| Transverse Particle Resonances | Transverse Particle Resonances: Outline |
| :---: | :---: |
| with Application to Circular Accelerators* <br> Prof. Steven M. Lund <br> Physics and Astronomy Department <br> Facility for Rare Isotope Beams (FRIB) <br> Michigan State University (MSU) <br> US Particle Accelerator School (USPAS) Lectures on <br> "Beam Physics with Intense Space-Charge" <br> Steven M. Lund, John J. Barnard, and Arun Persaud <br> US Particle Accelerator School Winter Session <br> UC San Diego, 13-24 January, 2020 <br> * Research supported by: <br> (Version 20200119) <br> FRIB/MSU, 2014 onward via: U.S. Department of Energy O ce of Science Cooperative Agreement DE-SC0000661and National Science Foundation Grant No. PHY-1102511 and <br> LLNL/LBNL, before 2014 via: US Dept. of Energy Contract Nos. DE-AC52-07NA27344 and DE-AC02-05CH11231 | Overview <br> Floquet Coordinates and Hill's Equation <br> Perturbed Hill's Equation in Floquet Coordinates <br> Sources of and Forms of Perturbation Terms <br> Solution of the Perturbed Hill's Equation: Resonances <br> Machine Operating Points: Tune Restrictions Resulting from Resonances Space-Charge E ects on Particle Resonances <br> References |
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## S1: Overview

In our treatment of transverse single particle orbits of lattices with s-varying focusing, we found that Hill's Equation describes the orbits to leading-order approximation:

$$
\begin{aligned}
& x^{\prime \prime}(s)+\kappa_{x}(s) x(s)=0 \\
& y^{\prime \prime}(s)+\kappa_{y}(s) y(s)=0
\end{aligned}
$$

where $\kappa_{x}(s), \quad \kappa_{y}(s)$ are functions that describe linear applied focusing forces of the lattice
*Focusing functions can also incorporate linear space-charge forces

- Self-consistent for special case of a KV distribution

In analyzing Hill's equations we employed phase-amplitude methods
$\rightarrow$ See: S.M. Lund lectures on Transverse Particle Dynamics, S8, on the betatron form of the solution

$$
\begin{array}{crl}
x(s)=A_{x i} \sqrt{\beta_{x}(s)} \cos \psi_{x}(s) & A_{x i}=\mathrm{const} \\
\frac{1}{2} \beta_{x}(s) \beta_{x}^{\prime \prime}(s)-\frac{1}{4} \beta_{x}^{\prime 2}(s)+\kappa_{x}(s) \beta_{x}^{2}(s)=1 & \psi_{x}(s)=\psi_{x i}+\int_{s_{i}}^{s} \frac{d \bar{s}}{\beta_{x}(\bar{s})} \\
\beta_{x}\left(s+L_{p}\right)=\beta_{x}(s) \quad \beta_{x}(s)>0 &
\end{array}
$$

$$
\begin{array}{lll}
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\hline
\end{array}
$$

These transforms will help us more simply understand the action of perturbations (from applied eld nonlinearities, ....) acting on the particle orbits:

$$
\begin{gathered}
x^{\prime \prime}(s)+\kappa_{x}(s) x(s)=\mathcal{P}_{x}\left(s ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right) \\
y^{\prime \prime}(s)+\kappa_{y}(s) y(s)=\mathcal{P}_{y}\left(s ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right) \\
\mathcal{P}_{x}, \mathcal{P}_{y}=\text { Perturbations } \\
\vec{\delta}=\text { Extra Coupling Variables }
\end{gathered}
$$

For simplicity, we restrict analysis to:

$$
\begin{aligned}
\gamma_{b} \beta_{b} & =\text { const } & & \text { No Acceleration } \\
\delta & =0 & & \text { No Axial Momentum Spread } \\
\phi & =0 & & \text { Neglect Space-Charge }
\end{aligned}
$$

- Acceleration can be incorporated using transformations
(see Transverse Particle Dynamics, S10)
$\rightarrow$ A limited analysis of space-charge e ects will be made in S7
We also take the applied focusing lattice to be periodic with:

$$
\begin{aligned}
& \kappa_{x}\left(s+L_{p}\right)=\kappa_{x}(s) \\
& \kappa_{y}\left(s+L_{p}\right)=\kappa_{y}(s)
\end{aligned} \quad L_{p}=\text { Lattice Period }
$$

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This formulation simpli ed identi cation of the Courant-Snyder invariant

$$
\begin{aligned}
\left(\frac{x}{w_{x}}\right)^{2}+\left(w_{x} x^{\prime}-w_{x}^{\prime} x\right)^{2} & =A_{x}^{2} \equiv \epsilon_{x}=\mathrm{const} \\
\frac{1+\beta_{x}^{\prime 2} / 4}{\beta_{x}} x^{2}-\beta_{x} \beta_{x}^{\prime} x x^{\prime}+\beta_{x} x^{\prime 2}=A_{x}^{2} & =\epsilon_{x} \\
\gamma_{x} x^{2}+2 \alpha_{x} x x^{\prime}+\beta_{x} x^{\prime 2} & =
\end{aligned}
$$

which helped to interpret the dynamics.
We will now exploit this formulation to better (analytically!) understand resonant instabilities in periodic focusing lattices. This is done by choosing coordinates such that stable unperturbed orbits described by Hill's equation:

$$
x^{\prime \prime}(s)+\kappa_{x}(s) x(s)=0
$$

are mapped to a continuous oscillator

$$
\begin{gathered}
\tilde{x}^{\prime \prime}(\tilde{s})+\tilde{k}_{\beta 0}^{2} \tilde{x}(\tilde{s})=0 \\
\tilde{k}_{\beta 0}^{2}=\mathrm{const}>0
\end{gathered} \quad . \tilde{\sim}=\text { Transformed Coordinate }
$$

$\rightarrow$ Because the linear lattice is designed for single particle stability this transformation can be e ected for any practical machine operating point SM Lund, USPAS, 2020

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For a ring we also always have the superperiodicity condition:

$$
\begin{gathered}
\mathcal{P}_{x}\left(s+\mathcal{C} ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right)=\mathcal{P}_{x}\left(s ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right) \\
\mathcal{P}_{y}\left(s+\mathcal{C} ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right)=\mathcal{P}_{y}\left(s ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right) \\
\mathcal{C}=\mathcal{N} L_{p}=\text { Circumference Ring } \\
\mathcal{N} \equiv \text { Superperiodicity }
\end{gathered}
$$

Perturbations can be Random and/or Systematic:
Random Errors in a ring will be felt once per particle lap in the ring rather than every lattice period


Systematic Errors can occur in both linear machines and rings and e ect every lattice period in the same manner.
Example: FODO Lattice with the same error in each dipole of pair


We will nd that perturbations arising from both random and systematic error can drive resonance phenomena that destabilize particle orbits and limit machine performance

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## Comment:

$\varphi$ can be interpreted as a normalized angle measured in the particle betatron phase advance:

$$
\begin{array}{cl}
\begin{array}{c}
\text { Ring: } \\
(\mathcal{N}=\text { Superperiod \#) }
\end{array} & \Longrightarrow \varphi \text { advances by } 2 \pi \text { on one transit } \\
\text { around ring for analysis of Random Errors } \\
\text { Linac or Ring: } & \Longrightarrow \varphi \text { advances by } 2 \pi \text { on transit through one lattice } \\
(\mathcal{N}=1) & \begin{array}{l}
\text { period for analysis of Systematic Errors in } \\
\\
\end{array}
\end{array}
$$

Take $\varphi$ as the independent coordinate:

$$
u=u(\varphi)
$$

and de ne a new "momentum" phase-space coordinate

$$
\dot{u} \equiv \frac{d u}{d \varphi} \quad \cdot \equiv \frac{d}{d \varphi}
$$

These new variables will be applied to express the unpreturbed Hill's equation in a simpler (continuously focused oscillator) form

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## S2: Floquet Coordinates and Hill's Equation

