07 Transverse Nonlinear Resonances	Transverse Particle Resonances: Outline
with Application to Circular Accelerators*Prof. Steven M. LundPhysics and Astronomy DepartmentFacility for Rare Isotope Beams (FRIB)Michigan State University (MSU)MSU PHY 905 and US Particle Accelerator School"Accelerator Physics"	Overview Floquet Coordinates and Hill's Equation Perturbed Hill's Equation in Floquet Coordinates Sources of and Forms of Perturbation Terms Solution of the Perturbed Hill's Equation: Resonances Machine Operating Points: Tune Restrictions Resulting from Resonances References
Steven M. Lund and Yue Hao East Lansing, Michigan February, 2020 * Research supported by: (Version 2020220) FRIB/MSU: U.S. Department of Energy O ce of Science Cooperative Agreement DE- SC0000661 and National Science Foundation Grant No. PHY-1102511 SM Lund, MSU & USPAS, 2020 Accelerator Physics 1	SM Lund, MSU & USPAS, 2020 Accelerator Physics 2
Transverse Particle Resonances: Detailed Outline	Transverse Particle Resonances: Detailed Outline - 2
Section headings include embedded links that when clicked on will direct you to the section	<ul> <li>5) Solution of the Perturbed Hill's Equation: Resonances</li> <li>Fourier Expansion of Perturbations and Resonance Terms</li> <li>Resonance Conditions</li> </ul>
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2) Floquet Coordinates and Hill's Equation Transformation of Hill's Equation Phase-Space Structure of Solution	References

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Expression of the Courant-Snyder Invariant

Transformation Result for x-Equation

3) Perturbed Hill's Equation in Floquet Coordinates

4) Sources of and Forms of Perturbation Terms Power Series Expansion of Perturbations Connection to Multipole Field Errors

Phase-Space Area Transform

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### S1: Overview

In our treatment of transverse single particle orbits of lattices with s-varying focusing, we found that Hill's Equation describes the orbits to leading-order approximation:

$$x''(s) + \kappa_x(s)x(s) = 0$$
  
$$y''(s) + \kappa_y(s)y(s) = 0$$

where  $\kappa_x(s)$ ,  $\kappa_y(s)$  are functions that describe linear applied focusing forces of the lattice

Focusing functions can also incorporate linear space-charge forces

- Self-consistent for special case of a KV distribution

In analyzing Hill's equations we employed phase-amplitude methods

See: S.M. Lund lectures on Transverse Particle Dynamics, S8, on the betatron form of the solution

$$\begin{aligned} x(s) &= A_{xi}\sqrt{\beta_x(s)}\cos\psi_x(s) & A_{xi} = \text{const} \\ \frac{1}{2}\beta_x(s)\beta_x''(s) - \frac{1}{4}\beta_x'^2(s) + \kappa_x(s)\beta_x^2(s) = 1 & \psi_x(s) = \psi_{xi} + \int_{s_i}^s \frac{d\bar{s}}{\beta_x(\bar{s})} \\ \beta_x(s + L_p) &= \beta_x(s) & \beta_x(s) > 0 \end{aligned}$$
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These transforms will help us more simply understand the action of perturbations (from applied eld nonlinearities, ....) acting on the particle orbits:

$$\begin{aligned} x''(s) + \kappa_x(s)x(s) &= \mathcal{P}_x(s; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) \\ y''(s) + \kappa_y(s)y(s) &= \mathcal{P}_y(s; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) \\ \mathcal{P}_x, \ \mathcal{P}_y = \text{Perturbations} \\ \vec{\delta} &= \text{Extra Coupling Variables} \end{aligned}$$

For simplicity, we restrict analysis to:

 $\gamma_b \beta_b = \text{const}$  No Acceleration  $\delta = 0$  No Axial Momentum Spread

 $\phi = 0$  Neglect Space-Charge

Weak acceleration can be incorporated using transformations

Lattice Period

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We also take the applied focusing lattice to be periodic with:  

$$\kappa (s + I_{c}) - \kappa (s)$$

$$\frac{\kappa_x(s+L_p) - \kappa_x(s)}{\kappa_y(s+L_p) = \kappa_y(s)} \quad L_p =$$

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This formulation simpli ed identi cation of the Courant-Snyder invariant:

$$\left(\frac{x}{w_x}\right)^2 + (w_x x' - w'_x x)^2 = A_x^2 \equiv \epsilon_x = \text{const}$$

$$\frac{1 + \beta_x'^2/4}{\beta_x} x^2 - \beta_x \beta_x' x x' + \beta_x x'^2 = A_x^2 = \epsilon_x$$

$$\gamma_x x^2 + 2\alpha_x x x' + \beta_x x'^2 =$$

$$w_x = \sqrt{\beta_x}$$

which helped to interpret the dynamics.

