k^3 k_z,$$

$$\xi_{x,z} = -n\gamma^{-2} \sum_m S^J S^{E_z} \csc^2 \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] \\ (k\Delta t - \zeta_z \sin(k\Delta t)) \Delta t [\omega] k_x k'_z / 4k^3,$$

$$\xi_{x,x} = -n \sum_m S^J S^{E_x} \csc \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] \eta_x k'_x / 2k^3 k_x,$$

$$\xi_{x,y} = nv \sum_m S^J S^{B_y} \csc \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] \eta_x k'_x / 2k^3 k_x,$$

\*B. B. Godfrey, J. L. Vay, I. Haber, J. Comp. Phys. 248 (2013)



# Numerical dispersion relation of full-PIC algorithm (III)

$$\eta_z = \cot \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] (k k_z^2 \Delta t + \zeta_z k_x^2 \sin(k \Delta t)) \sin \left( k'_z v \frac{\Delta t}{2} \right) + (k \Delta t - \zeta_x \sin(k \Delta t)) k_z^2 \cos \left( k'_z v \frac{\Delta t}{2} \right),$$

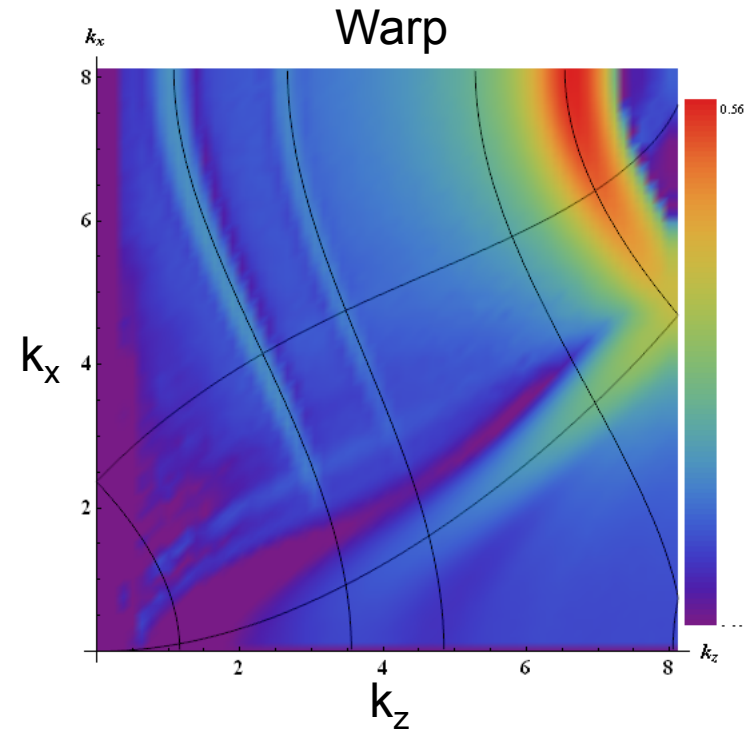
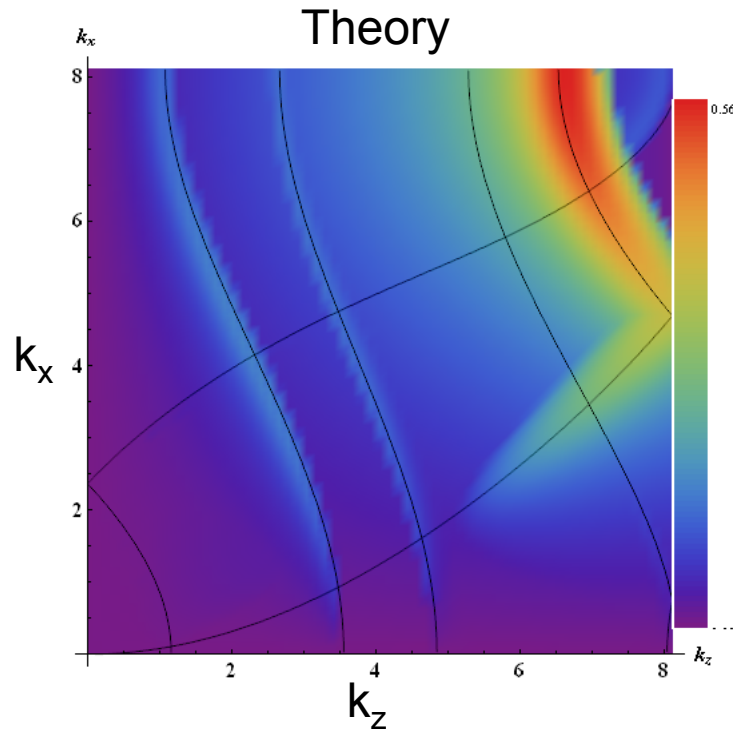
$$\eta_x = \cot \left[ (\omega - k'_z v) \frac{\Delta t}{2} \right] (k \Delta t - \zeta_z \sin(k \Delta t)) k_x^2 \sin \left( k'_z v \frac{\Delta t}{2} \right) + (k k_x^2 \Delta t + \zeta_x k_z^2 \sin(k \Delta t)) \cos \left( k'_z v \frac{\Delta t}{2} \right).$$

