Neutron-unbound states in ³¹Ne

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Background: The island of inversion near the N = 20 shell gap is home to nuclei with reordered single-particle energy levels compared with the spherical shell model. Studies of ³¹Ne have revealed that its ground state has a halo component characterized by a valence neutron orbiting a deformed ³⁰Ne core. This lightly bound nucleus with a separation energy of only $S_n = 170$ keV is expected to have excited states that are neutron unbound.

Purpose: The purpose of this experiment was to investigate the low-lying excited states in ³¹Ne that decay by the emission of a single neutron.

Methods: An 89 MeV/nucleon ³³Mg beam impinged on a segmented Be reaction target. Neutron-unbound states in ³¹Ne were populated via a two-proton knockout reaction. The ³⁰Ne fragment and associated neutron from the decay of ³¹Ne were detected by the MoNA-LISA-Sweeper experimental setup at the National Superconducting Cyclotron Laboratory. Invariant-mass spectroscopy was used to reconstruct the two-body decay energy (³⁰Ne +*n*).

Results: The two-body decay energy spectrum exhibits two features: a low-lying peak at 0.30 ± 0.17 MeV and a broad enhancement at 1.50 ± 0.33 MeV, each fit with an energy-dependent asymmetric Breit-Wigner lineshape representing a resonance in the continuum. Accompanying shell-model calculations using the FSU interaction within NUSHELLX, combined with cross-section calculations using the eikonal reaction theory, indicate that these peaks in the decay energy spectrum are caused by multiple resonant states in ³¹Ne.

Conclusions: Excited states in ³¹Ne were observed for the first time. Transitions from calculated shell-model final states in ³¹Ne to bound states in ³⁰Ne are in good agreement with the measured decay energy spectrum. Cross-section calculations for the two-proton knockout populating ³¹Ne states as well as spectroscopic factors pertaining to the decay of ³¹Ne into ³⁰Ne are used to examine the results within the context of the shell-model expectations.

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I. INTRODUCTION

Advances in radioactive beam facilities allow access to regions of the nuclear chart far from the valley of stability. Studies of exotic nuclei near the neutron drip line have led to the discovery of interesting structural phenomena such as the islands of inversion [1], in which traditional shell-model gaps are diminished or shifted, and halo structures [2], described by one or more loosely bound nucleons around a deformed core, with extended nuclear radii.

One signature of a halo nucleus is the large interaction cross section, which implies an extended-matter radius such as that observed in ¹¹Li [3]. Due to the centrifugal barrier, neutron-halo nuclei require *s*- or *p*-wave valence neutrons in order to account for the large observed interaction cross sections. For N > 20 nuclei, the spherical shell model

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suggests the valence neutron(s) would occupy the $f_{7/2}$ orbital; this higher-angular-momentum orbital is not characterized by an extended tail, implying that a halo structure in this region is indicative of level inversion with respect to the spherical shell model.

Halo structures are more common among weakly bound neutron-rich nuclei with a handful of halo nuclei identified, ³⁷Mg being the heaviest observed [4]. According to large-basis shell-model calculations, ³¹Ne, as well as its nearest neighbors ³⁰Ne, ³²Ne, and ³²Na, all belong to the N = 20 island of inversion [1].

Particle stability of ³¹Ne was established in 1996 [5] through a projectile fragmentation study. Neutron removal cross sections were measured at the RIBF (RIKEN) by Nakamura *et al.* [6]. In this study, the authors presented reaction and large-basis shell-model calculations and identified the ground-state spin-parity to be $\frac{3}{2}^{-}$, with the halo component formed by a weakly bound *p*-wave neutron having about 30% of a single-nucleon strength. The significant contribution of a weakly bound low-*l* orbital (*l* = 0 or 1) indicates halo structure for the ³¹Ne ground state, and by association indicates the level inversion from the independent-particle shell-model ordering.