We will now exploit this formulation to better (analytically!) understand resonant instabilities in periodic focusing lattices. This is done by choosing coordinates such that *stable* unperturbed orbits described by Hill's equation:

$$x''(s) + \kappa_x(s)x(s) = 0$$

are mapped to a continuous oscillator

$$\tilde{x}''(\tilde{s}) + \tilde{k}_{\beta 0}^2 \tilde{x}(\tilde{s}) = 0$$
  
$$\tilde{k}_{\beta 0}^2 = \text{const} > 0$$
  
$$\tilde{\cdots} = \text{Transformed Coordinate}$$

• Because the linear lattice is designed for single particle stability this transformation can be e ected for any practical machine operating point

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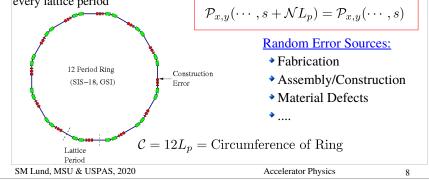
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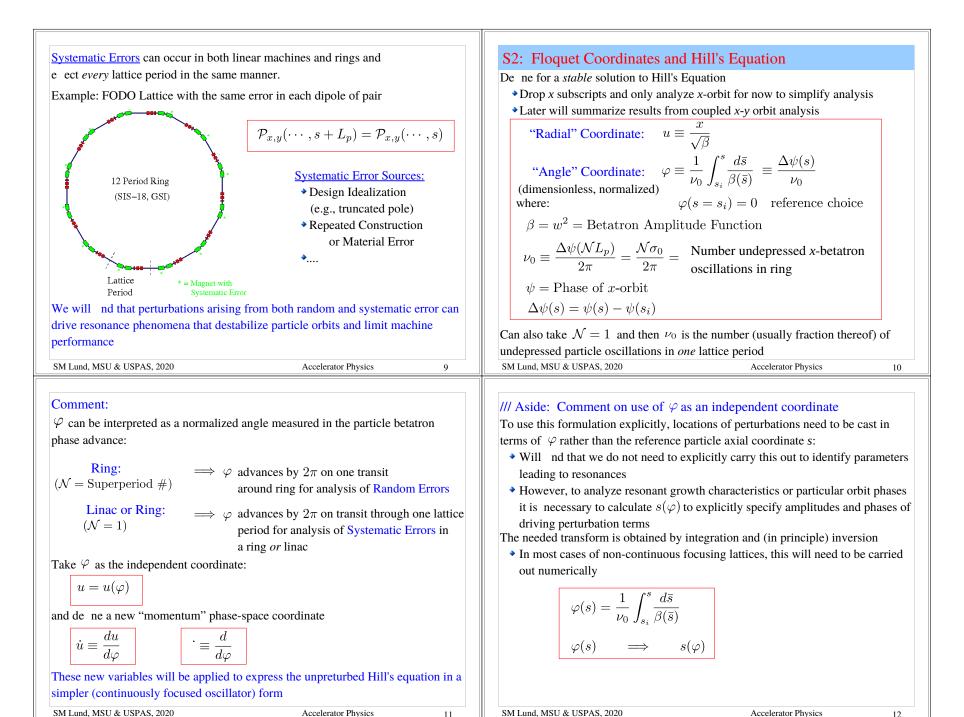
For a ring we also always have the superperiodicity condition:  

$$\begin{array}{l} \mathcal{P}_x(s + \mathcal{C}; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) = \mathcal{P}_x(s; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) \\ \mathcal{P}_y(s + \mathcal{C}; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) = \mathcal{P}_y(s; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) \\ \mathcal{C} = \mathcal{N}L_p = \text{Circumference Ring} \\ \mathcal{N} \equiv \text{Superperiodicity} \end{array}$$

Perturbations can be Random and/or Systematic:

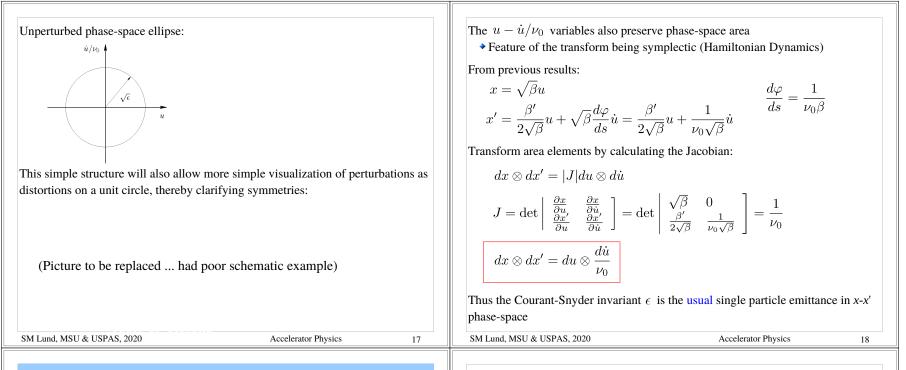
Random Errors in a ring will be felt once per particle lap in the ring rather than every lattice period