Then simplify and solve with Mathematica...

\*B. B. Godfrey, J. L. Vay, I. Haber, J. Comp. Phys. 248 (2013)



# Growth rates from theory match Warp simulations



Warp run uses uniform drifting plasma with periodic BC.  
Yee finite difference, energy conserving gather ( $c\Delta t/\Delta x=0.7$ )

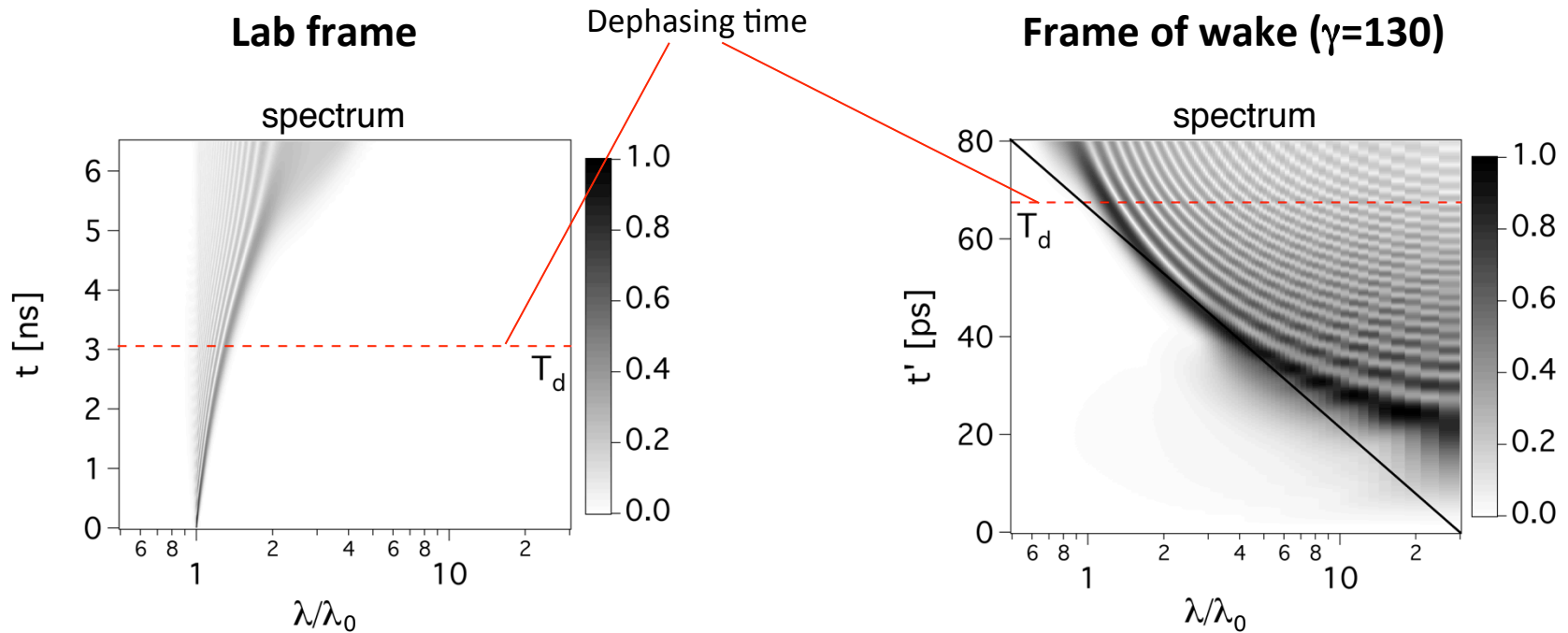
**Latest theory has led to ne insight and the development of very effective methods to mitigate the instability.**