## De ne for a stable solution to Hill's Equation

$\rightarrow$ Drop $x$ subscripts and only analyze $x$-orbit for now to simplify analysis
$\rightarrow$ Later will summarize results from coupled $x-y$ orbit analysis

$$
\begin{aligned}
& \text { "Radial" Coordinate: } u \equiv \frac{x}{\sqrt{\beta}} \\
& \text { "Angle" Coordinate: } \varphi \equiv \frac{1}{\nu_{0}} \int_{s_{i}}^{s} \frac{d \bar{s}}{\beta(\bar{s})} \equiv \frac{\Delta \psi(s)}{\nu_{0}} \\
& \text { (dimensionless, normalized) } \\
& \text { where: }
\end{aligned} \quad \varphi\left(s=s_{i}\right)=0 \quad \text { reference choice } \quad l
$$

$$
\beta=w^{2}=\text { Betatron Amplitude Function }
$$

$$
\nu_{0} \equiv \frac{\Delta \psi\left(\mathcal{N} L_{p}\right)}{2 \pi}=\frac{\mathcal{N} \sigma_{0}}{2 \pi}=\begin{aligned}
& \text { Number undepressed } x \text {-betatron } \\
& \text { oscillations in ring }
\end{aligned}
$$

$$
\psi=\text { Phase of } x \text {-orbit }
$$

$$
\Delta \psi(s)=\psi(s)-\psi\left(s_{i}\right)
$$

Can also take $\mathcal{N}=1$ and then $\nu_{0}$ is the number (usually fraction thereof) of undepressed particle oscillations in one lattice period
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/// Aside: Comment on use of $\varphi$ as an independent coordinate
To use this formulation explicitly, locations of perturbations need to be cast in terms of $\varphi$ rather than the reference particle axial coordinate s:
$\rightarrow$ Will nd that we do not need to explicitly carry this out to identify parameters leading to resonances

* However, to analyze resonant growth characteristics or particular orbit phases it is necessary to calculate $s(\varphi)$ to explicitly specify amplitudes and phases of driving perturbation terms
The needed transform is obtained by integration and (in principle) inversion
$\rightarrow$ In most cases of non-continuous focusing lattices, this will need to be carried out numerically

$$
\begin{aligned}
& \varphi(s)=\frac{1}{\nu_{0}} \int_{s_{i}}^{s} \frac{d \bar{s}}{\beta(\bar{s})} \\
& \varphi(s) \quad \Longrightarrow \quad s(\varphi)
\end{aligned}
$$

$$
\varphi(s) \equiv \frac{1}{\nu_{0}} \int_{s_{i}}^{s} \frac{d \bar{s}}{\beta(\bar{s})} \quad \Longrightarrow \quad \frac{d \varphi}{d s}=\frac{1}{\nu_{0} \beta}
$$

Rate of change in s not constant except for continuous focusing lattices
Continuous Focusing: Simplest case

$$
\begin{aligned}
\kappa_{x} & =k_{\beta 0}^{2}=\mathrm{const} \\
\frac{d \varphi}{d s} & =\frac{2 \pi}{\mathcal{C}}
\end{aligned}=\mathrm{const} \quad \Longrightarrow \varphi(s)=\frac{2 \pi}{\mathcal{C}}\left(s-s_{i}\right)
$$

Periodic Focusing: Simple FODO lattice to illustrate

## Add numerical example/plot

in future version of notes.

## From the de nition

$$
u \equiv \frac{x}{\sqrt{\beta}}
$$

Rearranging this and using the chain rule with $u=u(\varphi), \quad \beta=\beta(s)$

$$
\begin{array}{rlr}
x & =\sqrt{\beta} u & \\
x^{\prime} & =\frac{\beta^{\prime}}{2 \sqrt{\beta}} u+\sqrt{\beta} \frac{d u}{d \varphi} \frac{d \varphi}{d s} & \frac{d}{d s}=\frac{d \varphi}{d s} \frac{d}{d \varphi}
\end{array}
$$

From:

$$
\varphi \equiv \frac{1}{\nu_{0}} \int_{s_{i}}^{s} \frac{d \bar{s}}{\beta(\bar{s})} \quad \Longrightarrow \frac{d \varphi}{d s}=\frac{1}{\nu_{0} \beta}
$$

we obtain

$$
\begin{aligned}
x^{\prime} & =\frac{\beta^{\prime}}{2 \sqrt{\beta}} u+\frac{1}{\nu_{0} \sqrt{\beta}} \dot{u} \\
x^{\prime \prime} & =\frac{d}{d s} x^{\prime}=\frac{\beta^{\prime \prime}}{2 \sqrt{\beta}} u-\frac{\beta^{\prime 2}}{4 \beta^{3 / 2}} u+\frac{\beta^{\prime}}{2 \nu_{0} \beta^{3 / 2}} \dot{u}-\frac{\beta^{\prime}}{2 \nu_{0} \beta^{3 / 2}} \dot{u}+\frac{1}{\nu_{0}^{2} \beta^{3 / 2}} \ddot{u}
\end{aligned}
$$

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The general solution to the unperturbed simple harmonic oscillator equation can be expressed as:

$$
\begin{array}{ll}
u(\varphi)=u_{i} \cos \left(\nu_{0} \varphi\right)+\frac{\dot{u}_{i}}{\nu_{0}} \sin \left(\nu_{0} \varphi\right) \\
\frac{\dot{u}(\varphi)}{\nu_{0}}=-u_{i} \sin \left(\nu_{0} \varphi\right)+\frac{\dot{u}_{i}}{\nu_{0}} \cos \left(\nu_{0} \varphi\right) \\
u(\varphi=0)=u_{i}=\mathrm{const} & \mathrm{u}_{i} \text { and } \dot{u}_{i} \text { set by } x, x^{\prime} \\
\dot{u}(\varphi=0)=\dot{u}_{i}=\mathrm{const} & \text { initial conditions at } s=s_{i} \\
& \text { (phase choice } \varphi=0 \text { at } s=s_{i}
\end{array}
$$

Floquet representation simpli es interpretation of the Courant-Snyder invariant:
$u^{2}+\left(\frac{\dot{u}}{\nu_{0}}\right)^{2}=u_{i}^{2}\left[\sin ^{2}\left(\nu_{0} \varphi\right)+\underset{1}{\operatorname{qos}^{2}\left(\nu_{0} \varphi\right)+\left(\frac{\dot{u}_{i}}{\nu_{0}}\right)^{2}\left[\sin ^{2}\left(\nu_{0} \varphi\right)+\cos ^{2}\left(\nu_{0} \varphi\right)\right]}\right.$
$+u_{i} \frac{\dot{u}_{i}}{\nu_{0}}\left[\sin \left(\nu_{0} \varphi\right) \cos \left(\nu_{0} \varphi\right)-\sin \left(\nu_{0} \varphi\right) \cos \left(\nu_{0} \varphi\right)\right]$
$\Longrightarrow u^{2}+\left(\frac{\dot{u}}{\nu_{0}}\right)^{2}=u_{i}^{2}+\left(\frac{\dot{u}_{i}}{\nu_{0}}\right)^{2} \equiv \epsilon=\mathrm{const}$
$\rightarrow$ Unperturbed phase-space in $u-\dot{u} / \nu_{0}$ variables is a circle of area $\pi \epsilon$ !
$\rightarrow$ Relate this area to $x-x^{\prime}$ phase-space area shortly

- Preview: areas are equal due to the transform being symplectic
- Same symbols used for area as in Transverse Particle Dynamics is on purpose equation to a simple harmonic oscillator! SM Lund, USPAS, 2020


## Unperturbed phase-space ellipse:



This simple structure will also allow more simple visualization of perturbations as distortions on a unit circle, thereby clarifying symmetries:
(Picture to be replaced ... had poor schematic example)

## S3: Perturbed Hill's Equation in Floquet Coordinates

Return to the perturbed Hill's equation in S1:

$$
\begin{gathered}
x^{\prime \prime}(s)+\kappa_{x}(s) x(s)=\mathcal{P}_{x}\left(s ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right) \\
y^{\prime \prime}(s)+\kappa_{y}(s) y(s)=\mathcal{P}_{y}\left(s ; \mathbf{x}_{\perp}, \mathbf{x}_{\perp}^{\prime}, \vec{\delta}\right) \\
\mathcal{P}_{x}, \mathcal{P}_{y}=\text { Perturbations } \\
\vec{\delta}=\text { Extra Coupling Variables }
\end{gathered}
$$