These observed changes in nuclear structure near the drip line have provided important phenomena for testing theoretical models. Studies of excited states of loosely bound nuclei can help to provide a better understanding of the changing shell structure in exotic nuclei near the neutron drip line. Experimental information on the structure of ³¹Ne is limited thus far to studies of the ground state, which is very weakly bound with a low evaluated one-neutron separation energy $S_n = 170 \pm 130$ keV [7]. It is predicted that the excited states of ³¹Ne are neutron unbound [8]. The present study investigates the excited states of ³¹Ne for the first time via the method of invariant-mass spectroscopy.

II. EXPERIMENTAL SETUP

The experiment to investigate ³¹Ne excited states took place at the National Superconducting Cyclotron Laboratory (NSCL) at Michigan State University where a 140 MeV/nucleon ⁴⁸Ca beam was accelerated by the K500 and K1200 coupled cyclotrons. The beam was incident on an 849 mg/cm² Be target to produce ³³Mg nuclei via projectile fragmentation. A 736 mg/cm² aluminum wedge was used in the A1900 fragment separator to facilitate isotopic separation of the isotope of interest, which was transported at a magnetic rigidity of 3.8 Tm to the experimental setup.

³³Mg entered the experimental hall that housed the MoNA-LISA neutron detector array [9,10] and the large-gap superconducting sweeper dipole magnet [11] with an energy of 88.8 ± 0.17 MeV/nucleon (see Fig. 1). The beam first passed through a 420 μ m plastic timing scintillator before impinging on the segmented reaction target [12]. The latter was configured of an array of four 140 μ m silicon detectors and three interleaved beryllium targets with thicknesses of 199 mg/cm² each.

The ³¹Ne nuclei were produced from a ${}^{33}Mg(-2p)$ knockout reaction in one of the three Be targets. The ³¹Ne nuclei



FIG. 1. Schematic layout of the ${}^{33}Mg(-2p)$ experiment. See text for details.

produced may have populated the bound ground state or neutron-unbound excited states that decay immediately into a ³⁰Ne fragment and a neutron. The ³⁰Ne fragment was centered on the face of the first of two cathode readout drift chambers (CRDCs) after traveling through the large-gap (14 cm) "sweeper" superconducting dipole magnet [11]. The neutron associated with the decay was detected in one of the $10 \text{ cm} \times 10 \text{ cm} \times 200 \text{ cm}$ bars from the MoNA-LISA plastic scintillator array [9,10]. The position and energy loss of the ³⁰Ne charged fragment were measured by using the two CRDCs and an ionization chamber, respectively. The fragment time of flight was measured by using a plastic scintillator at the target position and another at the end of the charged particle detector setup. The position and time of flight of the forward-moving neutrons were measured in coincidence within an accuracy of \approx 7 cm and 100 ps, respectively.

III. ANALYSIS

The time of flight of all the fragments (e.g., secondary beams) transmitted to the experimental hall were measured from the A1900 fragment separator focal plane to the target position. This information can be used in combination with the energy loss measured in the first Si detector of the segmented target to create a particle identification plot for the incoming beams. The ³³Mg isotope is selected by placing appropriate gates on these parameters.

The charged ³⁰Ne fragments were detected and subsequently identified by using a suite of charged-particle detectors following the sweeper magnet. An element identification can be made based on the energy loss in the ionization chamber and the time of flight between two plastic timing detectors located just upstream from the reaction target and at the end of the sweeper charged-particle detectors, as illustrated in Fig. 1.

After isolating the events corresponding to Z = 10, an identification of mass number A is required to select the desired ³⁰Ne isotope. This identification is achieved by a combination of the flight time, position, and angle measurements. A neutron hit is then required in coincidence with the ³⁰Ne fragment to ensure that these events are attributed to the decay of an unbound excited state of ³¹Ne.