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$$\begin{aligned} \varphi(s) &= \frac{1}{\nu_0} \int_{s}^{s} \frac{ds}{\partial(s)} \implies \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta} \\ & \text{Rate of change in s not constant} \\ & \text{except for continuous focusing lattices} \\ & \kappa_x = k_{\beta n}^2 = \text{const} \\ & \frac{d\varphi}{ds} = \frac{2\pi}{c} (s - s_i) \\ & \text{Periodic Focusing: Simple FODD lattice to illustrate} \\ & \text{Add numerical example/plot} \\ & \text{in future version of notes.} \\ & \text{Multiple FODD lattice to illustrate} \\ & \text{Add numerical example/plot} \\ & \text{in future version of notes.} \\ & \frac{d\varphi}{ds} = \frac{2\pi}{c\sqrt{\beta}} u + \frac{1}{\nu_0\sqrt{\beta}} u \frac{d\varphi}{ds} = \frac{1}{c\sqrt{\beta}} du \frac{d\varphi}{ds} \frac{d\varphi}{ds} = \frac{d\varphi}{ds} \frac{d\varphi}{ds} \\ & \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta} \int_{s_1}^{s} \frac{d\varphi}{d(s)} \implies u = \frac{1}{\nu_0\beta} \\ & \text{Multiple FODD lattice to illustrate} \\ & \text{Multiple for some of notes.} \\ & \frac{d'}{ds} = \frac{\beta'}{2\sqrt{\beta}} u + \frac{1}{\nu_0\sqrt{\beta}} u \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta} \\ & \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta} \int_{s_1}^{s} \frac{d\varphi}{d(s)} \implies u = \frac{1}{\nu_0\beta\beta} \frac{d\varphi}{ds} u + \frac{1}{\nu_0\beta\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{2\nu_0\beta^{3/2}} u + \frac{d\varphi}{ds} \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{2\sqrt{\beta}} u + \frac{1}{\nu_0\beta\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{2\nu_0\beta\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{\nu_0\beta\beta^{3/2}} u \\ & \frac{d\varphi}{ds} = \frac{1}{2\nu_0\beta\beta^{3/2}} u \\ & \frac{1}{\nu_0\beta\beta^{3/2}} u \\ & \frac{1}{\nu_0\beta\beta^{$$



#### S3: Perturbed Hill's Equation in Floquet Coordinates

Return to the perturbed Hill's equation in S1:

 $\begin{aligned} x''(s) + \kappa_x(s)x(s) &= \mathcal{P}_x(s; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) \\ y''(s) + \kappa_y(s)y(s) &= \mathcal{P}_y(s; \mathbf{x}_{\perp}, \mathbf{x}'_{\perp}, \vec{\delta}) \\ \mathcal{P}_x, \ \mathcal{P}_y = \text{Perturbations} \\ \vec{\delta} &= \text{Extra Coupling Variables} \end{aligned}$ 

Drop the extra coupling variables and apply the Floquet transform in S2 and consider only transverse multipole magnetic eld perturbations

- Examine only *x*-equation, *y*-equation analogous
- From S4 in Transverse Particle Dynamics terms  $B_x$ ,  $B_y$  only have variation in *x*,*y*. If solenoid magnetic eld errors are put in, terms with x', y'dependence will also be needed
- Drop *x*-subscript in  $\mathcal{P}_x$  to simplify notation

$$\ddot{u}+\nu_0^2 u=\nu_0^2\beta^{3/2}\mathcal{P}$$

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$$\mathcal{P} = \mathcal{P}(s(\varphi), \sqrt{\beta u, y, \delta})$$

Transform *y* similarly to *x* If analyzing general orbit with *x* and *y* motion Accelerator Physics 19 Expand the perturbation in a power series:

- Can be done for *all* physical applied eld perturbations
- Multipole symmetries can be applied to restrict the form of the perturbations
- Perturbations can be random (once per lap; in ring) or systematic (every lattice period; in ring or in linac)

$$\mathcal{P}(x, y, s) = \mathcal{P}_0(y, s) + \mathcal{P}_1(y, s)x + \mathcal{P}_2(y, s)x^2 + \cdots$$
$$= \sum_{n=0}^{\infty} \mathcal{P}_n(y, s)x^n$$

Take:  $x = \sqrt{\beta u}$ 

to obtain:

$$\ddot{u} + \nu_0^2 u = \nu_0^2 \sum_{n=0}^{\infty} \beta^{\frac{n+3}{2}} \mathcal{P}_n(y,s) u^n$$