Drop the extra coupling variables and apply the Floquet transform in S2 and consider only transverse multipole magnetic eld perturbations

* Examine only $x$-equation, $y$-equation analogous
* From S4 in Transverse Particle Dynamics terms $B_{x}, B_{y}$ only have variation in $x, y$. If solenoid magnetic eld errors are put in, terms with $x^{\prime}, y^{\prime}$
dependence will also be needed
- Drop $x$-subscript in $\mathcal{P}_{x}$ to simplify notation

$$
\ddot{u}+\nu_{0}^{2} u=\nu_{0}^{2} \beta^{3 / 2} \mathcal{P}
$$

$$
\mathcal{P}=\mathcal{P}(s(\varphi), \sqrt{\beta} u, y, \vec{\delta})
$$

Transform $y$ similarly to $x$ If analyzing general orbit with $x$ and $y$ motion
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The $u-\dot{u} / \nu_{0}$ variables also preserve phase-space area
$\rightarrow$ Feature of the transform being symplectic (Hamiltonian Dynamics)
From previous results:

$$
\begin{array}{rlr}
x & =\sqrt{\beta} u & \frac{d \varphi}{d s}=\frac{1}{\nu_{0} \beta} \\
x^{\prime} & =\frac{\beta^{\prime}}{2 \sqrt{\beta}} u+\sqrt{\beta} \frac{d \varphi}{d s} \dot{u}=\frac{\beta^{\prime}}{2 \sqrt{\beta}} u+\frac{1}{\nu_{0} \sqrt{\beta}} \dot{u} &
\end{array}
$$

Transform area elements by calculating the Jacobian:

$$
\begin{aligned}
& d x \otimes d x^{\prime}=|J| d u \otimes d \dot{u} \\
& \left.\left.J=\operatorname{det} \left\lvert\, \begin{array}{ll}
\frac{\partial x}{\partial u} & \frac{\partial x}{\partial \dot{u}} \\
\frac{\partial x^{\prime}}{\partial u} & \frac{\partial x^{\prime}}{\partial \dot{u}}
\end{array}\right.\right]=\operatorname{det} \left\lvert\, \begin{array}{ll}
\sqrt{\beta} & 0 \\
\frac{\beta^{\prime}}{2 \sqrt{\beta}} & \frac{1}{\nu_{0} \sqrt{\beta}}
\end{array}\right.\right]=\frac{1}{\nu_{0}} \\
& d x \otimes d x^{\prime}=d u \otimes \frac{d \dot{u}}{\nu_{0}}
\end{aligned}
$$

Thus the Courant-Snyder invariant $\epsilon$ is the usual single particle emittance in $x-x^{\prime}$ phase-space; see lectures on Transverse Dynamics, S7
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## Expand the perturbation in a power series

$\rightarrow$ Can be done for all physical applied eld perturbations

- Multipole symmetries can be applied to restrict the form of the perturbations
- See: S4 in these notes and S3 in Transverse Particle Dynamics
$\rightarrow$ Perturbations can be random (once per lap; in ring) or systematic (every lattice period; in ring or in linac)

$$
\begin{aligned}
\mathcal{P}(x, y, s) & =\mathcal{P}_{0}(y, s)+\mathcal{P}_{1}(y, s) x+\mathcal{P}_{2}(y, s) x^{2}+\cdots \\
& =\sum_{n=0}^{\infty} \mathcal{P}_{n}(y, s) x^{n}
\end{aligned}
$$

Take: $\quad x=\sqrt{\beta} u$
to obtain:

$$
\ddot{u}+\nu_{0}^{2} u=\nu_{0}^{2} \sum_{n=0}^{\infty} \beta^{\frac{n+3}{2}} \mathcal{P}_{n}(y, s) u^{n}
$$

A similar equation applies in the $y$-plane.
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## S4: Sources of and Forms of Perturbation Terms

Within a 2D transverse model it was shown that transverse applied magnetic components entering the equations of motion can be expanded as:

- See: S3, Transverse Particle Dynamics: 2D components axial integral 3D components
* Applied electric elds can be analogously expanded

$$
\underline{B}^{*}(\underline{z})=B_{x}^{a}(x, y)-i B_{y}^{a}(x, y)=\sum_{n=1}^{\infty} \underline{b}_{n}\left(\frac{\underline{z}}{r_{p}}\right)^{n-1}
$$

$$
\underline{b}_{n}=\text { const }(\text { complex }) \equiv \mathcal{A}_{n}-i \mathcal{B}_{n} \quad \underline{z}=x+i y \quad i=\sqrt{-1}
$$

$$
n=\text { Multipole Index } \quad r_{p}=\text { Aperture "Pipe" Radius }
$$

$\mathcal{B}_{n} \Longrightarrow$ "Normal" Multipoles
$\mathcal{A}_{n} \Longrightarrow$ "Skew" Multipoles

| Index | Name | Normal ( $\mathcal{A}_{n}=0$ ) |  | Skew ( $\mathcal{B}_{n}=0$ ) |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | $B_{x} r_{p}^{n-1} / \mathcal{B}_{n}$ | $B_{y} r_{p}^{n-1} / \mathcal{B}_{n}$ | $B_{x} r_{p}^{n-1} / \mathcal{A}_{n}$ | $B_{y} r_{p}^{n-1} / \mathcal{A}_{n}$ |
| 1 | Dipole | 0 | 1 | 1 |  |
| 2 | Quadrupole | $y$ | $x$ | $x$ | $-y$ |
| 3 | Sextupole | $2 x y$ | $x^{2}-y^{2}$ | $x^{2}-y^{2}$ | $-2 x y$ |
| 4 | Octupole | $3 x^{2} y-y^{3}$ | $x^{3}-3 x y^{2}$ | $x^{3}-3 x y^{2}$ | $-3 x^{2} y+y^{3}$ |
| 5 | Decapole | $4 x^{3} y-4 x y^{3}$ | $x^{4}-6 x^{2} y^{2}+y^{4}$ | $x^{4}-6 x^{2} y^{2}+y^{4}$ | $-4 x^{3} y+4 x y^{3}$ |
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// Reminder: Particle equations of motion from Transverse Particle
Dynamics lecture notes
Transverse particle equations of motion in explicit component form:

$$
\begin{aligned}
x^{\prime \prime}+\frac{\left(\gamma_{b} \beta_{b}\right)^{\prime}}{\left(\gamma_{b} \beta_{b}\right)} x^{\prime} & =\frac{q}{m \gamma_{b} \beta_{b}^{2} c^{2}} E_{x}^{a}-\frac{q}{m \gamma_{b} \beta_{b} c} B_{y}^{a}+\frac{q}{m \gamma_{b} \beta_{b} c} B_{z}^{a} y^{\prime} \\
& -\frac{q}{m \gamma_{b}^{3} \beta_{b}^{2} c^{2}} \frac{\partial \phi}{\partial x} \\
y^{\prime \prime}+\frac{\left(\gamma_{b} \beta_{b}\right)^{\prime}}{\left(\gamma_{b} \beta_{b}\right)} y^{\prime} & =\frac{q}{m \gamma_{b} \beta_{b}^{2} c^{2}} E_{y}^{a}+\frac{q}{m \gamma_{b} \beta_{b} c} B_{x}^{a}-\frac{q}{m \gamma_{b} \beta_{b} c} B_{z}^{a} x^{\prime} \\
& -\frac{q}{m \gamma_{b}^{3} \beta_{b}^{2} c^{2}} \frac{\partial \phi}{\partial y}
\end{aligned}
$$