The two-body decay energy is reconstructed from the measured four-momenta of the decay products of ³¹Ne^{*} (i.e., ³⁰Ne +n) using the invariant-mass technique. Reconstructing



FIG. 2. Two-body decay energy spectrum of ³¹Ne^{*}. Data are represented by the open circles and the sum of the simulated resonances by a solid black line. The dotted dashed line and dashed line are the individual asymmetric Breit-Wigner resonant lineshapes, which have been subjected to acceptance and efficiency effects.

the four-momentum for the neutron at the target is straightforward using time of flight and position. For the charged fragment, the measured positions and angles in the chargedparticle detectors are used with a third-order ion-optical matrix transformation to determine the angles and positions of the charged fragments at the reaction target. The ion-optical matrix was calculated from the magnetic-field map by using COSY INFINITY [13]. It is assumed that the two-proton knockout and the neutron emission happen at the midpoint of one of the three beryllium targets and the appropriate energy is added back to account for the energy loss of the exiting charged reaction product. The segmented target allows one to properly identify the target segment where the reaction took place [12], and subsequently the added back energy specific to where the two-proton knockout process took place.

Monte Carlo simulations that include the experimental setup and energy-dependent Breit-Wigner lineshapes with theoretical decay energy distributions (including decay modes and reaction mechanisms), beam parameters, and the sweeper magnetic field as well as the detector acceptances, efficiencies, and resolutions were performed. These simulated data sets are directly compared with the experimental data. The neutron interactions were simulated by using the GEANT4 Monte Carlo toolkit [14,15] with the physics class of MENATE_R [16]. The simulations were benchmarked against several experimentally reconstructed observables, i.e., positions, kinetic energies, and relative velocities.

IV. RESULTS AND DISCUSSION

The experimentally reconstructed two-body decay energy spectrum for ³¹Ne* is shown in Fig. 2. The simulation assumes that the emitted decay products are isotropic in the center-of-



FIG. 3. Two-proton knockout cross sections to the first 30 negative-parity unbound states of each spin as a function of the 31 Ne^{*} energy above threshold.

mass frame and a weighted log-likelihood fit is used to extract the resonance parameters represented by asymmetric Breit-Wigner lineshapes.

Two resonant lineshapes at minimum are needed to describe the shape of the two-body decay energy spectrum. The fit using two resonances results in decay energies of 0.30 ± 0.17 and 1.50 ± 0.33 MeV. The uncertainties quoted represent the uncertainties of the log-likelihood fit and include the detector resolutions and acceptance of the experimental setup. The energy resolutions of the experimental setup at these energies are roughly 150 and 370 keV, respectively.

The two-proton removal cross sections populating the negative parity 31 Ne(J^{π}) final states at 88.8 MeV/nucleon on a ⁹Be target are calculated using the eikonal, direct reaction model approach of Ref. [17]. The ground-state to groundstate two-proton separation energy is 40.80 MeV [18], driving the direct nature of the 2p reaction mechanism [19]. Both the $\langle 2p, {}^{31}\text{Ne}(J^{\pi})|^{33}\text{Mg}(3/2^{-})\rangle$ overlaps that generate the two-nucleon amplitudes (TNAs) of the removed sd-shell protons and the spectrum of unbound excited ³¹Ne final states are taken from the FSU interaction [20] shell-model calculations made using the code NUSHELLX [21]. The FSU Hamiltonian uses the sd-pf model space with the constraint that the protons are in the sd shell. The neutron configuration for ³³Mg and ³¹Ne is $(sd)^{(-2)}(fp)^3$, and the neutron configuration for ³⁰Ne is $(sd)^{(-2)}(fp)^2$. Hartree-Fock calculations were used to constrain both, the density of the ³¹Ne residues for the calculation of the residue-target interaction, and the geometry parameters of the real Woods-Saxon potentials used to bind the single-particle orbitals of the removed protons. This follows the method discussed in Ref. [22]. Analogous shell-model calculations are used to compute the $\langle n, {}^{30}\text{Ne}(J_f^{\pi})|^{31}\text{Ne}(J^{\pi})\rangle$ overlaps and spectroscopic factors for the neutron decays of the particle-unbound final states.