A similar equation applies in the *y*-plane. SM Lund, MSU & USPAS, 2020

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<b>S4: So</b>	S4: Sources of and Forms of Perturbation Terms					
Within a	2D transver	se model it w	vas shown that t	ransverse applie	d magnetic	eld
				be expanded as:		
_	-	_		-		
	1 - C		-	omponents axial inte	egral 3D compo	nents
Appl	ied electric	elds can be	analogously exp	panded	-	
	$\underline{B}^*(\underline{z}) = B_x^a(x,y) - iB_y^a(x,y) = \sum_{n=1}^{\infty} \underline{b}_n \left(\frac{\underline{z}}{r_p}\right)^{n-1}$					
	n=1 (* $p$ )					
$\underline{b}_n =$	$\underline{b}_n = \text{const} (\text{complex}) \equiv \mathcal{A}_n - i\mathcal{B}_n \qquad \underline{z} = x + iy \qquad i = \sqrt{-1}$					
n =	$n = $ Multipole Index $r_p = $ Aperture "Pipe" Radius					
	$\mathcal{B}_n \Longrightarrow$ "Normal" Multipoles					
	$\mathcal{A}_n \Longrightarrow$ "Skew" Multipoles					
Cartesia	Cartesian projections: $\overline{B_x} - i\overline{B_y} = (\mathcal{A}_n - i\mathcal{B}_n)(x + iy)^{n-1}/r_p^{n-1}$			_		
Index	Name	Norma	$l \left( \mathcal{A}_n = 0 \right)$	Skew (B	$B_n = 0$	
n		$B_x r_p^{n-1} / \mathcal{B}_n$	$B_y r_p^{n-1} / \mathcal{B}_n$	$\frac{B_x r_p^{n-1} / \mathcal{A}_n}{1}$	$B_y r_p^{n-1} / \mathcal{A}_n$	_
1	Dipole	0	1	1		
2	Quadrupole	y	$x_{-}$		-y	
3	Sextupole	2xy	$x^2 - y^2$	$x^2 - y^2$	-2xy	
4	Octupole	$3x^2y - y^3$	$x^{3} - 3xy^{2}$	$ \begin{vmatrix} x \\ x^2 - y^2 \\ x^3 - 3xy^2 \\ x^4 - 6x^2y^2 + y^4 \end{vmatrix} $	$-3x^2y + y^3$	
5	Decapole	$4x^3y - 4xy^3$	$x^4 - 6x^2y^2 + y^4$	$x^4 - 6x^2y^2 + y^4$	$-4x^3y + 4xy^3$	_
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#### // Reminder: Particle equations of motion from Transverse Particle Dynamics lecture notes

Transverse particle equations of motion in explicit component form:

$$\begin{aligned} x'' + \frac{(\gamma_b \beta_b)'}{(\gamma_b \beta_b)} x' &= \frac{q}{m \gamma_b \beta_b^2 c^2} E_x^a - \frac{q}{m \gamma_b \beta_b c} B_y^a + \frac{q}{m \gamma_b \beta_b c} B_z^a y' \\ &- \frac{q}{m \gamma_b^3 \beta_b^2 c^2} \frac{\partial \phi}{\partial x} \\ y'' + \frac{(\gamma_b \beta_b)'}{(\gamma_b \beta_b)} y' &= \frac{q}{m \gamma_b \beta_b^2 c^2} E_y^a + \frac{q}{m \gamma_b \beta_b c} B_x^a - \frac{q}{m \gamma_b \beta_b c} B_z^a x' \\ &- \frac{q}{m \gamma_b^3 \beta_b^2 c^2} \frac{\partial \phi}{\partial y} \end{aligned}$$

Equations previously derived under assumptions:

• No bends ( xed *x-y-z* coordinate system with no local bends)

• Paraxial equations (  $x'^2, y'^2 \ll 1$  )

• No dispersive e ects ( $\beta_b$  same all particles), acceleration allowed ( $\beta_b \neq \text{const}$ )

 $\bullet$  Electrostatic and leading-order (in  $\ \beta_b$  ) self-magnetic interactions

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Trace back how the applied magnetic eld terms enter the *x*-plane equation of motion:

\* See: S2, Transverse Particle Dynamics and reminder on next page

• Apply equation in S2 with:  $\beta_b = \text{const}, \ \phi \simeq \text{const}, \ E_x^a \simeq 0, \ B_z^a \simeq 0$ 

• To include axial  $(B_z^a \neq 0)$  eld errors, follow a similar pattern to generalize

$$x'' = -\frac{q}{m\gamma_b\beta_bc}B_y^a$$

Express this equation as:

$$\begin{aligned} x'' + \kappa_x(s)x &= -\frac{q}{m\gamma_b\beta_bc} \left[ B^a_y(x,y,s) - B^a_y(x,y,s) \Big|_{\text{lin }x\text{-foc}} \right] \\ & \bigstar \\ \text{Nonlinear focusing terms only in []} \end{aligned}$$

\* "Normal" part of linear applied magnetic eld contained in focus function  $\kappa_x$ 

Compare to the form of the perturbed Hill's equation:

$$x'' + \kappa_x x = \mathcal{P}_x = \sum_{n=0}^{\infty} \mathcal{P}_n(y, s) x^n$$
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Reduce the x-plane equation to our situation:  

$$x'' + \frac{(\gamma_b\beta_b)'}{(\gamma_b\beta_b)}x' = \frac{q}{m\gamma_b\beta_b^2c^2} E_x^a - \frac{q}{m\gamma_b\beta_bc} B_y^a + \frac{q}{m\gamma_b\beta_bc} B_z^a y'$$

$$- \frac{q}{m\gamma_b^3\beta_b^2c^2} \frac{\partial \phi}{\partial x}$$

$$0 \text{ No Space-Charge}$$

Giving the equation we are analyzing:

$$\implies \qquad x'' = -\frac{q}{m\gamma_b\beta_bc}B_y^a$$

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Analyze the solution of the perturbed orbit equation:

$$\ddot{u} + \nu_0^2 u = \nu_0^2 \sum_{n=0}^{\infty} \beta^{\frac{n+3}{2}} \mathcal{P}_n(y,s) u^n$$

derived in S4.