Equations previously derived under assumptions:
$\rightarrow$ No bends ( xed $x-y-z$ coordinate system with no local bends)
$\rightarrow$ Paraxial equations ( $x^{\prime 2}, y^{\prime 2} \ll 1$ )
$\rightarrow$ No dispersive e ects ( $\beta_{b}$ same all particles), acceleration allowed ( $\beta_{b} \neq$ const)
$\rightarrow$ Electrostatic and leading-order (in $\beta_{b}$ ) self-magnetic interactions
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Trace back how the applied magnetic eld terms enter the $x$-plane equation of motion:

* See: S2, Transverse Particle Dynamics and reminder on next page
$\rightarrow$ Apply equation in S2 with: $\beta_{b}=$ const, $\phi \simeq$ const, $E_{x}^{a} \simeq 0, B_{z}^{a} \simeq 0$
$\rightarrow$ To include axial $\left(B_{z}^{a} \neq 0\right) \quad$ eld errors, follow a similar pattern to generalize

$$
x^{\prime \prime}=-\frac{q}{m \gamma_{b} \beta_{b} c} B_{y}^{a}
$$

Express this equation as:

$$
\begin{array}{r}
x^{\prime \prime}+\kappa_{x}(s) x=-\frac{q}{m \gamma_{b} \beta_{b} c}\left[B_{y}^{a}(x, y, s)-\left.B_{y}^{a}(x, y, s)\right|_{\operatorname{lin} x \text {-foc }}\right] \\
\text { Nonlinear focusing terms only in [] }
\end{array}
$$

* "Normal" part of linear applied magnetic eld contained in focus function $\kappa_{x}$

Compare to the form of the perturbed Hill's equation:

$$
x^{\prime \prime}+\kappa_{x} x=\mathcal{P}_{x}=\sum_{n=0}^{\infty} \mathcal{P}_{n}(y, s) x^{n}
$$

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Reduce the $x$-plane equation to our situation:

$$
40 \text { No accel } 40 \text { No E-Focus Boz }
$$

$$
x^{\prime \prime}+\frac{\left(\gamma_{b} \beta_{b}\right)^{\prime}}{\left(\gamma_{b} \beta_{b}\right)} x^{\prime}=\frac{q}{m \gamma_{b} \beta_{b}^{2} c^{2}} E_{x}^{a}-\frac{q}{m \gamma_{b} \beta_{b} c} B_{y}^{a}+\frac{q}{m \gamma_{b} \beta_{b} c} B_{z}^{a} y^{\prime}
$$

40 No Space-Charge

$$
-\frac{q}{m \gamma_{b}^{3} \beta_{b}^{2} c^{2}} \frac{\partial \phi}{\partial x}
$$

Giving the equation we are analyzing:

$$
\Longrightarrow \quad x^{\prime \prime}=-\frac{q}{m \gamma_{b} \beta_{b} c} B_{y}^{a}
$$

## Gives:

$$
\Longrightarrow \quad \mathcal{P}_{x}=-\frac{q}{m \gamma_{b} \beta_{b} c}\left[B_{y}^{a}-\left.B_{y}^{a}\right|_{\operatorname{lin} x \text {-foc }}\right]
$$

where the $y$ - eld components can be obtained from the multipole expansion as:

$$
\begin{aligned}
& B_{y}^{a}=-\operatorname{Im}\left[\underline{B}^{*}\right] \\
& \left.B_{y}^{a}\right|_{\text {lin } x \text {-focus }}=-\operatorname{Im}\left[\left.\underline{B}^{*}\right|_{n=1 \text { term }}\right]
\end{aligned} \quad \underline{B}^{*}=\sum_{n=1}^{\infty} \underline{b}_{n}\left(\frac{x+i y}{r_{p}}\right)^{n-1}
$$

* Use multipole eld components of magnets to obtain explicit form of eld component perturbations consistent with the Maxwell equations
$\rightarrow$ Need to subtract o design component of linear led from $\mathcal{P}_{x}$ perturbation term since it is included in $\kappa_{x}$
$\rightarrow$ Similar steps employed to identify $y$-plane perturbation terms, perturbations from axial eld components, and perturbations for applied electric eld components


## S5: Solution of the Perturbed Hill's Equation: Resonances

Analyze the solution of the perturbed orbit equation:

$$
\ddot{u}+\nu_{0}^{2} u=\nu_{0}^{2} \sum_{n=0}^{\infty} \beta^{\frac{n+3}{2}} \mathcal{P}_{n}(y, s) u^{n}
$$

derived in S4.
To more simply illustrate resonances, we analyze motion in the $x$-plane with:

$$
y(s) \equiv 0
$$

* Essential character of general analysis illustrated most simply in one plane
$\rightarrow$ Can generalize by expanding $\mathcal{P}_{n}(y, s)$ in a power series in $y$ and generalizing notation to distinguish between Floquet coordinates in the $x$ - and $y$-planes
- Results in coupled $x$ - and $y$-equations of motion

Caution: Multipole index $n$ and power series index $n$ in $\mathcal{P}_{x}$ expansion not the same (notational overuse: wanted analogous symbol)

- Multipole Expansion for $B_{x}^{a}, B_{y}^{a}$ :

$$
\begin{array}{llll}
n=1 & \text { Dipole } & n=3 & \text { Sextupole } \\
n=2 & \text { Ouadrunole } & n= &
\end{array}
$$

- Power Series Expansion for $\mathcal{P}_{x}$ :

$$
\begin{array}{lll}
x \text {-plane Motion }(y=0) & x-y \text { plane motion } \\
n=0 & \text { Dipole } & \text { Depends on form of } y \text {-coupling } \\
n=1 & \text { Quadrupole } & \\
n=2 & \text { Sextupole } &
\end{array}
$$

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Each n-labeled perturbation expansion coe cient is periodic with period of the ring circumference (random perturbations) or lattice period (systematic):

$$
\begin{aligned}
& L_{p}=\text { Lattice Period } \\
& \mathcal{C}=\mathcal{N} L_{p}=\text { Ring Circumference } \\
& \beta\left(s+L_{p}\right)=\beta(s) \\
& \beta\left(s+\mathcal{N} L_{p}\right)=\beta(s)
\end{aligned}
$$

Random Perturbation:

$$
\begin{aligned}
& \mathcal{P}_{n}\left(y, s+\mathcal{N} L_{p}\right)=\mathcal{P}_{n}(y, s) \\
& \Longrightarrow \beta^{\frac{n+3}{2}}\left(s+\mathcal{N} L_{p}\right) \mathcal{P}_{n}\left(y, s+\mathcal{N} L_{p}\right)=\beta^{\frac{n+3}{2}}(s) \mathcal{P}_{n}(y, s)
\end{aligned}
$$

Systematic Perturbation:

$$
\begin{aligned}
& \mathcal{P}_{n}\left(y, s+L_{p}\right)=\mathcal{P}_{n}(y, s) \\
& \Longrightarrow \beta^{\frac{n+3}{2}}\left(s+L_{p}\right) \mathcal{P}_{n}\left(y, s+L_{p}\right)=\beta^{\frac{n+3}{2}}(s) \mathcal{P}_{n}(y, s)
\end{aligned}
$$

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Expand each $n$-labeled perturbation expansion coe cient in a Fourier series as:

$$
\begin{aligned}
& \beta^{\frac{n+3}{2}} \mathcal{P}_{n}(y=0, s)=\sum_{k=-\infty}^{k=\infty} C_{n, k} e^{i k p \varphi} \\
& i \equiv \sqrt{-1} \\
& C_{n, k}=\int_{-\pi / p}^{1,} \begin{array}{ll}
\text { Random perturbation } \\
\text { (once per lap in ring) }
\end{array} \\
& \mathcal{N}, \\
& \quad \begin{array}{l}
\text { Systematic perturbation } \\
\text { (every lattice period) }
\end{array} \\
& s=s(\varphi) \quad \varphi=\int_{s_{0}}^{s} \frac{d \varphi}{C_{0}} \frac{d \tilde{s}}{\beta(\tilde{s})}
\end{aligned}
$$

