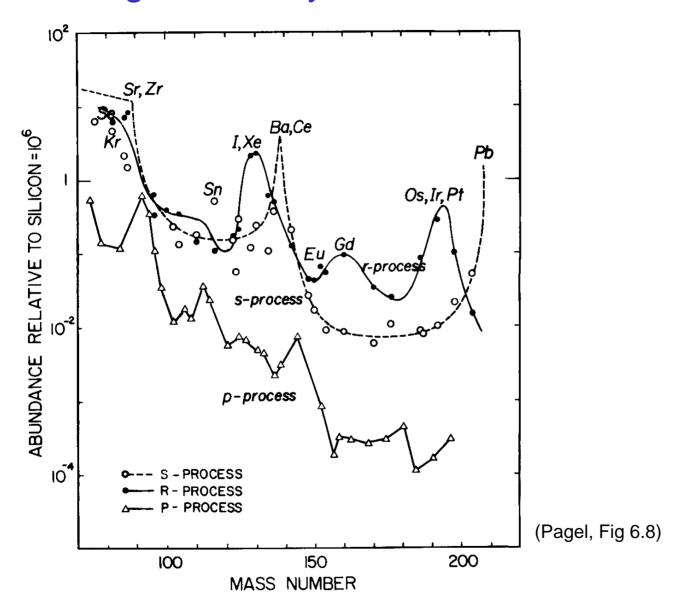
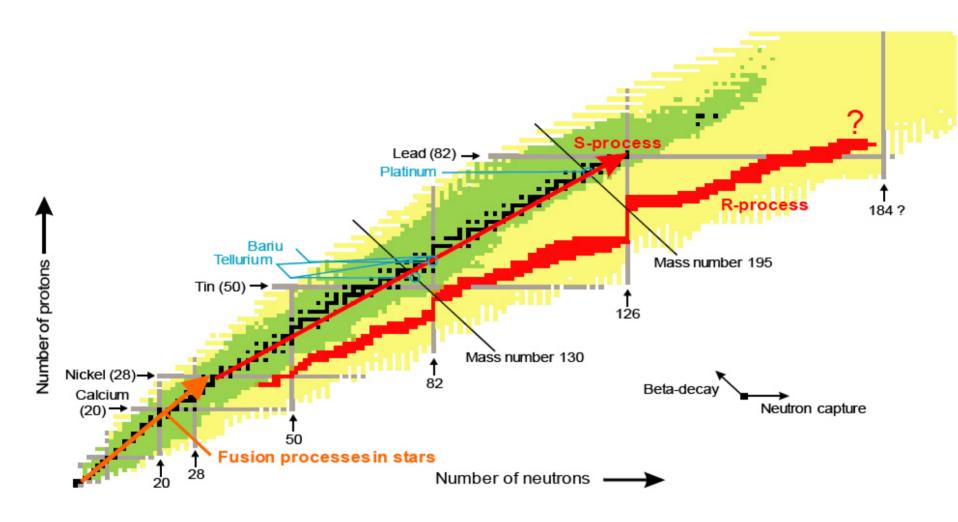
### The origin of heavy elements in the solar system



each process contribution is a mix of many events!



~35 p-nuclei cannot be made in s- or r-process: 74Se, 78Kr, 84Sr, 92Mo, 94Mo, 96Ru, 98Ru, 102Pd, 106Cd, 108Cd, 112Sn,113In, 114Sn, 115Sn, 120Te, 124Xe, 126Xe, 130Ba, 132Ba,136Ce, 138La, 138Ce, 144Sm, 152Gd, 156Dy, 158Dy, 162Er,164Er, 168Yb, 174Hf, 180Ta, 180W, 184Os, 190Pt, and 196Hg