To more simply illustrate resonances, we analyze motion in the *x*-plane with:

$$y(s)\equiv 0$$

- Essential character of general analysis illustrated most simply in one plane
- Can generalize by expanding  $\mathcal{P}_n(y,s)$  in a power series in y and generalizing notation to distinguish between Floquet coordinates in the x- and y-planes
- Results in coupled x- and y-equations of motion

ring circumference (random perturbations) or lattice period (systematic):

$$L_p = \text{Lattice Period}$$
$$\mathcal{C} = \mathcal{N}L_p = \text{Ring Circumference}$$
$$\beta(s + L_p) = \beta(s)$$
$$\beta(s + \mathcal{N}L_p) = \beta(s)$$

$$\mathcal{P}_n(y, s + \mathcal{N}L_p) = \mathcal{P}_n(y, s)$$
$$\implies \int \beta^{\frac{n+3}{2}}(s + \mathcal{N}L_p)\mathcal{P}_n(y, s + \mathcal{N}L_p) = \beta^{\frac{n+3}{2}}(s)\mathcal{P}_n(y, s)$$

$$\begin{aligned} \mathcal{P}_n(y,s+L_p) &= \mathcal{P}_n(y,s) \\ &\implies \boxed{\beta^{\frac{n+3}{2}}(s+L_p)\mathcal{P}_n(y,s+L_p) = \beta^{\frac{n+3}{2}}(s)\mathcal{P}_n(y,s)} \end{aligned}$$
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Expand each <i>n</i> -labeled perturbation e	xpansion coe cient in a Fourier ser	ries as:	The perturbed equation of
$\beta^{\frac{n+3}{2}}\mathcal{P}_n(y=0,s) = \sum_{k=\infty}^{k=\infty} 0$	$C_{n,k}e^{ikparphi}$		$\ddot{u} + \nu_0^2 u = \nu_0^2 \sum_{i=1}^{\infty}$
$k=-\infty$	$\left( 1,  \text{Random perturbation} \right)$		
	(once per lap in ring)		Expand the solution as:
$i \equiv \sqrt{-1} \qquad p$ $C_{n,k} = \int_{-\pi/p}^{\pi/p} \frac{d\varphi}{2\pi/p} e^{-ikp\varphi}$ $s = s(\varphi) \qquad \varphi = \int_{0}^{s}$	$\equiv \begin{cases} 1, & \text{Random perturbation} \\ & (\text{once per lap in ring}) \end{cases} \\ \mathcal{N}, & \text{Systematic perturbation} \\ & (\text{every lattice period}) \end{cases} \\ \mathcal{B}^{\frac{n+3}{2}}(s)\mathcal{P}_n(y=0,s) &= \text{const} \\ & \text{complex-} \\ \frac{1}{\nu_0} \frac{d\tilde{s}}{\beta(\tilde{s})} \end{cases}$	ı •valued)	$u = u_0 + \delta u$ where $u_0$ is the solution absence of perturbations: $\ddot{u}_0 + \nu_0^2 u_0 = 0$ Assume small-amplitude
<ul> <li>Can apply to Rings for random period</li> </ul>	0 /- (-)		$ u_0  \gg  \delta u $
or systematic perturbations (with	$p = \mathcal{N}$ )		
<ul> <li>Can apply to linacs for periodic p</li> </ul>	erturbations (every lattice period) with	ith $p = 1$	Then to leading order, the
<ul> <li>Does not apply to random perturb</li> <li>In linac random perturbations will</li> </ul>	pations in a linac vary every lattice period and drive rando	om walk	$\ddot{\delta u} + \nu_0^2 \delta u \simeq \nu_0^2$
type e ects but not resonances SM Lund, MSU & USPAS, 2020	Accelerator Physics	29	SM Lund, MSU & USPAS, 2020

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To obtain the perturbed equation of motion, the unperturbed solution  $u_0$  is inserted on the RHS terms

• Gives simple harmonic oscillator equation with driving terms Solution of the unperturbed orbit is simply expressed as:

of motion becomes:  $k{=}\infty$  $\sum C_{n,k} e^{ikp\varphi} u^n$  $=0 \ k = -\infty$  $u_0 =$ unperturbed solution  $\delta u = \text{perturbation due to errors}$ on to the simple harmonic oscillator equation in the Unperturbed equation of motion e perturbations so that he equation of motion for  $\delta u$  is: Perturbed  $\sum^{k=\infty} C_{n,k} e^{ipk\varphi} u_0^n$ equation of motion  $n=0 \ k=-\infty$ Accelerator Physics 30 iu, mou & uspas, 202

Using this expansion the linearized perturbed equation of motion becomes:  $\vec{\delta u} + \nu_0^2 \delta u \simeq \nu_0^2 \sum_{n=0}^{\infty} \sum_{k=-\infty}^{k=\infty} \sum_{m=0}^n \binom{n}{m} \frac{C_{n,k}}{2^n} e^{i[(n-2m)\nu_0 + pk]\varphi} e^{i(n-2m)\varphi_i}$ 