- Can apply to Rings for random perturbations (with $p=1$ )
or systematic perturbations (with $p=\mathcal{N}$ )
$\rightarrow$ Can apply to linacs for periodic perturbations (every lattice period) with $p=1$
- Does not apply to random perturbations in a linac
- In linac random perturbations will vary every lattice period and drive random walk type e ects but not resonances
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To obtain the perturbed equation of motion, the unperturbed solution $u_{0}$ is inserted on the RHS terms
$\rightarrow$ Gives simple harmonic oscillator equation with driving terms
Solution of the unperturbed orbit is simply expressed as:

$$
\left.\begin{array}{c}
u_{0}=u_{0 i} \cos \left(\nu_{0} \varphi+\varphi_{i}\right)=u_{0 i} \frac{e^{i\left(\nu_{0} \varphi+\varphi_{i}\right)}+e^{-i\left(\nu_{0} \varphi+\varphi_{i}\right)}}{2} \\
u_{0 i}=\text { const } \\
\varphi_{i}=\text { const }
\end{array}\right\} \begin{aligned}
& \text { Set by particle initial conditions: } \\
& x\left(s_{i}\right)=x_{i}, \quad x^{\prime}\left(s_{i}\right)=x_{i}^{\prime}
\end{aligned}
$$

$$
\begin{aligned}
& \text { Then binomial expand: } \\
& \qquad \begin{aligned}
u_{0}^{n} & =u_{0 i}^{n}\left(\frac{e^{i\left(\nu_{0} \varphi+\varphi_{i}\right)}+e^{-i\left(\nu_{0} \varphi+\varphi_{i}\right)}}{2}\right)^{n} \\
& =\frac{u_{o i}^{n}}{2^{n}} \sum_{m=0}^{n}\binom{n}{m} e^{i(n-m)\left(\nu_{0} \varphi+\varphi_{i}\right)} e^{-i m\left(\nu_{0} \varphi+\varphi_{i}\right)} \\
& =\frac{u_{o i}^{n}}{2^{n}} \sum_{m=0}^{n}\binom{n}{m} e^{i(n-2 m) \nu_{0} \varphi} e^{i(n-2 m) \varphi_{i}}
\end{aligned}
\end{aligned}
$$

where $\quad\binom{n}{m} \equiv \frac{n!}{m!(n-m)!} \quad$ is a binomial coe cient
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The perturbed equation of motion becomes:

$$
\ddot{u}+\nu_{0}^{2} u=\nu_{0}^{2} \sum_{n=0}^{\infty} \sum_{k=-\infty}^{k=\infty} C_{n, k} e^{i k p \varphi} u^{n}
$$

Expand the solution as:

$$
\begin{array}{ll}
u=u_{0}+\delta u & u_{0}=\text { unperturbed solution } \\
\delta u=\text { perturbation due to errors }
\end{array}
$$

where $u_{0}$ is the solution to the simple harmonic oscillator equation in the absence of perturbations:

$$
\ddot{u}_{0}+\nu_{0}^{2} u_{0}=0 \quad \begin{aligned}
& \text { Unperturbed } \\
& \text { equation of motion }
\end{aligned}
$$

Assume small-amplitude perturbations so that

$$
\left|u_{0}\right| \gg|\delta u|
$$

Then to leading order, the equation of motion for $\delta u$ is:

| $\ddot{\delta u}+\nu_{0}^{2} \delta u \simeq \nu_{0}^{2} \sum_{n=0}^{\infty} \sum_{k=-\infty}^{k=\infty} C_{n, k} e^{i p k \varphi} u_{0}^{n} \quad$Perturbed <br> equation of motion |
| :--- | :--- |
| Particle Resonances |

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Using this expansion the linearized perturbed equation of motion becomes:

$$
\ddot{\delta u}+\nu_{0}^{2} \delta u \simeq \nu_{0}^{2} \sum_{n=0}^{\infty} \sum_{k=-\infty}^{k=\infty} \sum_{m=0}^{n}\binom{n}{m} \frac{C_{n, k}}{2^{n}} e^{i\left[(n-2 m) \nu_{0}+p k\right] \varphi} e^{i(n-2 m) \varphi_{i}}
$$

The solution for $\delta u$ can be expanded as:

$$
\begin{aligned}
& \delta u= \delta u_{h}+ \\
& \delta u_{p} \\
& \delta u_{h}= \\
& \text { homogenous solution } \\
& \quad \text { General solution to: } \quad \ddot{\delta} u_{h}+\nu_{0}^{2} \delta u_{h}=0 \\
& \delta u_{p}= \text { particular solution } \\
& \text { Any solution with: } \quad \delta u \rightarrow \delta u_{p}
\end{aligned}
$$

$\rightarrow$ Can drop homogeneous solution because it can be absorbed in unperturbed solution $u_{0}$

- Exception: some classes of linear amplitude errors in adjusting magnets
* Only a particular solution need be found, take:
$\delta u=\delta u_{p}$

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$$
\ddot{\delta u}+\nu_{0}^{2} \delta u \simeq \nu_{0}^{2} \sum_{n=0}^{\infty} \sum_{k=-\infty}^{k=\infty} \sum_{m=0}^{n}\binom{n}{m} \frac{C_{n, k}}{2^{n}} e^{i\left[(n-2 m) \nu_{0}+p k\right] \varphi} e^{i(n-2 m) \varphi_{i}}
$$

Equation describes a driven simple harmonic oscillator with periodic driving terms on the RHS:
$\rightarrow$ Homework problem reviews that solution of such an equation will be unstable when the driving term has a frequency component equal to the restoring term

- Resonant exchange and amplitude grows linearly (not exponential!) in $\varphi$
- Parameters meeting resonance condition will lead to instabilities with particle oscillation amplitude growing in $\varphi(s)$
Resonances occur when:

$$
(n-2 m) \nu_{0}+p k= \pm \nu_{0}
$$

is satis ed for the operating tune $\nu_{0}$ and some values of:

$$
n=0,1,2, \cdots \quad m=0,1,2, \cdots, n
$$

$k=-\infty, \cdots,-1,0,1, \cdots, \infty$
$p \equiv \begin{cases}1, & \text { Random perturbation } \\ \mathcal{N}, & \text { Systematic perturbation }\end{cases}$
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## S6: Machine Operating Points:

Tune Restrictions Resulting from Resonances
Examine situations where the $x$-plane motion resonance condition:

$$
(n-2 m) \nu_{0}+p k= \pm \nu_{0}
$$

is satis ed for the operating tune $\nu_{0}$ and some values of:

| Multipole Order Index: <br> $n=0,1,2, \cdots$ | Specify to understand class of <br> perturbations |
| :--- | :--- |
| Particle Binomial Expansion Index:  <br> $m=0,1,2, \cdots, n$  | Linear superposition for <br> multiple perturbations |

Periodicity Fourier Series Expansion Index:

$$
k=-\infty, \cdots,-1,0,1, \cdots, \infty
$$

Perturbation Symmetry Factor:

$$
p \equiv \begin{cases}1, & \text { Random perturbation } \\ \mathcal{N}, & \text { Systematic perturbation }\end{cases}
$$

Resonances can be analyzed one at a time using linear superposition
$\rightarrow$ Analysis valid for small-amplitudes
Analyze resonance possibilities starting with index $n<==>$ Multipole Order SM Lund, USPAS, 2020

If growth rate is su ciently large, machine operating points satisfying the resonance condition will be problematic since particles will be lost (scraped) by the machine aperture due to increasing oscillation amplitude:
$\rightarrow$ Machine operating tune ( $\nu_{0}$ ) can be adjusted to avoid
$\rightarrow$ Perturbation can be actively corrected to reduce amplitude of driving term
Low order resonance terms with smaller $n, k, m$ magnitudes are expected to be more dangerous because:
$\rightarrow$ Less likely to be washed out by e ects not included in model
$\rightarrow$ Amplitude coe cients expected to be stronger
More detailed theories consider coherence length, nite amplitude, and nonlinear term e ects. Such treatments and numerical analysis concretely motivate importance/strength of terms. A standard reference on analytic theory is:
$\rightarrow$ Kolomenskii and Lebedev, Theory of Circular Accelerators, North-Holland (1966)
We only consider lowest order e ects in these notes.
In the next section we will examine how resonances restrict possible machine operating parameters.
$\rightarrow$ After establishing clear picture of e ect on single particle orbits we will then add space-charge
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$$
\begin{aligned}
& n=0, \underline{\text { Dipole Perturbations: }} \quad(n-2 m) \nu_{0}+p k= \pm \nu_{0} \\
& n=0, \Longrightarrow m=0
\end{aligned}
$$

and the resonance condition gives a single constraint:

$$
\begin{aligned}
\nu_{0}= \pm p k \quad p k=\text { integer } & k=-\infty, \cdots,-1,0,1, \cdots, \infty \\
& p= \begin{cases}1, & \text { Random perturbation } \\
\mathcal{N}, & \text { Systematic perturbation }\end{cases}
\end{aligned}
$$