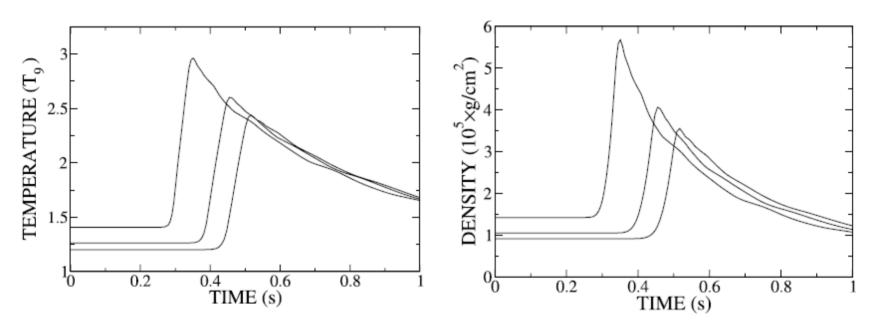
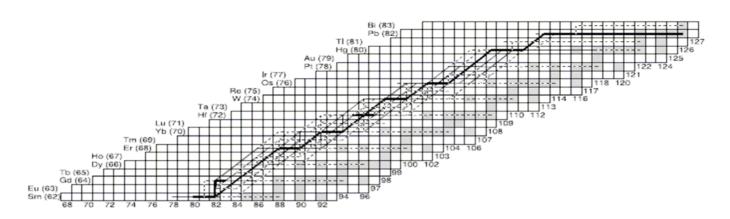
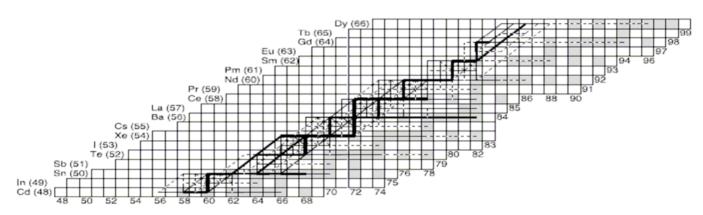
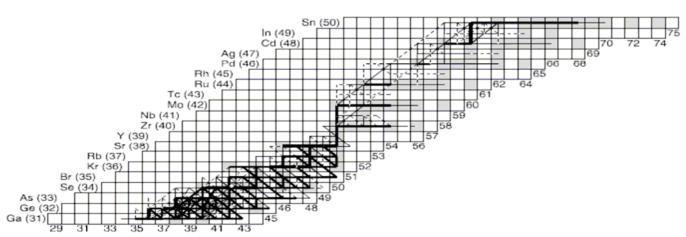


Fig. 2.—Temperature and density profiles of the SN shock front traversing the Ne/O layer of the pre-SN star.







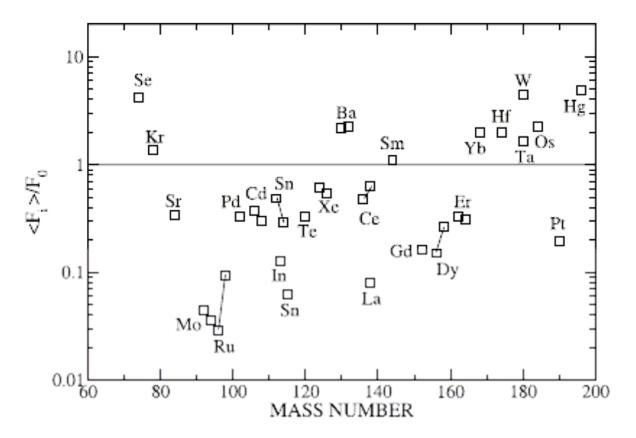
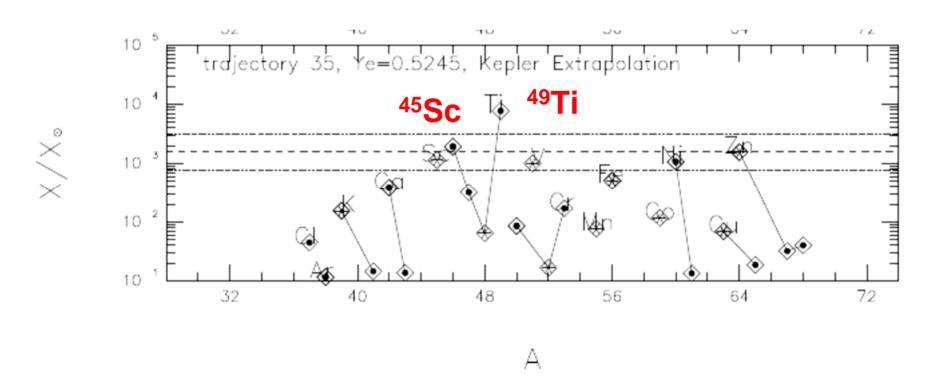


Fig. 4.—Averaged normalized overproduction factor for the proton-rich p-nuclei from network calculation with standard reaction rates (see text).