The solution for  $\delta u$  can be expanded as:

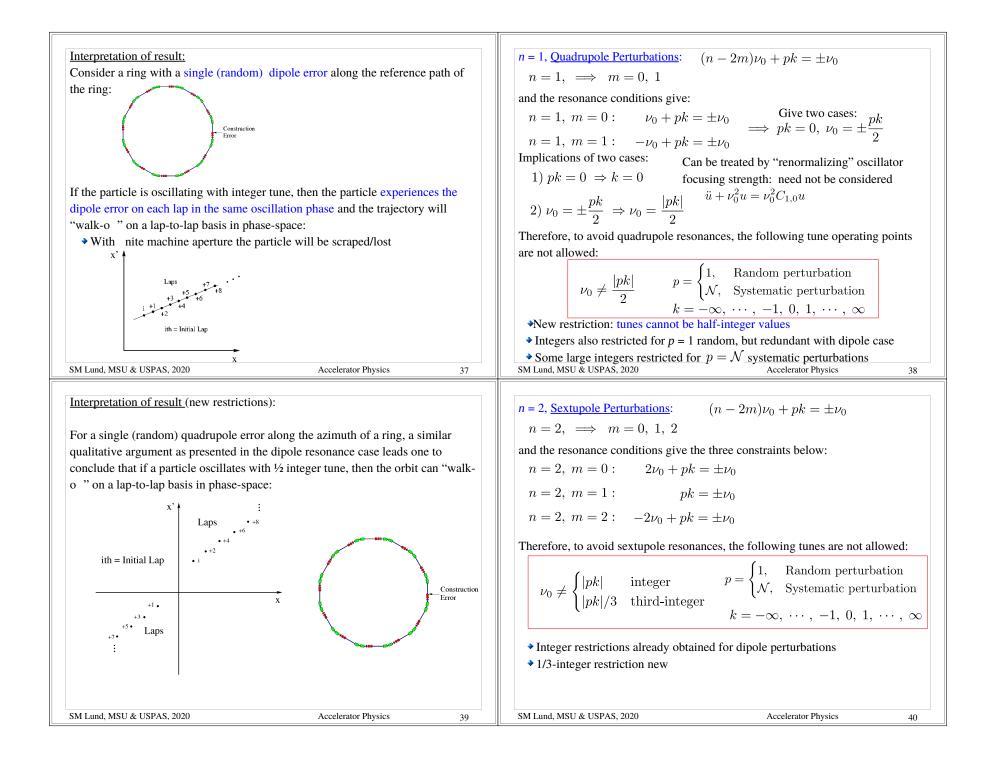
$$\begin{split} \delta u &= \delta u_h + \delta u_p \\ \delta u_h &= \text{homogenous solution} \\ & \text{General solution to:} \quad \ddot{\delta u}_h + \nu_0^2 \delta u_h = 0 \\ \delta u_p &= \text{particular solution} \\ & \text{Any solution with:} \quad \delta u \to \delta u_p \end{split}$$
• Can drop homogeneous solution because it can be absorbed in unperturbed solution  $u_0$ 
- Exception: some classes of linear amplitude errors in adjusting magnets
• Only a particular solution need be found, take:

 $\delta u = \delta u_p$ 

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$\begin{split} & \dot{\delta u} + \nu_0^2 \delta u \simeq \nu_0^2 \sum_{n=0}^{\infty} \sum_{k=-\infty}^{k=\infty} \sum_{m=0}^n \binom{n}{m} \frac{C_{n,k}}{2^n} e^{i[(n-2m)\nu_0 + pk]\varphi} e^{i(n-2m)\varphi_i} \\ & \text{Equation describes a driven simple harmonic oscillator with periodic driving terms on the RHS:} \\ \bullet  Homework problem reviews that solution of such an equation will be unstable when the driving term has a frequency component equal to the restoring term - Resonant exchange and amplitude grows linearly (not exponential!) in \varphi - Parameters meeting resonance condition will lead to instabilities with particle oscillation amplitude growing in \varphi(s) Resonances occur when:\begin{aligned} & (n-2m)\nu_0 + pk = \pm \nu_0 \\ \text{is satis ed for the operating tune } \nu_0 \text{ and some values of:} \\ & n = 0, 1, 2, \cdots, m = 0, 1, 2, \cdots, n \\ & k = -\infty, \cdots, -1, 0, 1, \cdots, \infty \\ & p \equiv \begin{cases} 1, & \text{Random perturbation} \\ \mathcal{N}, & \text{Systematic perturbation} \end{cases} \end{split}$	<ul> <li>If growth rate is su ciently large, machine operating points satisfying the resonance condition will be problematic since particles will be lost (scraped) by the machine aperture due to increasing oscillation amplitude:</li> <li>Machine operating tune (ν<sub>0</sub>) can be adjusted to avoid</li> <li>Perturbation can be actively corrected to reduce amplitude of driving term</li> <li>Low order resonance terms with smaller <i>n</i>, <i>k</i>, <i>m</i> magnitudes are expected to be more dangerous because:</li> <li>Less likely to be washed out by e ects not included in model</li> <li>Amplitude coe cients expected to be stronger</li> <li>More detailed theories consider coherence length, nite amplitude, and nonlinear term e ects. Such treatments and numerical analysis concretely motivate importance/strength of terms. A standard reference on analytic theory is:</li> <li>Kolomenskii and Lebedev, <i>Theory of Circular Accelerators</i>, North-Holland (1966)</li> <li>We only consider lowest order e ects in these notes.</li> <li>In the next section we will examine how resonances restrict possible machine operating parameters.</li> </ul>
Image: Provide state of the state of th	SM Lund, MSU & USPAS, 2020 Accelerator Physics 34
S6: Machine Operating Points: Tune Restrictions Resulting from ResonancesExamine situations where the x-plane motion resonance condition: $(n-2m)\nu_0 + pk = \pm \nu_0$ is satis ed for the operating tune $\nu_0$ and some values of: $(n-2m)\nu_0 + pk = \pm \nu_0$ is satis ed for the operating tune $\nu_0$ and some values of: 	$\begin{array}{l} \begin{array}{l} \textbf{n}=\textbf{0}, \mbox{Dipole Perturbations:} & (n-2m)\nu_0+pk=\pm\nu_0\\ n=0, \implies m=0\\ \mbox{and the resonance condition gives a single constraint:}\\ \nu_0=\pm pk \qquad pk=\mbox{ integer } k=-\infty,\ \cdots,\ -1,\ 0,\ 1,\ \cdots,\ \infty\\ p=\begin{cases} 1, & \mbox{Random perturbation}\\ \mathcal{N}, & \mbox{Systematic perturbation} \end{cases} \\ \mbox{Therefore, to avoid dipole resonances integer tunes} operating points not allowed:}\\ \hline p=1 & \mbox{Random Perturbation} & \nu_0\neq 1,\ 2,\ 3,\ \cdots\\ p=\mathcal{N} & \mbox{Systematic Perturbation} & \nu_0\neq\mathcal{N},\ 2\mathcal{N},\ 3\mathcal{N},\ \cdots\\ \end{array} \\ \hline \mbox{Systematic errors are signi cantly less restrictive on machine operating points} for large \ \mathcal{N}\\ \mbox{-Illustrates why high symmetry is desirable} \end{array}$
Resonances can be analyzed one at a time using linear superposition Analysis valid for small-amplitudes	<ul> <li>Racetracks with N = 2 can be problem</li> <li>Multiply random perturbation tune restrictions by N to obtain the systematic perturbation case</li> </ul>
Analyze resonance possibilities starting with index $n \leq =>$ Multipole Order	perturbation case



$n = 3, \text{ Octupole Perturbations:} (n - 2m)\nu_0 + pk = \pm\nu_0$ $n = 3, \implies m = 0, 1, 2, 3$ and the resonance conditions give the three constraints below: $n = 3, m = 0: \qquad 3\nu_0 + pk = \pm\nu_0$	Higher-order $(n > 2)$ cases analyzed analogously: $(n - 2m)\nu_0 + pk = \pm \nu_0$ • Produce more constraints but expected to be weaker as order increases • Will always generate a new constraint $\nu_0 \neq \frac{ pk }{n+1}$
$n = 3, m = 1:  \nu_0 + pk = \pm \nu_0$ $n = 3, m = 2:  -\nu_0 + pk = \pm \nu_0$ $n = 3, m = 3:  -3\nu_0 + pk = \pm \nu_0$ Therefore, to avoid octupole resonances, the following tunes are not allowed: $\nu_0 \neq \begin{cases}  pk /2 & \text{half-integer} \\  pk /4 & \text{quarter-integer} \end{cases} p = \begin{cases} 1, & \text{Random perturbation} \\ \mathcal{N}, & \text{Systematic perturbation} \\ k = -\infty, \cdots, -1, 0, 1, \cdots, \infty \end{cases}$	
<ul> <li>1/2 - integer restrictions already obtained for quadrupole perturbations</li> <li>1/4-integer restriction new</li> <li>Higher-order (n &gt; 3) cases analyzed analogously as n increases</li> <li>Resonances expected to be weaker as order increases</li> </ul>	
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#### General form of resonance condition

The general resonance condition (all *n*-values) for *x*-plane motion can be summarized as:

 $M\nu_0 = N$  M, N =Integers of same sign |M| ="Order" of resonance

• Higher order numbers *M* are typically less dangerous

- Longer coherence length for validity of theory: e ects not included can "wash-out" the resonance

- Coe cients generally smaller

Particle motion is not (measure zero) really restricted to the *x*-plane, and a more complete analysis taking into account coupled *x*- and *y*-plane motion shows that the generalized resonance condition is:

 Place unperturbed y-orbit in rhs perturbation term, then leading-order expand analogously to x-case to obtain additional driving terms