Therefore, to avoid dipole resonances integer tunes operating points not allowed:

| $p=1$ | Random Perturbation | $\nu_{0} \neq 1,2,3, \cdots$ |
| :--- | :--- | :--- |
| $p=\mathcal{N}$ | Systematic Perturbation | $\nu_{0} \neq \mathcal{N}, 2 \mathcal{N}, 3 \mathcal{N}, \cdots$ |

$\rightarrow$ Systematic errors are signi cantly less restrictive on machine operating points for large $\mathcal{N}$

- Illustrates why high symmetry is desirable
- Racetracks with $\mathcal{N}=2$ can be problem
$\rightarrow$ Multiply random perturbation tune restrictions by $\mathcal{N}$ to obtain the systematic perturbation case


## Interpretation of result:

Consider a ring with a single (random) dipole error along the reference path of the ring:


If the particle is oscillating with integer tune, then the particle experiences the dipole error on each lap in the same oscillation phase and the trajectory will "walk-o " on a lap-to-lap basis in phase-space
$\rightarrow$ With nite machine aperture the particle will be scraped/lost


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## Interpretation of result (new restrictions):

For a single (random) quadrupole error along the azimuth of a ring, a similar qualitative argument as presented in the dipole resonance case leads one to conclude that if a particle oscillates with $1 / 2$ integer tune, then the orbit can "walk
o " on a lap-to-lap basis in phase-space:



$$
\begin{aligned}
& n=1, \text { Quadrupole Perturbations: } \quad(n-2 m) \nu_{0}+p k= \pm \nu_{0} \\
& n=1, \Longrightarrow m=0,1
\end{aligned}
$$

and the resonance conditions give:

$$
\begin{array}{rr}
n=1, m=0: \\
n=1, m=1: & \nu_{0}+p k= \pm \nu_{0}
\end{array} \quad \Longrightarrow \quad \begin{aligned}
& \text { Give two cases: } \\
& \\
& n=p k=0, \nu_{0}= \pm \frac{p k}{2}
\end{aligned}
$$

Implications of two cases: Can be treated by "renormalizing" oscillator

$$
\begin{aligned}
& \text { 1) } p k=0 \Rightarrow k=0 \quad \text { focusing strength: need not be considered } \\
& \text { 2) } \nu_{0}= \pm \frac{p k}{2} \Rightarrow \nu_{0}=\frac{|p k|}{2} \quad \ddot{u}+\nu_{0}^{2} u=\nu_{0}^{2} C_{1,0} u
\end{aligned}
$$

Therefore, to avoid quadrupole resonances, the following tune operating points are not allowed:

$$
\begin{aligned}
\nu_{0} \neq \frac{|p k|}{2} & p
\end{aligned} \quad\left\{\begin{array}{ll}
1, & \text { Random perturbation } \\
\mathcal{N}, & \text { Systematic perturbation }
\end{array}\right\}
$$

$*$ New restriction: tunes cannot be half-integer values
Integers also restricted for $p=1$ random, but redundant with dipole case
$\rightarrow$ Some large integers restricted for $p=\mathcal{N}$ systematic perturbations SM Lund, USPAS, 2020

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$$
n=2, \text { Sextupole Perturbations: } \quad(n-2 m) \nu_{0}+p k= \pm \nu_{0}
$$

$$
n=2, \quad \Longrightarrow \quad m=0,1,2
$$

and the resonance conditions give the three constraints below:

$$
\begin{array}{rrr}
n=2, m=0: & 2 \nu_{0}+p k= \pm \nu_{0} \\
n=2, m=1: & p k= \pm \nu_{0} \\
n=2, m=2: & -2 \nu_{0}+p k= \pm \nu_{0}
\end{array}
$$

Therefore, to avoid sextupole resonances, the following tunes are not allowed:

$$
\begin{aligned}
\nu_{0} \neq\left\{\begin{array}{lll}
|p k| & \text { integer } & p= \begin{cases}1, & \text { Random perturbation } \\
\mathcal{N}, & \text { Systematic perturbation } \\
|p k| / 3 & \text { third-integer }\end{cases} \\
k=-\infty, \cdots,-1,0,1, \cdots, \infty
\end{array}\right. \\
\hline k=\cdots, \cdots, \cdots
\end{aligned}
$$

Integer restrictions already obtained for dipole perturbations
$\rightarrow 1 / 3$-integer restriction new

Higher-order $(n>2)$ cases analyzed analogously: $(n-2 m) \nu_{0}+p k= \pm \nu_{0}$
$\rightarrow$ Produce more constraints but expected to be weaker as order increases
$\rightarrow$ Will always generate a new constraint $\nu_{0} \neq \frac{|p k|}{n+1}$

## Restrictions on machine operating points

Tune restrictions are typically plotted in $\nu_{0 x}-\nu_{0 y}$ space order-by-order up to a max order value to nd allowed tunes where the machine can safely operate
$\rightarrow$ Often $3^{\text {rd }}$ order is chosen as a maximum to avoid
$\rightarrow$ Cases for random $(p=1)$ and systematic $(p=\mathcal{N})$ perturbations considered
Machine operating points chosen as far as possible from low order resonance lines

## Random Perturbations

Systematic Perturbations
$p=1 \quad$ Adapted from Wiedemann $\quad p=\mathcal{N}=4$



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## General form of resonance condition

The general resonance condition (all $n$-values) for $x$-plane motion can be summarized as:

$$
M \nu_{0}=N \quad \begin{aligned}
M, N & =\text { Integers of same sign } \\
|M| & =\text { "Order" of resonance }
\end{aligned}
$$

$\rightarrow$ Higher order numbers $M$ are typically less dangerous

- Longer coherence length for validity of theory: e ects not included can
"wash-out" the resonance
- Coe cients generally smaller

Particle motion is not (measure zero) really restricted to the $x$-plane, and a more complete analysis taking into account coupled $x$ - and $y$-plane motion shows that the generalized resonance condition is:
$\rightarrow$ Place unperturbed y-orbit in rhs perturbation term, then leading-order expand analogously to x-case to obtain additional driving terms

$$
\begin{array}{cl}
M_{x} \nu_{0 x}+M_{y} \nu_{0 y}=N & M_{x}, M_{y}, N=\text { Integers of same sign } \\
\nu_{0 x}=x \text {-plane tune } & \left|M_{x}\right|+\left|M_{y}\right|=\text { "Order" of resonance } \\
\nu_{0 y}=y \text {-plane tune } &
\end{array}
$$

$\rightarrow$ Lower order resonances are more dangerous analogously to $x$-case SM Lund, USPAS, 2020

## Discussion: Restrictions on machine operating points

## Random Errors:

* Errors always present and give low-order resonances
* Usually have weak amplitude coe cients
- Can be corrected/compensated to reduce e ects


## Systematic Errors:

$\rightarrow$ Lead to higher-order resonances for large $\mathcal{N}$ and a lower density of resonance lines (see plots on previous slide comparing the equal boxed red areas)

Large symmetric rings with high $\mathcal{N}$ values have less operating
restrictions from systematic errors
Practical issues such as construction cost and getting the beam
into and out of the ring can lead to smaller $\mathcal{N}$ values (racetrack lattice)
$\rightarrow$ BUT systematic error Amplitude coe cients can be large
-Systematic e ects accumulate in amplitude period by period
Resonances beyond $3^{\text {rd }}$ order rarely need be considered