The two-proton knockout cross sections to the first 30 negative-parity states of each spin are shown in Fig. 3. A summary of the more strongly populated levels and their cross sections that include spectroscopic factors can be seen in

TABLE I. Calculated cross sections σ for the ³³Mg(-2p)³¹Ne and ³¹Ne^{*} \rightarrow ³⁰Ne + *n* decays. This cross section folds in spectroscopic factors pertaining to a specific decay. E_x is the energy of the ³¹Ne level, E_n is decay energy, and Γ_t is the decay width for each of the spin-parity states J^{π} .

2 J ^π	E_x (MeV)	σ (mb)	J_f^π	E_n (MeV)	Γ_t (MeV)
5-	0.43	0.029	0^+	0.26	0.004
7-	0.70	0.053	0^+	0.53	0.157
9-	1.49	0.097	2^{+}	0.52	0.167
11^{-}	2.13	0.069	2^{+}	1.17	0.367
7-	2.52	0.022	2^{+}	1.56	0.541
5-	3.11	0.027	2^{+}	2.15	0.698
7-	3.72	0.032	4^{+}	1.32	0.423
7-	4.27	0.086	4^{+}	1.86	0.464
7-	4.36	0.022	4+	1.95	0.275
5-	4.53	0.027	2^{+}	3.57	0.468
5-	4.53	0.027	4^{+}	2.13	0.468
7-	4.95	0.054	4^{+}	2.54	0.129
11-	4.99	0.023	4+	2.59	0.215

Table I. Only levels in ³¹Ne with a cross section for the oneneutron decay to ³⁰Ne +*n* larger than 0.020 mb are shown. A level scheme based on experimental data and levels from shell-model calculations can be seen in Fig. 4. The two extracted decay energies from the fit to data are color coded in the figure: red (300 keV) and blue (1.50 MeV). Due to the absence of γ -ray detection, decays to excited states in ³⁰Ne cannot be discriminated but must be considered. The colorcoded lines in the center column illustrate possible levels based on decays to accessible states in the daughter nucleus ³⁰Ne. For example, the red lines are located 300 keV above each ³⁰Ne level.



FIG. 4. Decay energy and level scheme for the decay of ³¹Ne into ³⁰Ne +*n*. In the center are levels potentially observed by the present experiment, color coded by the measured decay energies (see text for detail). The arrows specify potential decay paths to states in ³⁰Ne with colors indicating the measured decay energies. The right spectrum shows the NUSHELLX calculations. The length of the bars (excluding the ground state) indicate a population cross section from the two-proton knockout. Energies listed are relative to the one-neutron separation energy of ³¹Ne.

The measured two-body decay energy represents the energy released when a neutron-unbound excited state of ³¹Ne decays to a state in ³⁰Ne. Determining the energy of a potentially observed level in ³¹Ne relative to its ground state requires knowledge of the one-neutron separation energy, which at present has a large uncertainty, at $S_n = 170 \pm 130$ keV [7]. Potential decay paths from ³¹Ne considered include a decay to either the ground state of ³⁰Ne, the 2⁺ state at 0.79 MeV, or the 4⁺ state at 2.23 MeV [23].

The fit to the decay energy spectrum considers two resonances, which is the minimum needed to describe the overall shape of the measured spectrum. However, shell-model calculations predict multiple strong decays that could possibly be observed with improved resolution. See Table I for details on the decays discussed below.

The first feature of the two-body decay energy is a lowlying peak, fit with a resonant lineshape at $E_n = 0.30 \pm$ 0.17 MeV. The log-likelihood is not sensitive to the width of the lineshape but indicates it is broader than the experimental resolution. The likelihood indicates that a width of at least 1 MeV is needed to fit the decay energy spectrum. The shell-model predicts three states with associated decay energies consistent with this measurement. The $\frac{9}{2}^{-}$ state at 1.49 MeV has the largest cross section from the two-proton knockout at 0.097 mb and the decay to the 2^+ state in 30 Ne would give an E_n of 0.52 MeV. The $\frac{7}{2}^-$ state at 0.70 MeV has a cross section of 0.053 mb and a decay to the ground state of ³⁰Ne would result in an E_n of 0.53 MeV. These two decays are close in energy but populate different daughter states in ³⁰Ne. The absence of γ -ray detection in our setup means these two decays cannot be resolved. A third decay consistent with this energy is of the first-excited state of ³¹Ne, with a spin-parity of $\frac{5}{2}^{-}$ at 0.43 MeV, decaying to the ground state of ³⁰Ne with a decay energy of 0.26 MeV. The widths of these individual decays are narrow (less than 200 keV, see Table I). The large width of the fit used to describe this feature of the spectrum strengthens the interpretation as a group of unresolved resonances.