# Physics in boosted frame also allows the use of wideband filtering

Time history of laser spectrum (relative to laser  $\lambda_0$  in vacuum)

Spectrum very different in lab and boosted frames



Content concentrated around  $\lambda_0$

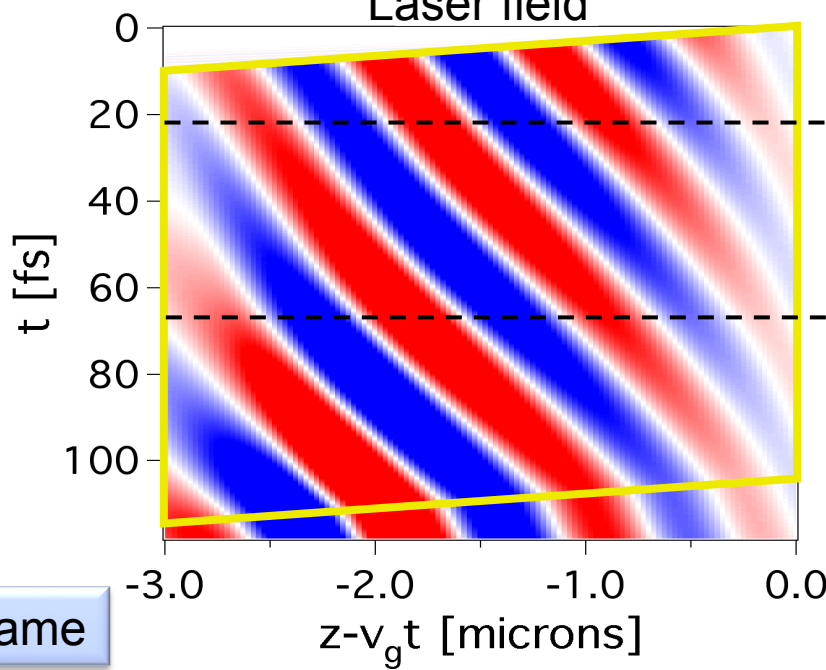
Content concentrated at much larger  $\lambda$

➔ More filtering possible without altering physics\*.