* E ects outside of model assumed tend to wash-out higher order resonances

More detailed treatments calculate amplitudes/strengths of resonant terms
$\rightarrow$ See accelerator physics references:
Further info: Wiedemann, Particle Accelerator Physics (2007)
Amplitudes/Strengths: Kolomenskii and Lebedev, $\begin{aligned} & \text { Theory of Circular } \\ & \text { Particle Resonances }\end{aligned}$ Lund, USPAS, 2020

## S7: Space-Charge E ects on Particle Resonances

## S7A: Introduction

Ring operating points are chosen to be far from low-order particle resonance lines in $x-y$ tune space. Processes that act to shift particle resonances closer towards the low-order lines can prove problematic:
$\rightarrow$ Oscillation amplitudes increase (spoiling beam quality and control)
$\star$ Particles can be lost
Tune shift limits of machine operation are often named "Laslett Limits" in honor of Jackson Laslett who rst calculated tune shift limits for various processes:
$\rightarrow$ Image charges
$\rightarrow$ Image currents

- Internal beam self- eld
* 

Processes shifting resonances can be grouped into two broad categories:

| Coherent | Same for every particle in distribution  <br>  $\rightarrow$ Usually most dangerous: full beam resonant |
| :--- | ---: |
| Incoherent | Dierent for particles in separate parts of the distribution <br>  |
|  | $\rightarrow$ Usually less dangerous: only e ects part of beam |

* Usually less dangerous: only e ects part of beam

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Here we will analyze Laslett tune shift limits induced by coherent space-charge taking a KV distribution model for linear space-charge
$\rightarrow$ In KV model, space-charge forces interior to the beam are coherent because all particles have the same depressed tune
$\rightarrow$ Warning: some call space-charge tune shifts even within the KV model as "incoherent" used as a name taken to be synonymous with space-charge. This may not make good logical sense.

We will not analyze Laslett limits for other processes in these lectures. But the logical procedure is similar to the space-charge case.

## S7B: Laslett Space-Charge Limit

Laslett rst obtained a space-charge limit for rings by assuming that the beam space-charge is uniformly distributed as in a KV model and thereby acts as a coherent shift to previously derived resonance conditions. Denote:
$\nu_{0 x} \equiv x$-tune (bare) in absence of space-charge $\nu_{x} \equiv x$-tune (depressed) with uniform density beam

$$
\Delta \nu_{x} \equiv \nu_{0 x}-\nu_{x}=\text { Space-charge tune shift } \quad \Delta \nu_{x} \geq 0
$$

Assume that dipole (integer) and quadrupole (half-integer) tunes only need be excluded when space-charge e ects are included.
$\rightarrow$ Space-charge likely induces more washing-out of higher-order resonances
If the bare tune operating point is chosen as far as possible from $1 / 2$-integer resonance lines, the maximum space-charge induced tune shift allowed is $1 / 4-$ integer, giving:

$$
\left.\Delta \nu_{x}\right|_{\max }=\frac{1}{4} \Longrightarrow \begin{gathered}
\text { Establishes maximum current } \\
\text { (use KV results in lectures on }
\end{gathered}
$$

Transverse Equilibrium Distributions)
$\rightarrow$ Analogous equation applies in the $y$-plane

- Identical restriction in lattices with equal $x$ - and $y$-focusing strengths

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Consider a symmetric ring (not race track for simple arguments) with

$$
\begin{aligned}
\mathcal{N} & =\text { Number Lattice Periods } \\
L_{p} & =\text { Lattice Period } \\
\sigma_{0} & =\text { Phase advance in } x, y \text {-directions }
\end{aligned}
$$

Gives bare (undepressed) tunes:

$$
\Longrightarrow \quad \nu_{0 x}=\nu_{0 y} \equiv \nu_{0}=\mathcal{N} \frac{\sigma_{0}}{2 \pi}
$$

De ning the depressed tune in the presence of KV model space-charge analogously to the bare tunes gives:

* Drop " 0 " from subscript to indicate in the presence of space-charge

$$
\nu_{x}=\nu_{y} \equiv \nu=\mathcal{N} \frac{\sigma}{2 \pi}
$$

Then the allowed space-charge depression $\sigma / \sigma_{0}$ for $\delta \nu=\nu_{0}-\nu=1 / 4$ is:

$$
\delta \nu=\frac{\mathcal{N} \sigma_{0}}{2 \pi}-\frac{\left.\mathcal{N} \sigma\right|_{\min }}{2 \pi}=1 /\left.4 \quad \Longrightarrow \quad \frac{\sigma}{\sigma_{0}}\right|_{\min }=1-\frac{\pi / 2}{\mathcal{N} \sigma_{0}}
$$

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## Estimate of Maximum Perveance/Current Allowed by Laslett Limit:

## Simple Continuous Focusing Estimate

Model the focusing as continuous and assume an unbunched, transverse matched

$$
\begin{array}{ll}
\text { KV distribution with: } \\
\qquad \begin{aligned}
\kappa_{x} & =\kappa_{y}=k_{\beta 0}^{2}=\text { const } \\
\varepsilon_{x} & =\varepsilon_{y} \equiv \varepsilon=\text { const } \\
Q & =\frac{q \lambda}{2 \pi \epsilon_{0} m \gamma_{b}^{3} \beta_{b}^{2} c^{2}}=\text { const }
\end{aligned} & \text { Focusing Strength } \quad k_{\beta 0}=\frac{\sigma_{0}}{L_{p}}=\frac{2 \pi \nu_{0}}{\mathcal{N} L_{p}} \\
& \text { Dimsensionless Perveance }
\end{array}
$$

The matched envelope equation gives:

$$
r_{x}=r_{y}=r_{b}=\mathrm{const}
$$

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$$
r_{b}^{\prime \prime}+k_{\beta 0}^{2} r_{b}-\frac{Q}{r_{b}}-\frac{\varepsilon^{2}}{r_{b}^{3}}=0
$$

$$
\Longrightarrow \quad r_{b}^{2}=\frac{Q+\sqrt{4 k_{\beta 0}^{2} \varepsilon^{2}+Q^{2}}}{2 k_{\beta 0}^{2}}
$$

Use this constraint in formulas for the depressed phase advance to connect to the tune depression limit.

Depressed phase advance per lattice period can then be calculated from formulas in lectures on Transverse Equilibrium Distributions as:

Two forms equivalent from envelope $\sigma=k_{\beta} L_{p}=\sqrt{k_{\beta 0}^{2}-\frac{Q}{r_{b}^{2}}} L_{p}=\int_{s_{i}}^{s_{i}+L_{p}} \frac{d s}{\beta}=\varepsilon \int_{s_{i}}^{s_{i}+L_{p}} \frac{d s}{r_{b}^{2}}$ equation


$$
\nu_{0}=\mathcal{N} \frac{\sigma_{0}}{2 \pi}=\frac{\mathcal{N} k_{\beta 0} L_{p}}{2 \pi}
$$

and previous formulas for $Q /\left(k_{\beta 0}^{2} r_{b}^{2}\right)$ gives:

$$
\nu=\nu_{0} \sqrt{1-\frac{2 Q}{Q+\sqrt{\frac{16 \pi^{2} \nu_{Q}^{2} \varepsilon^{2}}{\mathcal{N}^{2} L_{p}^{2}}+Q^{2}}}}
$$

Setting the phase shift to the Laslett current limit value

$$
\left.\nu\right|_{Q=Q_{\max }}=\nu_{0}-\frac{1}{4}
$$

gives a constraint for the maximum value of $Q=Q_{\max }$ to avoid
$1 / 2$-integer resonances:

$$
\frac{2 Q_{\max }}{Q_{\max }+\sqrt{\frac{16 \pi^{2} \nu_{0}^{2} \varepsilon^{2}}{\mathcal{N}^{2} L_{p}^{2}}+Q_{\max }^{2}}}=1-\left(\frac{\nu_{0}-1 / 4}{\nu_{0}}\right)^{2}=\frac{1}{2 \nu_{0}^{2}}\left(\nu_{0}-1 / 8\right)
$$