 $M_x \nu_{0x} + M_y \nu_{0y} = N \qquad M_x, \ M_y, \ N = \text{Integers of same sign}$  $\nu_{0x} = x \text{-plane tune} \qquad |M_x| + |M_y| = \text{"Order" of resonance}$  $\nu_{0y} = y \text{-plane tune}$ • Lower order resonances are more dangerous analogously to x-case

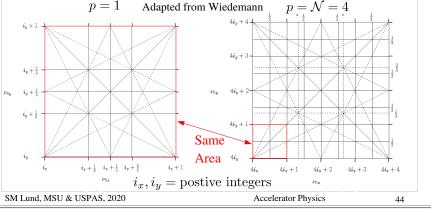
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#### Restrictions on machine operating points

Tune restrictions are typically plotted in  $\nu_{0x} - \nu_{0y}$  space order-by-order up to a max order value to nd allowed tunes where the machine can safely operate

- Often 3<sup>rd</sup> order is chosen as a maximum to avoid
- Cases for random (p = 1) and systematic (p = N) perturbations considered

# Machine operating points chosen as far as possible from low order resonance lines Random Perturbations Systematic Perturbations



<ul> <li>Errors always present and give low-order resonances</li> <li>Usually have weak amplitude coe cients <ul> <li>Can be corrected/compensated to reduce e ects</li> </ul> </li> <li>Systematic Errors: <ul> <li>Lead to higher-order resonances for large N and a lower density of resonance lines (see plots on previous slide comparing the equal boxed red areas) <ul> <li>Large symmetric rings with high N values have less operating restrictions from systematic errors</li> <li>Practical issues such as construction cost and getting the beam into and out of the ring can lead to smaller N values (racetrack lattice)</li> </ul> </li> <li>BUT systematic e cets accumulate in amplitude period by period</li> <li>Resonances beyond 3<sup>rd</sup> order rarely need be considered</li> <li>E ects outside of model assumed tend to wash-out higher order resonances</li> <li>More detailed treatments calculate amplitudes/strengths of resonant terms</li> <li>See accelerator physics references: <ul> <li>Further info: Wiedemann, Paricle Accelerator Physics (2007)</li> </ul> </li> </ul></li></ul>	Discussion: Restrictions on machine operating points	Notation/Nomenclature: Laslett Limits
SM Lund, MSU & USPAS, 2020 Accelerator Physics 45 SM Lund, MSU & USPAS, 2020 Accelerator Physics 46	<ul> <li>Errors always present and give low-order resonances</li> <li>Usually have weak amplitude coe cients <ul> <li>Can be corrected/compensated to reduce e ects</li> </ul> </li> <li>Systematic Errors: <ul> <li>Lead to higher-order resonances for large N and a lower density of resonance lines (see plots on previous slide comparing the equal boxed red areas) <ul> <li>Large symmetric rings with high N values have less operating restrictions from systematic errors</li> <li>Practical issues such as construction cost and getting the beam into and out of the ring can lead to smaller N values (racetrack lattice)</li> </ul> </li> <li>BUT systematic e ects accumulate in amplitude period by period</li> <li>Resonances beyond 3<sup>rd</sup> order rarely need be considered</li> <li>E ects outside of model assumed tend to wash-out higher order resonances</li> <li>More detailed treatments calculate amplitudes/strengths of resonant terms</li> <li>See accelerator physics references:</li> </ul> </li> </ul>	<ul> <li>Oscillation amplitudes increase (spoiling beam quality and control)</li> <li>Particles can be lost</li> <li>Tune shift limits of machine operation are often named "Laslett Limits" in honor of Jackson Laslett who rst calculated tune shift limits for various processes:</li> <li>Image charges</li> <li>Image currents</li> <li>Internal beam self- elds</li> <li></li> <li>Processes shifting resonances can be grouped into two broad categories:</li> <li>Coherent Same for every particle in distribution <ul> <li>Usually most dangerous: full beam resonant</li> <li>Incoherent Di erent for particles in separate parts of the distribution</li> </ul> </li> </ul>

## Corrections and suggestions for improvements welcome!

These notes will be corrected and expanded for reference and for use in future editions of US Particle Accelerator School (USPAS) and Michigan State University (MSU) courses. Contact:

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Please provide corrections with respect to the present archived version at:

https://people.nscl.msu.edu/~lund/msu/phy905\_2020/

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# References:

E. D. Courant and H. S. Snyder, "Theory of the Alternating Gradient Synchrotron," Annals of Physics **3**, 1 (1958)

H. Wiedemann, *Particle Accelerator Physics*, Third Edition, Springer-Verlag (2007)

A. Dragt, "Lectures on Nonlinear Orbit Dynamics," in *Physics of High Energy Accelerators*, edited by R.A. Carrigan, F.R. Hudson, and M. Month (AIP Conf. Proc. No. 87, 1982) p. 147

D. A. Edwards and M. J. Syphers, *An Introduction to the Physics of High Energy Accelerators*, Wiley (1993)

F. Sacherer, *Transverse Space-Charge E ects in Circular Accelerators*, Univ. of California Berkeley, Ph.D Thesis (1968)

A. A. Kolomenskii and A. N. Lebedev, *Theory of Circular Accelerators*, North-Holland (1966)