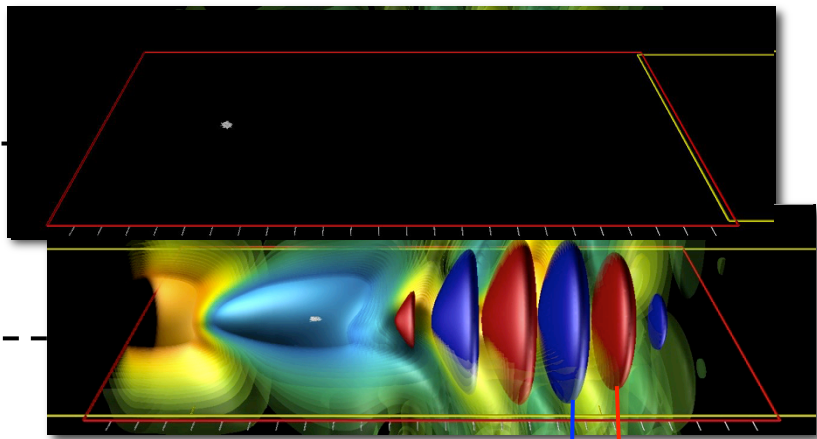


Time  
↓

### Laser field



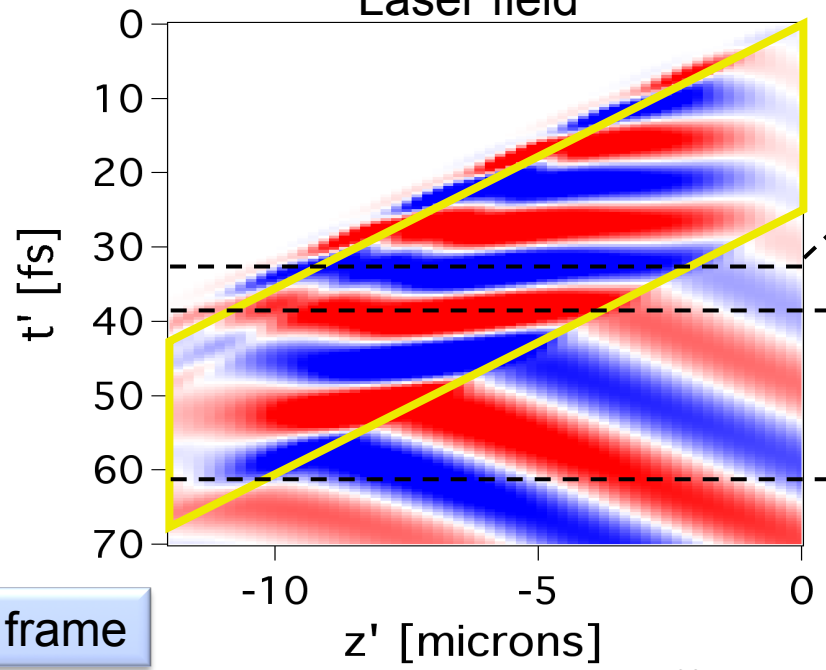
Lab frame



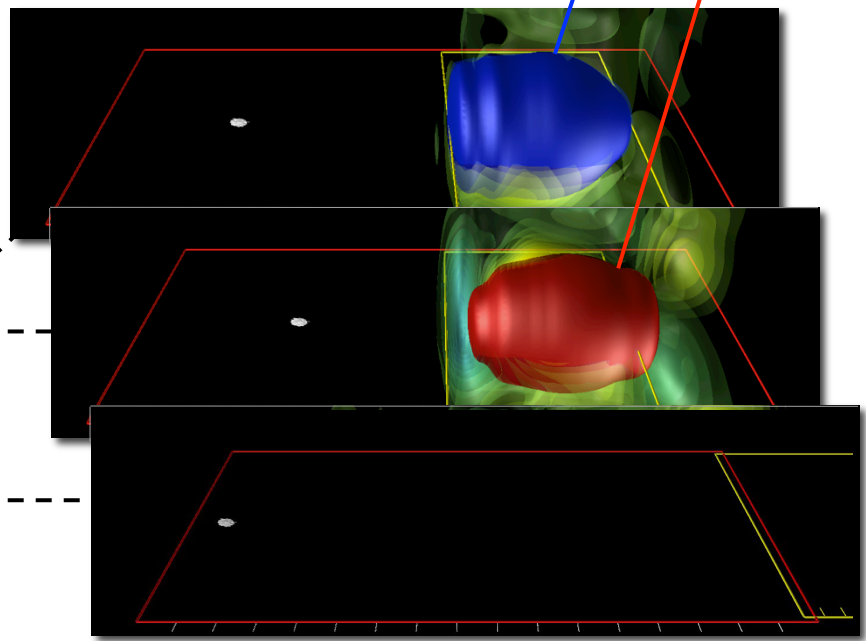
Hyperbolic rotation from Lorentz Transformation converts laser...  
...*spatial oscillations* into *time beating*