This can be arraigned into a quadratic equation for $Q_{\max }$ and solved to show that the Laslett "current" limit expressed in terms of max transportable perveance:

To express in terms

$$
\begin{aligned}
Q<Q_{\max } & =\frac{\pi \varepsilon}{\mathcal{N} L_{p}}\left(\frac{\nu_{0}-1 / 8}{\nu_{0}}\right) \frac{1}{\sqrt{1-\frac{1}{2 \nu_{0}}\left(\frac{\nu_{0}-1 / 8}{\nu_{0}}\right)}} \\
& \simeq \frac{\pi \varepsilon}{\mathcal{N} L_{p}}\left(1+\frac{1}{8 \nu_{0}}+\operatorname{Order}\left(1 / \nu_{0}^{2}\right)\right)
\end{aligned}
$$

$$
\begin{aligned}
Q & =\frac{q \lambda}{2 \pi \epsilon_{0} m \gamma_{b}^{3} \beta_{b}^{2} c^{2}} \\
& =\frac{q I}{2 \pi \epsilon_{0} m \gamma_{b}^{3} \beta_{b} c}
\end{aligned}
$$

// Example: Take (typical synchrotron numbers, represents peak charge in rf bunch) $\mathcal{N} L_{p}=\mathcal{C}=$ Ring Circumfrance $\sim 300 \mathrm{~m}$

$$
\begin{array}{ccc}
\varepsilon \sim 50 \mathrm{~mm}-\mathrm{mrad} & \Longrightarrow & Q<Q_{\max } \simeq \frac{\pi \varepsilon}{\mathcal{C}} \simeq 5 \times 10^{-7} \\
\text { Neglect } 1 / \nu_{0} \text { term } & \text { Not a lot of charge } \ldots . \quad \text { // } \\
\text { SM Lund, USPAS, 2020 } & \text { Particle Resonances }
\end{array}
$$

## Discussion:

Laslett limit may be overly restrictive:
$\rightarrow$ KV model assumes all particles in beam have the same tune

- Signi cant spectrum of particle tunes likely in real beam

Particularly if space-charge strong: see Transverse Equilibrium Dists, S7

$$
I=\text { Beam Current }
$$

- No equilibrium beam: core oscillates and space-charge may act
- No equilibrium beam: core oscillates and space
incoherently to e ectively wash-out resonances

For strong e:
$\rightarrow$ Frequency spread large and KV approx bad

- Does not work in spite of beam density being near uniform density for smooth distribution
 For weak space-charge $\rightarrow$ Frequency spread small and KV approx good
$\rightarrow$ Works in spite of beam density being far from uniform density for smooth distribution
- Simulations suggest Laslett limit poses little issues over 10s - 100s of laps in rings (Small Recirculator, LLNL) and in fast bunch compressions in rings

Longer simulations di cult to resolve: see USPAS Simulation Course
$\rightarrow$ Future experiments can hopefully address this issue

- University of Maryland electron ring?

SM Lund, USPAS, 2020 Particle Resonances

Frequency distribution for an idealized 1D thermal equilibrium beam suggest signi cant deviations from KV coherent picture when space-charge intensity is high.


Oscillation Frequency, $k_{\beta} / k_{\beta 0}$

|  |  | Statistical Measures |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
|  |  | Mean: | RMS: $\overline{\sigma_{F}}=$ | Width: | Relative Width: |
| $\sigma / \sigma_{0}$ | $\Delta$ | $\mu_{F}=\overline{k_{\beta}} / k_{\beta 0}$ | $\sqrt{\overline{k_{\beta}^{2}}-\overline{k_{\beta}}} / k_{\beta 0}$ | $F_{w}=2 \sqrt{3} \sigma_{F}$ | $F_{w} / \mu_{F}$ |
| 0.9 | 2.879 | 0.886 | 0.0176 | 0.0610 | 0.0689 |
| 0.8 | 1.093 | 0.774 | 0.0354 | 0.123 | 0.159 |
| 0.7 | 0.5181 | 0.663 | 0.0531 | 0.184 | 0.277 |
| 0.6 | 0.2500 | 0.557 | 0.0696 | 0.241 | 0.433 |
| 0.5 | 0.1097 | 0.456 | 0.0833 | 0.289 | 0.634 |
| 0.4 | $3.780 \times 10^{-2}$ | 0.361 | 0.0915 | 0.317 | 0.878 |
| 0.3 | $7.562 \times 10^{-3}$ | 0.274 | 0.0898 | 0.311 | 1.14 |
| 0.2 | $3.649 \times 10^{-4}$ | 0.190 | 0.0750 | 0.260 | 1.37 |
| 0.1 | $5.522 \times 10^{-8}$ | 0.102 | 0.0465 | 0.161 | 1.58 |
| SM Lund, USPAS, 2020 |  | Particle Resonances |  |  |  |

Discussion Continued:
$\rightarrow$ Even if internal resonances in the core of the beam are washed out due to nonlinear space-charge at high intensity, centroid resonances may still behave more as a single particle (see notes on Transverse Centroid and Envelope Descriptions of Beam Evolution) to limit beam control.

- Steering and correction can mitigate low order centroid instabilities
- Centroid will also have (likely weak if steering used)
image charge correction to the tune
$\rightarrow$ Caution: Terminology can be very bad/confusing on topic. Some researchers:
- Call KV Laslett space-charge shift an "incoherent tune shift"
limit in spite of it being (KV) coherent
- Call anything space-charge related "incoherent" regardless of mode
- Call beam transport near the KV Laslett space-charge shift limit a
"space charge dominated beam" even though space-charge defocusing
likely is only a small fraction of the applied focusing
More research on this topic is needed
* Higher intensities open new applications for energy and material processing
$\rightarrow$ Many possibilities to extend operating range of existing machines and make new use of developed technology
$\rightarrow$ Good area for graduate thesis projects!
SM Lund, USPAS, 2020


## Laslett Space Charge Limit for an Elliptical KV Beam

For more info, see material in the G. Franchetti lecture from the 2014 CERN Accelerator School in the reference below:
http://cas.web.cern.ch/sites/cas.web.cern.ch/ les/lectures/prague-2014/franchettisc.pdf

## S8: Limits Induced by Space-Charge Collective Modes

Add in future edition of notes here or in kinetic theory:

Review simple 1D theory results of Sacherer and implications for rings

## Corrections and suggestions for improvements welcome!

These notes will be corrected and expanded for reference and for use in future editions of US Particle Accelerator School (USPAS) and Michigan State University (MSU) courses. Contact:

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Please provide corrections with respect to the present archived version at:
https://people.nscl.msu.edu/~lund/uspas/bpisc_2020
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## References: Continued (2):

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A. Dragt, "Lectures on Nonlinear Orbit Dynamics," in Physics of High Energy Accelerators, edited by R.A. Carrigan, F.R. Hudson, and M. Month (AIP Conf. Proc. No. 87, 1982) p. 147
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F. Sacherer, Transverse Space-Charge E ects in Circular Accelerators, Univ. of California Berkeley, Ph.D Thesis (1968)
A. A. Kolomenskii and A. N. Lebedev, Theory of Circular Accelerators, NorthHolland (1966)

## References: For more information see:

These course notes are posted with updates, corrections, and supplemental material at: https://people.nscl.msu.edu/~lund/uspas/bpisc_2020
Materials associated with previous and related versions of this course are archived at: JJ Barnard and SM Lund, Beam Physics with Intense Space-Charge, USPAS: https://people.nscl.msu.edu/~lund/uspas/bpisc_2017 2017 Version https://people.nscl.msu.edu/~lund/uspas/bpisc_2015 2015 Version http://hifweb.lbl.gov/USPAS_2011 2011 Lecture Notes + Info http://uspas.fnal.gov/programs/past-programs.shtml (2008, 2006, 2004)
JJ Barnard and SM Lund, Interaction of Intense Charged Particle Beams with Electric and Magnetic Fields, UC Berkeley, Nuclear Engineering NE290H http://hifweb.lbl.gov/NE290H 2009 Lecture Notes + Info

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