# The vp process

J.Pruet et al. 2004: early neutrino driven winds in core collapse SNe are p-rich  $(Y_e \sim 0.5)$ 

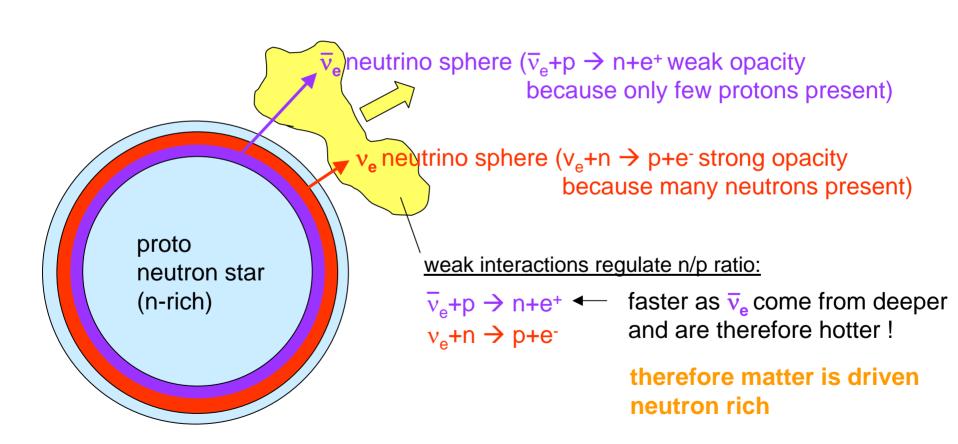


→ Origin of some isotopes not fully accounted for in SN?

## Recall: r-process in Supernovae?

Most favored scenario for high entropy:

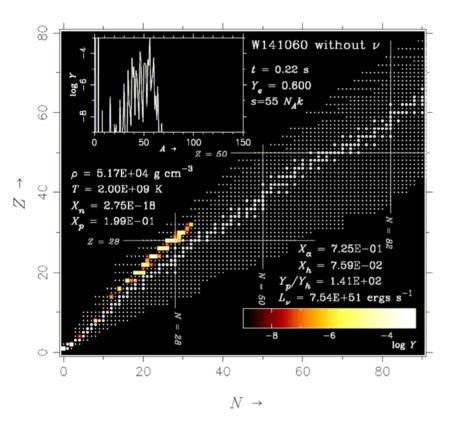
Neutrino heated wind evaporating from proto neutron star in core collapse

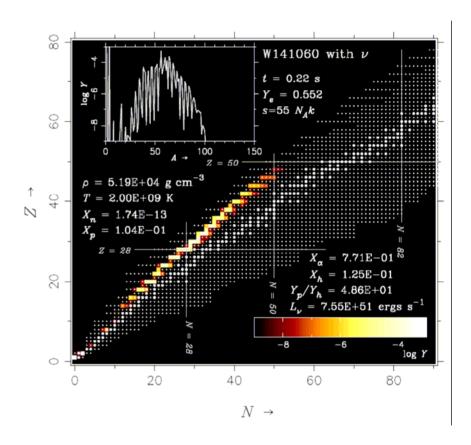


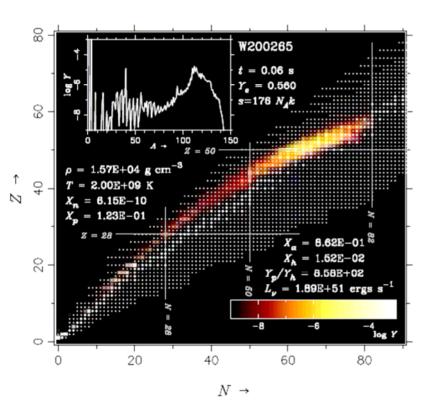
# The vp process – the role of neutrinos?

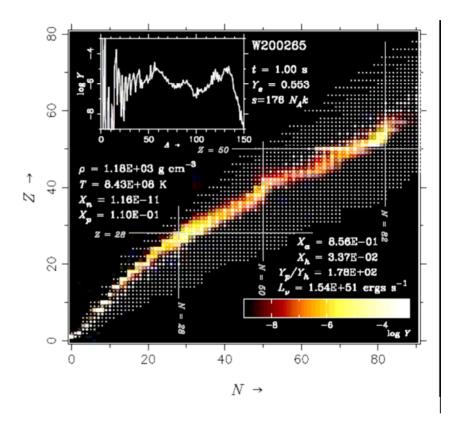
Froehlich et al. 2006 Pruet et al. 2006 Wanajo et al. 2006