Time  
↓

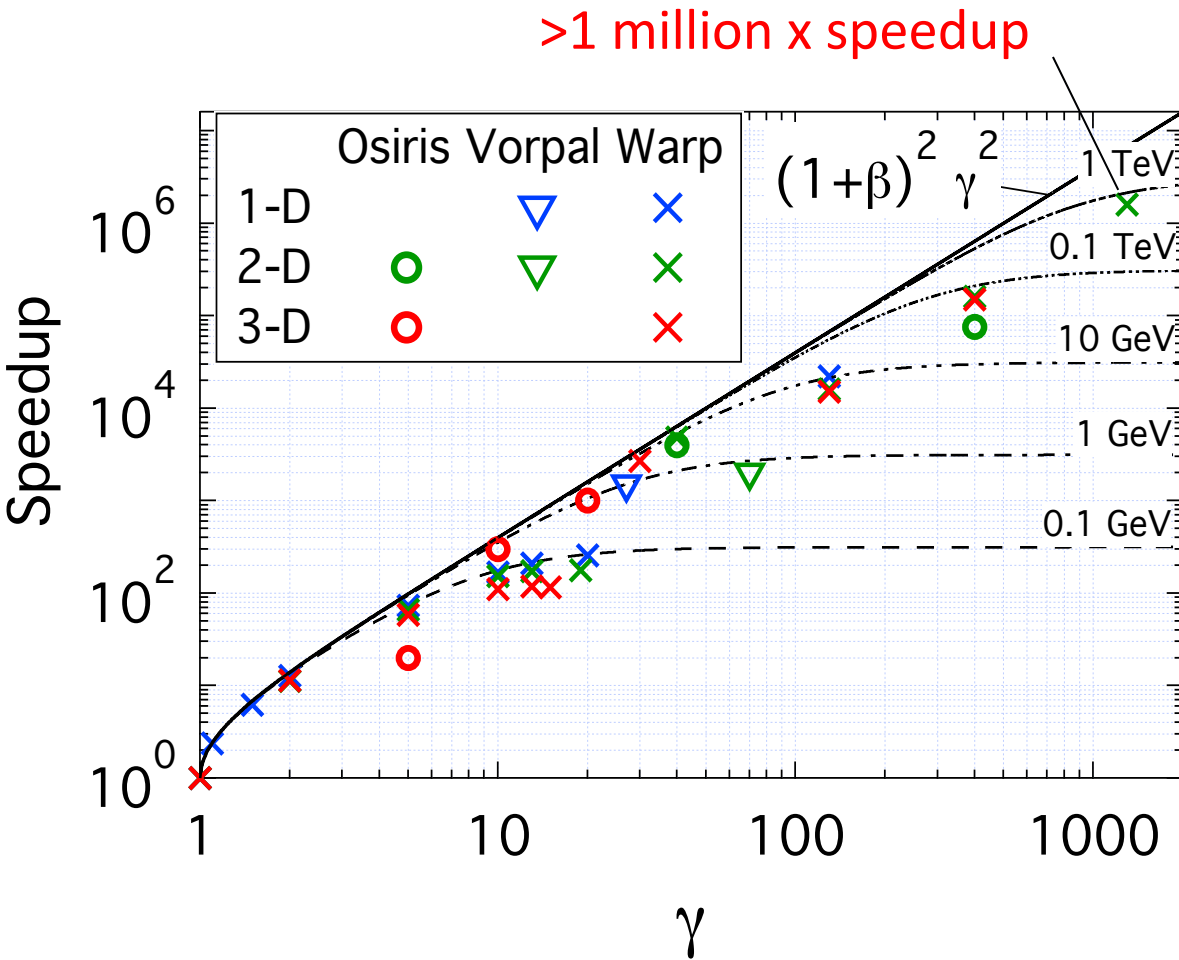
### Laser field



Wake frame



# Speedup verified by us and others to over a million



## Warp:

1. J.-L. Vay, et al., *Phys. Plasmas* **18** 123103 (2011)
2. J.-L. Vay, et al., *Phys. Plasmas (letter)* **18** 030701 (2011)
3. J.-L. Vay, et al., *J. Comput. Phys.* **230** 5908 (2011)
4. J.-L. Vay et al, PAC Proc. (2009)

## Osiris:

1. S. Martins, et al., *Nat. Phys.* **6** 311 (2010)
2. S. Martins, et al., *Comput. Phys. Comm.* **181** 869 (2010)
3. S. Martins, et al., *Phys. Plasmas* **17** 056705 (2010)
4. S. Martins et al, PAC Proc. (2009)