The second resonance in the maximum likelihood fit is at $E_n = 1.50 \pm 0.33$ MeV above the neutron threshold. There are multiple calculated decays that are consistent with this measurement. The $\frac{11}{2}^-$ at 2.13 MeV decaying to the 2⁺ state in ³⁰Ne is predicted to be a prominent decay with an E_n of 1.17 MeV. Also consistent with this measurement is another prominent decay of the $\frac{7}{2}^{-}$ state at 3.72 MeV to the 4⁺ state of ³⁰Ne with an E_n of 1.32 MeV. It is also worth noting that, although the decay from the first $\frac{9}{2}^{-}$ state in ³¹Ne decaying to the ground state of ³⁰Ne is consistent with this measured decay energy (see Fig. 4), it is outside the model space of the shell-model calculations performed. The single-particle decay width for an l = 5 decay is about 0.026 keV. The spectroscopic factor for mixing an $h_{9/2}$ orbital about 30 MeV above the $0f_{7/2}$ must be less than about 0.01. This indicates a width for the decay to 0^+ being less than 0.26 eV; which is not expected to compete with the decay to 2^+ which has a decay width of 0.167 MeV (see Table I).

The shell-model calculations listed in Table I are a subset of the calculated decays with two-proton knockout cross



FIG. 5. Solid lines represent theoretical decay energy spectra. The larger distribution (top line) assumes one neutron decay for all calculated ³¹Ne states. The smaller distribution (bottom line) includes decays only from states in ³¹Ne below $S_{2n} = 3.72$ MeV. The markers indicate the measured spectrum. The only free parameter for this comparison is the overall scaling.

sections greater than 0.02 mb. Figure 3 shows the cross section calculations have considerable relative strength up to 6.5 MeV, resulting in a large number of possible decays. When the various decays to daughter states in ³⁰Ne are calculated for states with cross sections down to 0.001 mb, there are 140 total decays to consider. Figure 5 shows a comparison of shell-model theory to the measured data. The top line indicates a one-neutron decay for all calculated ³¹Ne states. The bottom line includes only the decays from ³¹Ne states below $S_{2n} = 3.72 \text{ MeV}$ [24]. The shaded area is a difference between these two assumptions: either all ³¹Ne states decay via one neutron decay, or only states below S_{2n} decay via one-neutron decay. In reality, these processes compete and an accurate picture of this complex decay spectrum lies in between the two. The shape of the measured data spectrum is not inconsistent with the predictions of the shell-model calculations.

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V. SUMMARY AND OUTLOOK

This experiment studies the excited states of ³¹Ne for the first time, and the results are interpreted in the context of shell-model calculations. ³¹Ne is a prime example of a halo nucleus in the island of inversion and can give insight into other nuclei in this region.

The neutron decay calculations performed indicate a complex scenario. Calculating the two-proton knockout cross sections up to roughly 6.5 MeV gives more than 100 potential decays to consider. The experimentally measured decay energy spectrum can be fit with two resonant lineshapes, interpreted here as collections of relatively strong narrower decays. The neutron decay calculations show that the neutron decay spectrum is composed of many unresolved peaks, the shape of which is not inconsistent with the observed spectrum.

To get a better understanding of how ³¹Ne is related to other neutron-rich nuclei in the same region, further studies attempting to resolve the various resonances would be useful. The addition of γ -ray detection in a future MoNA-LISA experiment would be well suited to clarify the observation of specific energy levels in ³¹Ne and give a better understanding of the shell evolution in the region of the N = 20 island of inversion.

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