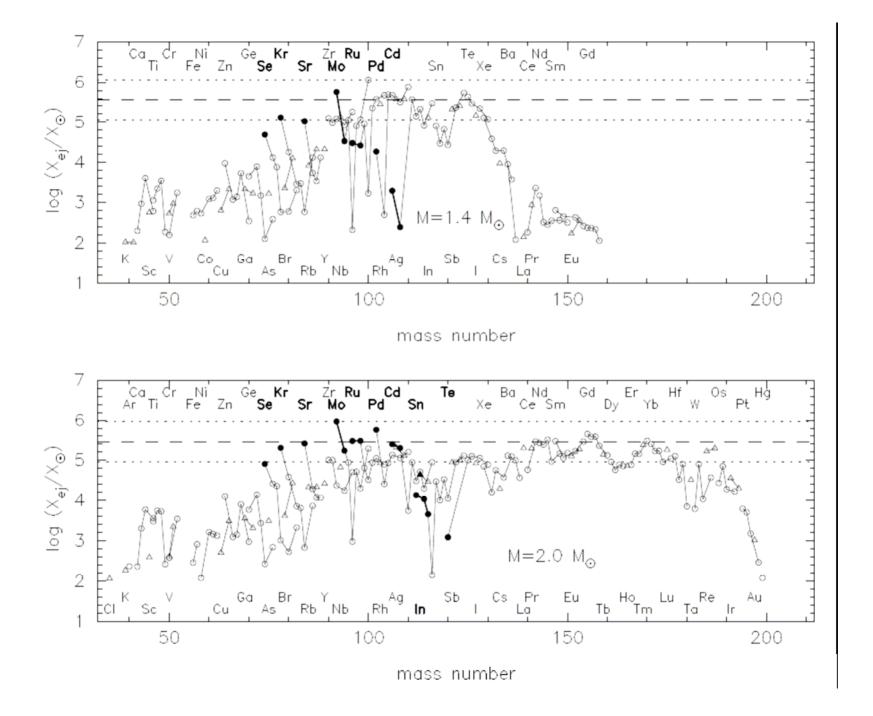
Recall  $\alpha$ -effect in r-process ...



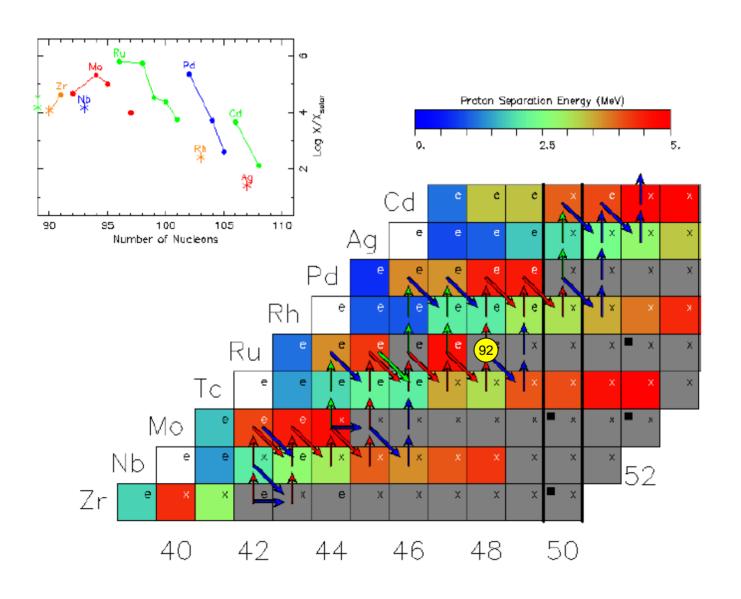




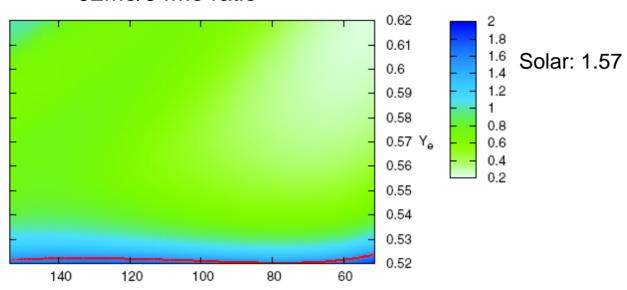




#### Fisker et al. 2009 – can one make <sup>92</sup>Mo/<sup>94</sup>Mo in solar proportions?

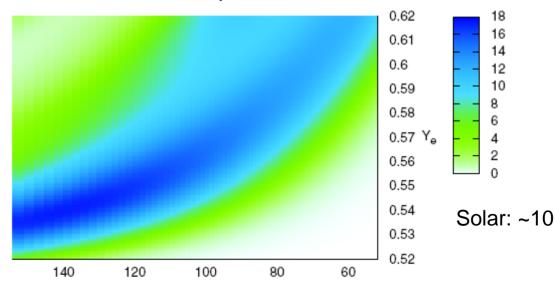


#### 92Mo/94Mo ratio



Entropy/baryon

#### 92Mo overproduction



Entropy/baryon