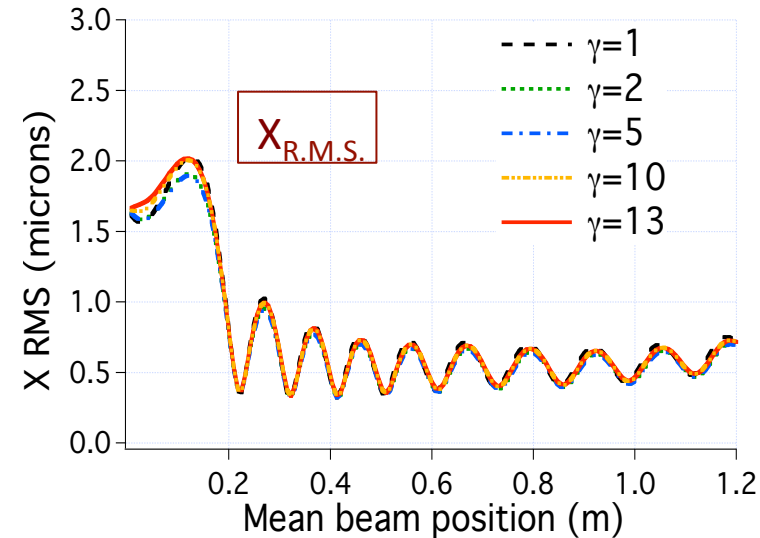
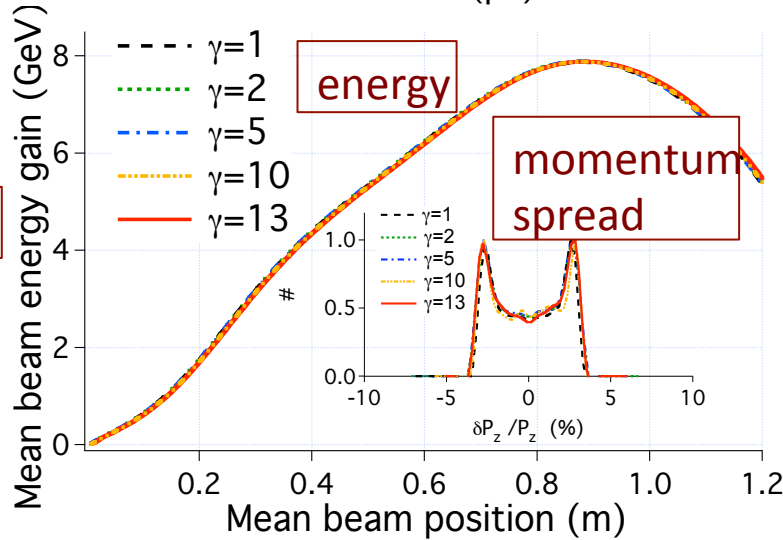
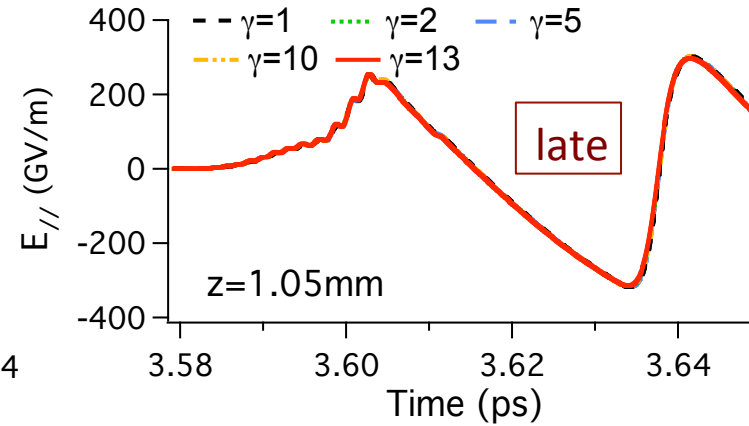
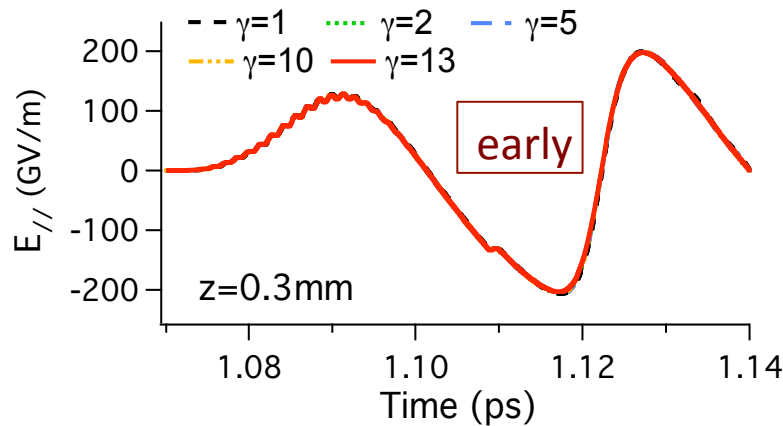
## Vorpal:

1. D. Bruhwiler, et al., *AIP Conf. Proc* **1086** 29 (2009)



# Very high precision validation of BF method with Warp

Simulations in various frames ( $\gamma=1,2,5,10,13$ ) are almost undistinguishable.



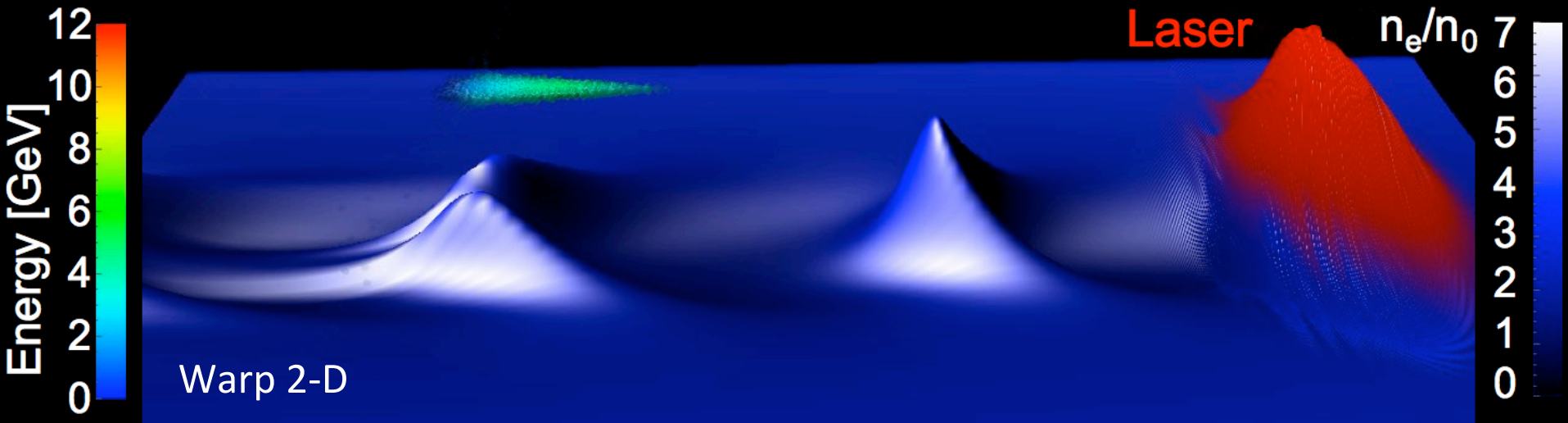
Warp-3D –  $a_0=1$ ,  $n_0=10^{19}\text{cm}^{-3}$  ( $\sim 100$  MeV) scaled to  $10^{17}\text{cm}^{-3}$  ( $\sim 10$  GeV). Detailed validation for  $a_0>1$  (non-linear regime) is underway.





# Enabling simulations that were previously untractable

Simulation of 10 GeV stage for BELLA project (LBNL)



State-of-the-art PIC simulations of 10 GeV stages:

2006 (lab) in 1D:  $\sim 5k$  CPU-hours  $\rightarrow$  2011 (boost) in 3D:  $\sim 1k$  CPU-hours

Current state-of-the-art in lab: 2-D RZ simulations in  $\sim 2$  weeks on thousands of cores.

# Special topics summary

- Modeling of relativistic beams/plasmas with full PIC may benefit from “non-standard” algorithms
  - Lorentz invariant particle pusher
  - Quasistatic approximation
  - Optimal Lorentz boosted frame
- Quasistatic is well established method, but requires writing dedicated code or module
- Boosted frame approach is newer and uses standard PIC at core, needing only extensions



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